

# 1. Symmetries in classical mechanics and classical field theory

## Principle of least action

$$L(\underbrace{q_1, \dots, q_f}_{\text{generalized coordinates}}, \underbrace{\dot{q}_1, \dots, \dot{q}_f}_{\text{generalized velocities}}, t) \equiv L(q, \dot{q}, t)$$

f... number of dof

$$S = \int_{t_1}^{t_2} dt L(q, \dot{q}, t) \quad \text{action (integral)}$$

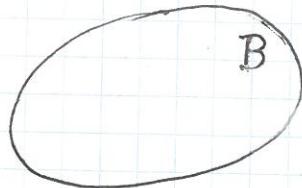
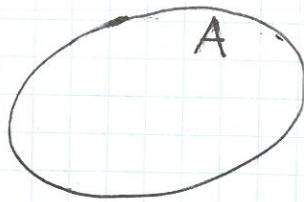
$$\delta S = 0 \Rightarrow \frac{d}{dt} \frac{\partial L}{\partial \dot{q}_i} = \frac{\partial L}{\partial q_i} \quad \begin{matrix} \text{equations of} \\ \text{motion} \end{matrix}$$

with  $\delta q(t_1) = 0$        $i=1, \dots, f$       (Euler-Lagrange)

remark:  $L \rightarrow L' = L + \frac{d}{dt} f(q, t)$  does not change EOM

$$\text{reason: } \delta \int_{t_1}^{t_2} dt \frac{d}{dt} f(q, t) = \delta [f(q(t_2), t_2) - f(q(t_1), t_1)] = 0$$

$L$  and  $L'$  are physically equivalent



closed subsystems A, B  
no interaction with each other

$$L \rightarrow L_A + L_B \quad \underline{\text{additivity}}$$

→ no individual rescaling of Lagrange functions of closed subsystems (only rescaling for all subsystems with the same factor)

point particle in inertial frame  $\vec{x}, \dot{\vec{x}} = \vec{v}$   
symmetry

inertial frame characterized by the following properties:

time homogeneous

space homogeneous and isotropic

$$\Rightarrow L = L(\vec{v}^2) \quad \text{EOM} \quad \frac{d}{dt} \frac{\partial L(\vec{v})}{\partial v_i} = 0 \Rightarrow v_i = \text{const.}$$

nonrelativistic mechanics → coordinates of inertial frames related by Galilei transformations

$$\begin{aligned} \text{universal time} \quad t' &= t & (\text{special case: no shift}) \\ \vec{x}' &= \vec{x} + \vec{u} t & (\text{special case: velocity transformation}) \end{aligned}$$

(3)

## Galileian principle of relativity -

→ L and L' equivalent (equal up to total time derivative)

$$L' = L(\vec{v}'^2) = L((\vec{v} + \vec{u})^2) = L(\vec{v}^2 + 2\vec{v} \cdot \vec{u} + \vec{u}^2)$$

consider infinitesimal  $\vec{u}$ :

$$\rightarrow L(\vec{v}'^2) = L(\vec{v}^2) + \underbrace{\frac{\partial L}{\partial \vec{v}^2} 2\vec{u} \cdot \vec{v}}_{\stackrel{\text{III}}{\rightarrow}} + \mathcal{O}(\vec{u}^2)$$

$\frac{d\vec{x}}{dt}$  already total time derivative

$$\Rightarrow \frac{\partial L}{\partial \vec{v}^2} = \text{const.} \Rightarrow L(\vec{v}^2) = \frac{m}{2} \vec{v}^2$$

constant m is called "mass"

$$\text{EOM} \quad \underbrace{\frac{d}{dt} \frac{\partial L}{\partial v_i}}_{p_i = m v_i \text{ momentum}} = \underbrace{\frac{\partial L}{\partial x_i}}_{\vec{p} = m \vec{v}} \quad i=1,2,3$$

$$\dot{p}_i = 0 \Rightarrow p_i = \text{const.} \Rightarrow v_i = \text{const.} \quad (\text{Newton I})$$

finite Galilei transformation:

$$L' = \frac{m}{2} \vec{v}'^2 = \frac{m}{2} (\vec{v} + \vec{u})^2 = \frac{m}{2} \vec{v}^2 + \underbrace{m \vec{v} \cdot \vec{u} + \frac{m}{2} \vec{u}^2}_{\frac{d}{dt} (m \vec{x} \cdot \vec{u} + \frac{m}{2} \vec{u}^2 t)}$$

system of noninteracting particles  $L = \sum_a \frac{m_a}{2} \vec{v}_a^2$  (additivity)

general case  $L = L(q_1, \dots, q_f, \dot{q}_1, \dots, \dot{q}_f, t)$

$$p_i = \frac{\partial L}{\partial \dot{q}_i} \quad \text{generalized momenta} \quad (i=1, \dots, f)$$

if  $L$  independent of  $q_i$  (symmetry!)  $\Rightarrow \underbrace{\frac{d}{dt} \frac{\partial L}{\partial \dot{q}_i}}_{p_i} = 0$   
 $\Rightarrow p_i = \text{const.}$

example: Lagrange function of free particle (in inertial frame) written in cylindrical coordinates

$$L = \frac{m}{2} (\dot{r}^2 + r^2 \dot{\varphi}^2 + \dot{z}^2)$$

$$\frac{\partial L}{\partial \dot{\varphi}} = 0 \quad \Rightarrow \quad p_\varphi \equiv L_z = mr^2 \dot{\varphi} = \text{const.}$$

rotational symmetry- component of angular momentum in  $z$ -direction

Time translation symmetry =

if  $\frac{\partial L}{\partial t} = 0$  (no explicit time-dependence)  $L = L(q, \dot{q})$

$$\Rightarrow \frac{d}{dt} L(q, \dot{q}) = \sum_{i=1}^f \left( \underbrace{\frac{\partial L}{\partial q_i}}_{\text{EOM}} \dot{q}_i + \underbrace{\frac{\partial L}{\partial \dot{q}_i}}_{\frac{d}{dt} \frac{\partial L}{\partial \dot{q}_i}} \ddot{q}_i \right) = \frac{d}{dt} \sum_{i=1}^f \underbrace{\frac{\partial L}{\partial \dot{q}_i}}_{\frac{d}{dt} \frac{\partial L}{\partial \dot{q}_i}} \dot{q}_i$$

(5)

$$\Rightarrow \text{At } t= \frac{d}{dt} \left( \underbrace{\sum_{i=1}^k \frac{\partial L}{\partial \dot{q}_i} \dot{q}_i - L}_{p_i} \right) = 0$$

$$\Rightarrow E := \sum_{i=1}^k \frac{\partial L}{\partial \dot{q}_i} \dot{q}_i - L = \text{const.}$$

called "energy"

example: point particle in inertial frame

$$E = m \vec{v} \cdot \vec{v} - \frac{m \vec{v}^2}{2} = \frac{m \vec{v}^2}{2} \quad \begin{matrix} \text{nonrelativistic} \\ \checkmark \text{Kinetic energy} \end{matrix}$$

Galilean invariant system of two point particles (closed system)

$$L = \frac{m_1 \vec{v}_1^2}{2} + \frac{m_2 \vec{v}_2^2}{2} - U(|\vec{x}_1 - \vec{x}_2|) \quad \text{EOM } m_a \ddot{\vec{v}}_a = - \underbrace{\frac{\partial U}{\partial \vec{x}_a}}_{\vec{F}_a} \quad a=1,2$$

$$0 = \frac{dL}{dt} \Rightarrow E = \frac{m_1 \vec{v}_1^2}{2} + \frac{m_2 \vec{v}_2^2}{2} + U(|\vec{x}_1 - \vec{x}_2|) \quad \begin{matrix} \text{total energy} \\ \underline{\text{conserved}} \end{matrix}$$

space homogeneous

$L$  invariant under  $\vec{x}_a \rightarrow \vec{x}_a + \vec{\epsilon}$ ,  $a=1,2$

$$\delta L = \sum_a \frac{\partial L}{\partial \vec{x}_a} \cdot \vec{\epsilon} = \vec{\epsilon} \cdot \sum_a \frac{\partial L}{\partial \vec{x}_a} = 0 \quad \begin{matrix} \vec{\epsilon} \text{ arbitrary} \\ \sum_a \frac{\partial L}{\partial \vec{x}_a} = 0 \end{matrix} \quad (\text{Newton III})$$

$$\Rightarrow \sum_a \frac{d}{dt} \frac{\partial L}{\partial \vec{v}_a} = \sum_a \frac{\partial L}{\partial \vec{x}_a} = 0 \quad \Rightarrow \underbrace{\frac{d}{dt} \sum_a \frac{\partial L}{\partial \vec{x}_a}}_{\vec{P}} = 0$$

here:  $\vec{P} = m_1 \vec{v}_1 + m_2 \vec{v}_2$

total momentum

(6)

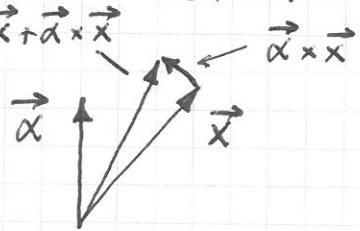
## space isotropic

L invariant under

$$\vec{x}_a \rightarrow \vec{x}_a + \underbrace{\vec{\alpha} \times \vec{x}_a}_{\delta \vec{x}_a}$$

$$\Rightarrow \vec{v}_a \rightarrow \vec{v}_a + \underbrace{\vec{\alpha} + \vec{v}_a}_{\delta \vec{v}_a}$$

infinitesimal rotation



$$\delta L = \sum_a \left( \underbrace{\frac{\partial L}{\partial \vec{x}_a} \cdot \delta \vec{x}_a}_{\text{EOM } \dot{\vec{p}}_a} + \underbrace{\frac{\partial L}{\partial \vec{v}_a} \cdot \delta \vec{v}_a}_{\vec{p}_a} \right) =$$

$$= \sum_a [ \dot{\vec{p}}_a \cdot (\vec{\alpha} \times \vec{x}_a) + \vec{p}_a \cdot (\vec{\alpha} \times \vec{v}_a) ]$$

$$= \vec{\alpha} \cdot \sum_a ( \vec{x}_a \times \dot{\vec{p}}_a + \dot{\vec{x}}_a \times \vec{p}_a )$$

$$= \vec{\alpha} \cdot \frac{d}{dt} \sum_a (\vec{x}_a \times \vec{p}_a) \xrightarrow{\vec{\alpha} \text{ arbitrary}} \frac{d}{dt} \underbrace{\sum_a (\vec{x}_a \times \vec{p}_a)}_{\vec{L}} = 0$$

total angular momentum conserved

$$\vec{L} = \sum_a (\vec{x}_a \times \vec{p}_a)$$

nonrelativistic:  $\vec{L} = \sum_a m_a (\vec{x}_a \times \vec{p} \vec{v}_a)$

(7)

## Relativistic mechanics

invariant  $ds^2 = c^2 dt^2 - d\vec{x}^2$  proper time:  $d\tau = ds/c$

$$\rightarrow S = -\alpha \int_1^2 ds = -\alpha c \int_{t_1}^{t_2} dt \sqrt{1 - \frac{\vec{v}^2}{c^2}}$$

$$\rightarrow L = -\alpha c \sqrt{1 - \frac{\vec{v}^2}{c^2}} \xrightarrow{|\vec{v}| \ll c} -\alpha c \left(1 - \frac{\vec{v}^2}{2c^2} + \dots\right)$$

$$= -\alpha c + \frac{\alpha}{c} \frac{\vec{v}^2}{2} + \dots$$

$$\Rightarrow \frac{\alpha}{c} = m$$

$$\rightarrow S = -mc \int ds = -mc^2 \int dt \sqrt{1 - \frac{\vec{v}^2}{c^2}}$$

$$L = -mc^2 \sqrt{1 - \frac{\vec{v}^2}{c^2}}$$

rel. momentum  $\vec{p} = \frac{\partial L}{\partial \vec{v}} = \frac{m\vec{v}}{\sqrt{1 - \frac{\vec{v}^2}{c^2}}}$

rel. energy  $E = \vec{p} \cdot \vec{v} - L = \frac{mc^2}{\sqrt{1 - \frac{\vec{v}^2}{c^2}}}$

$$\xrightarrow{|\vec{v}| \ll c} mc^2 + \frac{m\vec{v}^2}{2} + \dots$$

rest energy.

(8)

rel. energy-momentum relation  $\rightarrow$  Hamilton function

$$E = \sqrt{m^2 c^4 + \vec{p}^2 c^2} \xrightarrow{|\vec{p}| \ll mc} m c^2 + \frac{\vec{p}^2}{2m} + \dots$$

massless particle :  $E = |\vec{p}|c$

velocity :  $\vec{v} = \frac{\vec{p} c^2}{E} \xrightarrow{m=0} \vec{v} = \frac{\vec{p}}{|\vec{p}|} c$

$|\vec{v}| = c$  for massless particle

angular momentum :  $\vec{L} = \vec{x} \times \vec{p}$

Noether theorem

general form of a symmetry transformation

$$t \rightarrow t', q(t) \rightarrow q'(t')$$

EOM (form) invariant

sufficient condition for invariance :

$$\begin{aligned} S[q'(t')] &= S[q(t)] \\ \Leftrightarrow \int_{t_1'}^{t_2'} L(q'(t'), \dot{q}'(t'), t') dt' &= \int_{t_1}^{t_2} L(q(t), \dot{q}(t), t) dt \end{aligned}$$

classical physics: only continuous transformations lead to conserved quantities

relevant structure already contained in infinitesimal transformations

→ Lie groups

$$t \rightarrow t' = t + \alpha \tau(q, \dot{q}, t) + O(\alpha^2)$$

$$q(t) \rightarrow q'(t') = q(t) + \alpha g(q, \dot{q}, t) + O(\alpha^2)$$

action invariant:

$$\begin{aligned} 0 &= S[q'(t')] - S[q(t)] \\ &= \int_{t'_1}^{t'_2} dt' L(q'(t'), \dot{q}'(t'), t') - S[q(t)] \\ &= \int_{t_1}^{t_2} dt \left(1 + \alpha \frac{dt}{dt}\right) L[q(t) + \cancel{\alpha \tau(q, \dot{q}, t)}, \\ &\quad (\dot{q}(t) + \cancel{\alpha \frac{dq}{dt}}) (1 - \alpha \frac{dt}{dt}), \\ &\quad t + \alpha \tau] - S[q(t)] \\ &= \int_{t_1}^{t_2} dt \left(1 + \alpha \frac{dt}{dt}\right) L[q + \alpha g, \dot{q} + \alpha \frac{dg}{dt} - \dot{q}(t) \alpha \frac{dt}{dt}, \\ &\quad t + \alpha \tau] - S[q(t)] \end{aligned}$$

$$= \alpha \int_{t_1}^{t_2} dt \left[ \sum_{i=1}^l \underbrace{\frac{\partial L}{\partial \dot{q}_i}}_{\frac{d}{dt} \frac{\partial L}{\partial \dot{q}_i}} \dot{q}_i + \sum_{i=1}^l \frac{\partial L}{\partial \dot{q}_i} \left( \frac{d \dot{q}_i}{dt} - \dot{q}_i \frac{d \tau}{dt} \right) \right. \\ \left. + \frac{\partial L}{\partial t} \tau + L \frac{d \tau}{dt} \right]$$

$$\Rightarrow \frac{d}{dt} \sum_{i=1}^l \frac{\partial L}{\partial \dot{q}_i} \dot{q}_i + \left( L - \sum_{i=1}^l \frac{\partial L}{\partial \dot{q}_i} \dot{q}_i \right) \frac{d \tau}{dt} + \frac{\partial L}{\partial t} \tau = 0$$

$$\frac{d}{dt} \left( L - \sum_{i=1}^l \frac{\partial L}{\partial \dot{q}_i} \dot{q}_i \right) = \frac{\partial L}{\partial t} + \sum_{i=1}^l \cancel{\frac{\partial L}{\partial \dot{q}_i} \dot{q}_i} + \cancel{\frac{\partial L}{\partial \dot{q}_i} \ddot{q}_i}$$

$$- \cancel{\frac{\partial L}{\partial \dot{q}_i} \dot{q}_i} - \cancel{\frac{\partial L}{\partial \dot{q}_i} \ddot{q}_i} = \frac{\partial L}{\partial t}$$

$$\Rightarrow \frac{d}{dt} \left[ \sum_{i=1}^l \frac{\partial L}{\partial \dot{q}_i} \dot{q}_i + \left( L - \sum_{i=1}^l \frac{\partial L}{\partial \dot{q}_i} \dot{q}_i \right) \tau \right] = 0$$

action invariant under continuous transformation

$$t \rightarrow t' = t + \alpha \tau(q, \dot{q}, t) + O(\alpha^2)$$

$$q(t) \rightarrow q'(t') = q(t) + \alpha g(q, \dot{q}, t) + O(\alpha^2)$$

$\Rightarrow$  constant of motion

$$\sum_{i=1}^l \frac{\partial L}{\partial \dot{q}_i} \dot{q}_i + \left( L - \sum_{i=1}^l \frac{\partial L}{\partial \dot{q}_i} \dot{q}_i \right) \tau = \text{const.}$$

## generalized Noether theorem

$L$  only determined up to total derivative!

$$(1 + \alpha \frac{d\tau}{dt}) L [q(t) + \alpha g(q, \dot{q}, t), (\dot{q}(t) + \alpha \frac{dg}{dt}), (1 - \alpha \frac{d\tau}{dt}), t + \alpha \tau]$$

$$= L(q(t), \dot{q}(t), t) + \alpha \frac{d}{dt} f(q(t), t)$$

$$\Rightarrow \sum_{i=1}^f \frac{\partial L}{\partial \dot{q}_i} g_i + \left( L - \sum_{i=1}^f \frac{\partial L}{\partial \dot{q}_i} \dot{q}_i \right) \tau - f = \text{const.}$$

consider Galilei-invariant Lagrangian function

$$L(\vec{x}_1, \dots, \vec{x}_N, \dot{\vec{x}}_1, \dots, \dot{\vec{x}}_N) = \underbrace{\sum_a \frac{m_a \dot{\vec{x}}_a^2}{2}}_T - U(|\vec{x}_a - \vec{x}_b|)$$

Galilei group (symmetry group of nonrelativistic mechanics)

$$\text{Galilei transformation } \vec{x}'_a = R \vec{x}_a + \vec{u} t + \vec{\xi}$$

$$t' = t + t_0$$

$$R \in SO(3, \mathbb{R}) \quad R^T R = \mathbb{1}$$

$$\det R = 1$$

$$\vec{u} \in \mathbb{R}^3$$

$$\vec{\xi} \in \mathbb{R}^3$$

$$R, \vec{u}, \vec{\xi}, t_0 \quad 3+3+3+1 = 10 \text{ parameters}$$

$\rightarrow 10$  constants of motion

(i) invariance under spatial translations

$\vec{\xi} \rightarrow 3 \text{ parameters}$

$$\vec{x}_a(t) \rightarrow \vec{x}'_a(t) = \vec{x}_a(t) + \underbrace{\alpha \vec{d}}_{\vec{\xi}}$$

arbitrary constant vector

$$t' = t$$

$$\dot{\vec{x}}_a(t) = \dot{\vec{x}}'_a(t) \Rightarrow T \text{ invariant}$$

$U$  depends only on  $|\vec{x}_a - \vec{x}_b| \Rightarrow U \text{ invariant}$

$$\vec{g}_a = \vec{d}, \tau = 0$$

$$\Rightarrow \sum_{a=1}^N \vec{p}_a \cdot \vec{d} = \text{const}$$

$\vec{d}$  arbitrary  $\Rightarrow$  total momentum  $\sum_{a=1}^N \vec{p}_a =: \vec{P}$  conserved

(ii) invariance under time translation ( $t_0 \rightarrow 1 \text{ parameter}$ )

$$t' = t + \alpha$$

$$\vec{x}'_a(t') = \vec{x}'_a(t + \alpha) = \vec{x}_a(t)$$

$$\tau = 1, \vec{g}_a = 0$$

(13)

$$\frac{\partial U}{\partial t} = 0 \Rightarrow L \text{ invariant}$$

Noether

$$\Rightarrow \underbrace{L}_{T-U} - \underbrace{\sum_{a=1}^N m_a \dot{\vec{x}}_a^2}_{2T} = - (T+U) = -E \text{ conserved}$$

energy conservation

(iii) invariance under spatial rotations ( $R \rightarrow 3$  parameters)

$$t' = t \Rightarrow \tau = 0$$

$$\vec{x}'_a = \vec{x}_a + \alpha \underbrace{\vec{n} \times \vec{x}_a}_{\vec{g}_a}, \quad |\vec{n}| = 1$$

$$\dot{\vec{x}}'_a = R \dot{\vec{x}}_a \Rightarrow \dot{\vec{x}}'_a \cdot \dot{\vec{x}}'_a = \dot{\vec{x}}_a \cdot \dot{\vec{x}}_a \quad \text{because of } R^T R = \mathbb{1}$$

$\Rightarrow T$  invariant

$|\vec{x}_a - \vec{x}_b|$  invariant under rotations  $\Rightarrow U$  invariant

Noether

$$\Rightarrow \sum_a \vec{p}_a \cdot (\vec{n} \times \vec{x}_a) = \vec{n} \cdot \underbrace{\sum_a \vec{x}_a \times \vec{p}_a}_{\vec{L}} \text{ conserved}$$

$\vec{n}$  arbitrary unit vector  $\Rightarrow$  total angular momentum  $\vec{L}$  conserved

(14)

(iv) pure velocity transformation (boost)  
 $(\vec{u} \rightarrow 3 \text{ parameters})$

$$t' = t \Rightarrow \tau = 0$$

$$\vec{x}_a'(t) = \vec{x}_a(t) + \underbrace{\alpha \vec{n} t}_{\vec{u}}, |\vec{n}| = 1$$

$$\Rightarrow \vec{g}_a = \vec{n} t$$

$|\vec{x}_a - \vec{x}_b|$  invariant  $\Rightarrow U$  invariant

$$T' = \sum_{a=1}^N \frac{m_a \dot{\vec{x}}_a'^2}{2} = \sum_{a=1}^N \frac{m_a}{2} (\dot{\vec{x}}_a + \alpha \vec{n})^2$$

$$= T + \alpha \underbrace{\sum_{a=1}^N m_a \vec{x}_a \cdot \vec{n}}_{\frac{d}{dt} \vec{n} \cdot \sum_{a=1}^N m_a \vec{x}_a} + O(\alpha^2)$$

$$\underbrace{f(\vec{x}_1, \dots, \vec{x}_N)}$$

T invariant up to total derivative

generalized Noether theorem  $\Rightarrow$

$$\sum_{a=1}^N \vec{p}_a \cdot \vec{n} t - \vec{n} \cdot \sum_{a=1}^N m_a \vec{x}_a = \vec{n} \sum_{a=1}^N (\vec{p}_a t - m_a \vec{x}_a) = \text{const.}$$

$$CM \quad \vec{R} = \frac{1}{M} \sum_{a=1}^N m_a \vec{x}_a(t), \quad M = \sum_{a=1}^N m_a$$

$$\Rightarrow \vec{R}(t) = \vec{R}_0 + \frac{\vec{P}}{M} t \quad \begin{matrix} \text{movement of CM} \\ (\text{CM theorem}) \end{matrix}$$

remark: (spatial) translation invariance  $\Rightarrow U$  invariant under velocity transf.

$\Rightarrow$  conservation of total momentum implies CM theorem

$$\dot{\vec{R}}(t) = \frac{1}{M} \sum_{a=1}^N m_a \dot{\vec{x}}_a(t) = \frac{\vec{P}}{M} \xrightarrow{\text{conserved}} = \text{const.}$$

$$\xrightarrow{\text{integration}} \vec{R}(t) = \vec{R}_0 + \frac{\vec{P}}{M} t$$

## (Relativistic) field theory



action Lorentz invariant

$$x^0 = ct, \vec{x} \rightarrow x^\mu = (x^0, \vec{x})$$

from now on:  $c = 1$

Lagrangian density (Lagrangian)

$$S = \underbrace{\int d^4x \mathcal{L}(\phi_a(x), \phi_{a,\mu}(x), x)}_{\text{Lorentz scalar}}$$

Lorentz scalar

$\mathcal{L}$  Lorentz invariant  $\Rightarrow S$  Lorentz invariant

(inverse not necessary, but in (almost) all relevant cases:  $\mathcal{L}$  inv.)

## Principle of least action for fields

fields  $\phi(x_1^o, \vec{x})$  and  $\phi(x_2^o, \vec{x})$  fixed

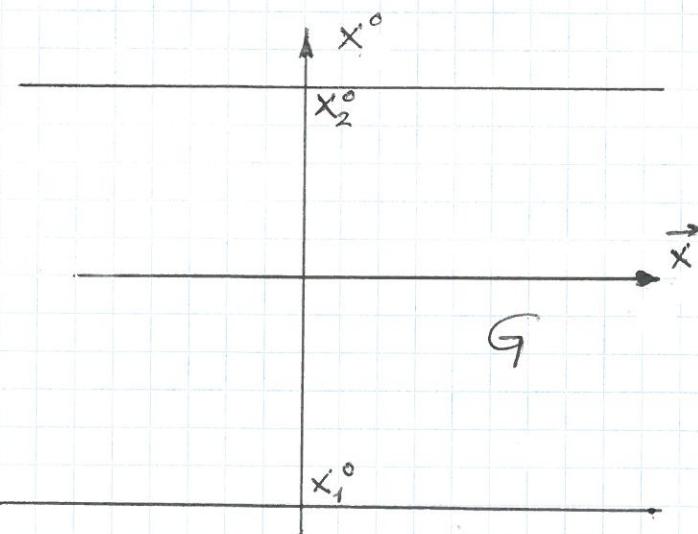
for  $x_1^o \leq x^o \leq x_2^o$ : choose  $\phi(x)$  such that  $S \rightarrow \min.$

assumption:  $\lim_{|\vec{x}| \rightarrow \infty} \phi(x^o, \vec{x}) = 0$

field variation  $\phi(x) \rightarrow \phi(x) + \eta(x)$

with  $\eta(x_1^o, \vec{x}) = \eta(x_2^o, \vec{x}) = 0$ ,  $\lim_{|\vec{x}| \rightarrow \infty} \eta(x^o, \vec{x}) = 0$

$\delta S = 0$  ( $S$  extremal)



$$\delta S = \int dx^o \int d^3x [L(\phi + \eta, \phi_\mu + \eta_\mu, x) - L(\phi, \phi_\mu, x)] \quad (\text{index } \alpha \text{ suppressed})$$

$$= \int_{x_1^o}^{x_2^o} dx^o \int d^3x \left[ \frac{\partial L}{\partial \phi} \eta + \frac{\partial L}{\partial \phi_\mu} \eta_\mu + O(\eta^2) \right]$$

$$\underbrace{\frac{\partial}{\partial x^\mu} \left( \frac{\partial L}{\partial \phi_\mu} \eta \right)}_{\left( \frac{\partial}{\partial x^\mu} \frac{\partial L}{\partial \phi_\mu} \right) \eta} - \left( \frac{\partial}{\partial x^\mu} \frac{\partial L}{\partial \phi_\mu} \right) \eta$$

$$= \int_{x_1^o}^{x_2^o} dx^o \int_{\mathbb{R}^3} d^3x \left[ \left( \frac{\partial \mathcal{L}}{\partial \phi} - \frac{\partial}{\partial x^\mu} \frac{\partial \mathcal{L}}{\partial \phi_{,\mu}} \right) \eta + \frac{\partial}{\partial x^\mu} \left( \frac{\partial \mathcal{L}}{\partial \phi_{,\mu}} \eta \right) \right] \quad (17)$$

$$\int_{x_1^o}^{x_2^o} dx^o \int_{\mathbb{R}^3} d^3x \frac{\partial}{\partial x^\mu} \left( \frac{\partial \mathcal{L}}{\partial \phi_{,\mu}} \eta \right) =$$

$$= \underbrace{\int_{\mathbb{R}^3} d^3x \frac{\partial \mathcal{L}}{\partial \phi_{,0}} \eta}_{\text{O}} \Big|_{x_1^o}^{x_2^o} + \underbrace{\int_{x_1^o}^{x_2^o} dx^o \int_{\partial \mathbb{R}^3} df_i \left( \frac{\partial \mathcal{L}}{\partial \phi_{,i}} \eta \right)}_{\text{O}}$$

because of  $\eta(x_2^o, \vec{x}) = 0$

because of  $\eta(x; \vec{x}) \rightarrow 0$   
 $|x| \rightarrow \infty$

$\eta$  arbitrary in  $G \Rightarrow$  Euler-Lagrange field equation

$$\boxed{\frac{\partial}{\partial x^\mu} \frac{\partial \mathcal{L}}{\partial \phi_{,\mu}} = \frac{\partial \mathcal{L}}{\partial \phi}}$$

remark: remember Gauss theorem in  $d=3$

$$\int_V d^3x \operatorname{div} \vec{A} = \int_{\partial V} \vec{df} \cdot \vec{A}$$

surface  $\vec{x}(u, v)$



oriented surface element  $\vec{df} = \frac{\partial \vec{x}}{\partial u} \times \frac{\partial \vec{x}}{\partial v} du dv$

$\partial V$  boundary of  $V$

$$df_i = \epsilon_{ijk} \frac{\partial x_j}{\partial u} \frac{\partial x_k}{\partial v} du dv$$

theorem of Gauss in  $d=4$

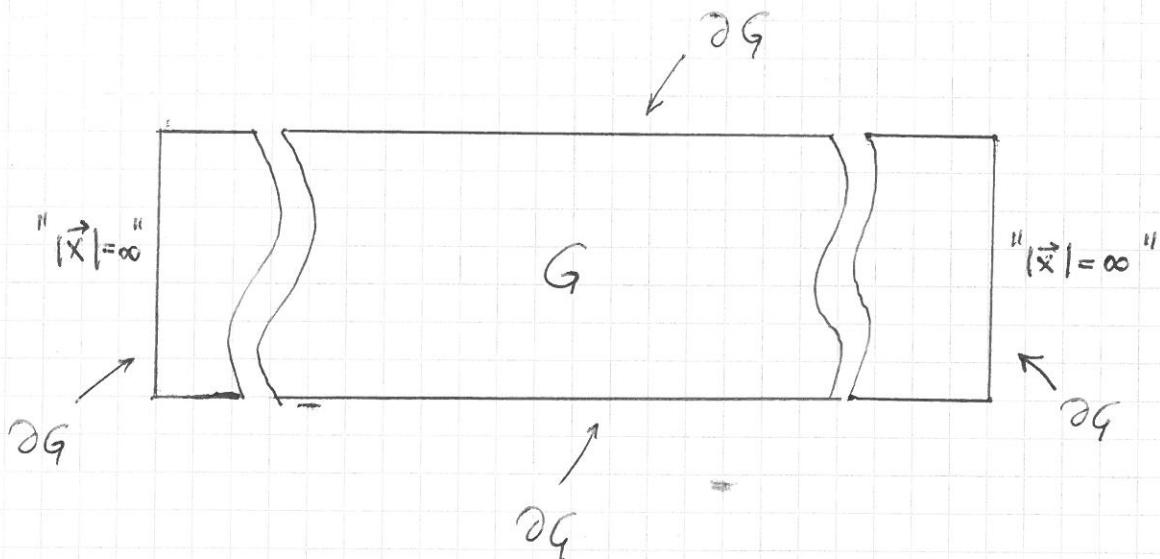
$$\int_G d^4x \partial_\mu A^\mu = \int_{\partial G} d\sigma_\mu A^\mu$$

3-dim. hypersurface in  $d=4$ :  $x(u, v, w)$

$$\rightarrow \text{oriented surface element } d\sigma_\mu = \epsilon_{\mu\nu\gamma\sigma} \frac{\partial x'}{\partial u} \frac{\partial x^g}{\partial v} \frac{\partial x^o}{\partial w} du dv dw$$

$\epsilon_{\mu\nu\gamma\sigma}$  totally antisymmetric, convention:  $\epsilon_{0123} = +1$

in our case:



$$\eta \Big|_{\partial G} = 0$$

some field equations:

(a) Klein-Gordon equation (free scalar field)

$$S = \int d^4x \frac{1}{2} (\partial^\mu \phi \partial_\mu \phi - m^2 \phi^2), \quad \phi^* = \phi$$

$$\Rightarrow (\square + m^2) \phi = 0, \quad \square = \partial_\mu \partial^\mu = \frac{\partial^2}{\partial t^2} - \Delta$$

↑  
d'Alembert operator

(b) Massive vector field (Proca equations)

real vector field  $A_\mu(x)$

$$F_{\mu\nu} := \partial_\mu A_\nu - \partial_\nu A_\mu, \quad m \neq 0 !!!$$

$$S = \int d^4x \left( -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + \frac{m^2}{2} A_\mu A^\mu \right)$$

$$\Rightarrow (\square + m^2) A_\mu = 0, \quad \partial_\mu A^\mu = 0$$

(c) Massless vector field (electrodynamics)

$$S = \int d^4x \left( -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} - j^\mu A_\mu \right)$$

$$F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu, \quad \partial_\mu j^\mu = 0$$

$\Rightarrow$  field equation  $\partial_\nu F^{\nu\mu} = j^\mu$  (Maxwell)

Conservation laws in relativistic field theory

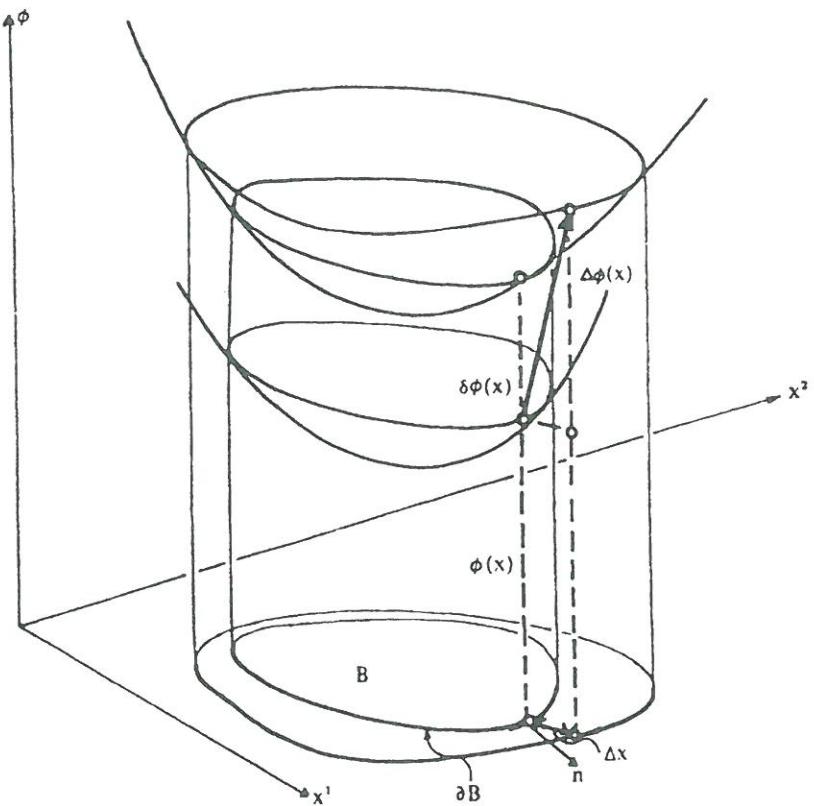
$$S = \int_D d^4x \mathcal{L}(\phi_a, \phi_{a,\mu}, x)$$

$D$  ... 4-dimensional domain in Minkowski space  
with boundary  $\partial D$

change of  $S$  under infinitesimal change  
of  $\phi_a$  as well as  $D$ ?

for every  $x$   $\phi_a(x)$  changed by  $\delta\phi_a(x)$

$x$  shifted by  $\Delta x$  (displacement vector)

Fig. 10.1. The graph of  $\phi_a = \phi_a(x)$  and its variation in  $(x, \phi)$ -space

$$\Delta S = \int_{\mathcal{D}} d^4x \left[ \frac{\partial \mathcal{L}}{\partial \phi_a} \delta \phi_a + \frac{\partial \mathcal{L}}{\partial \phi_{a,\mu}} \delta \phi_{a,\mu} \right] + \int_{\partial \mathcal{D}} d\sigma_\mu \mathcal{L} \Delta x^\mu$$

summation over  $a$ !

$$\text{Gauß} = \int_{\mathcal{D}} d^4x \left[ \frac{\partial \mathcal{L}}{\partial \phi_a} - \partial_\mu \frac{\partial \mathcal{L}}{\partial \phi_{a,\mu}} \right] \delta \phi_a$$

$$+ \int_{\partial \mathcal{D}} d\sigma_\mu \left[ \frac{\partial \mathcal{L}}{\partial \phi_{a,\mu}} \delta \phi_a + \mathcal{L} \Delta x^\mu \right]$$

$\Delta \phi_a = \phi_{a,\mu} \Delta x^\mu$

"vertical" variation  $\delta\phi_a(x)$

$$\begin{aligned} \text{"skew" variation } \Delta\phi_a(x) &= (\phi_a + \delta\phi_a)(x + \Delta x) - \phi_a(x) \\ &= \delta\phi_a(x) + \phi_{a,\mu}(x) \Delta x^\mu \end{aligned}$$

$$\Delta S = \int_D d^4x \left[ \frac{\partial L}{\partial \phi_a} - \partial_\mu \frac{\partial L}{\partial \phi_{a,\mu}} \right] \delta\phi_a$$

$$+ \int_{\partial D} d\sigma_\mu \left[ \frac{\partial L}{\partial \phi_{a,\mu}} \Delta\phi_a - \underbrace{\left( \frac{\partial L}{\partial \phi_{a,\mu}} \phi_{a,\nu} - L \delta^\mu_\nu \right)}_{\Theta^\mu_\nu} \Delta x^\nu \right]$$

we recover the Euler-Lagrange field equations

$\partial D$  and values on  $\partial D$  kept fixed : ( $\Delta x^\mu = 0, \delta\phi_a = 0$  on  $\partial D$ )

$$\Rightarrow \Delta S = 0 \text{ implies } \frac{\partial L}{\partial \phi_a} - \partial_\mu \frac{\partial L}{\partial \phi_{a,\mu}} = 0$$

(no influence of special choice of  $D$ )

$\rightarrow L$  defines an action principle (variation principle)

on the other hand:  $D$  specified + given boundary values

for  $\phi_a$  on  $\partial D$   $\rightarrow$  variational problem

transformation  $(x, \phi) \rightarrow (x', \phi')$

field equations form-invariant if action integral invariant

$$x^\mu \rightarrow x'^\mu = X^\mu(x, \phi)$$

$$\phi_a \rightarrow \phi'_a = \Phi_a(x, \phi)$$

$$\phi'_a(x') = \Phi_a(x(x'), \phi(x(x')))$$

$$x \in \mathcal{D} \rightarrow x' \in \mathcal{D}'$$

$$S' = \int_{\mathcal{D}'} d^4 x' \mathcal{L}(x', \phi'(x'), \frac{\partial \phi'(x')}{\partial x'^\mu}) =$$

$$= \int_{\mathcal{D}} d^4 x \mathcal{L}'(x, \phi(x), \frac{\partial \phi(x)}{\partial x^\mu})$$

$\mathcal{L}'$  defined by this equation

$$\text{if } \mathcal{L} \equiv \mathcal{L}' \Rightarrow S = S'$$

one-parameter group of transformations

$$x'^\mu = X^\mu(x, \phi; \alpha)$$

$$\phi'_a = \Phi_a(x, \phi; \alpha)$$

$\alpha = 0 \leftrightarrow \text{identity}$

consequences of  $S' = S$  for infinitesimal  $\alpha$ :

$$x'^\mu = X^\mu(x, \phi; \alpha) = \underbrace{X^\mu(x, \phi; 0)}_{x^\mu} + \underbrace{\frac{\partial X^\mu}{\partial \alpha} \Big|_{\alpha=0}}_{\Delta X^\mu} (\alpha + O(\alpha^2))$$

$$\Delta \phi_a = \phi'_a(x') - \phi_a(x) = \underbrace{\Phi_a(x, \phi; \alpha)}_{\phi_a(x)} - \phi_a(x) + O(\alpha^2)$$

$$= \underbrace{\Phi_a(x, \phi; 0)}_{\phi_a(x)} + \underbrace{\frac{\partial \Phi_a}{\partial \alpha} \Big|_{\alpha=0}}_{\Delta \phi_a} (\alpha) - \phi_a(x) + O(\alpha^2)$$

$$\Rightarrow \Delta \phi_a = \frac{\partial \Phi_a}{\partial \alpha} \Big|_{\alpha=0} \quad \text{skew variation!}$$

assumption:  $\phi$  satisfies field equations

$$\Rightarrow \int \limits_{\partial D} d\tilde{\sigma}_\mu \underbrace{\left[ \frac{\partial \mathcal{L}}{\partial \phi_{a,\mu}} \Delta \phi_a - \Theta^\mu_\nu \Delta x^\nu \right]}_{j^\mu \alpha} = 0$$

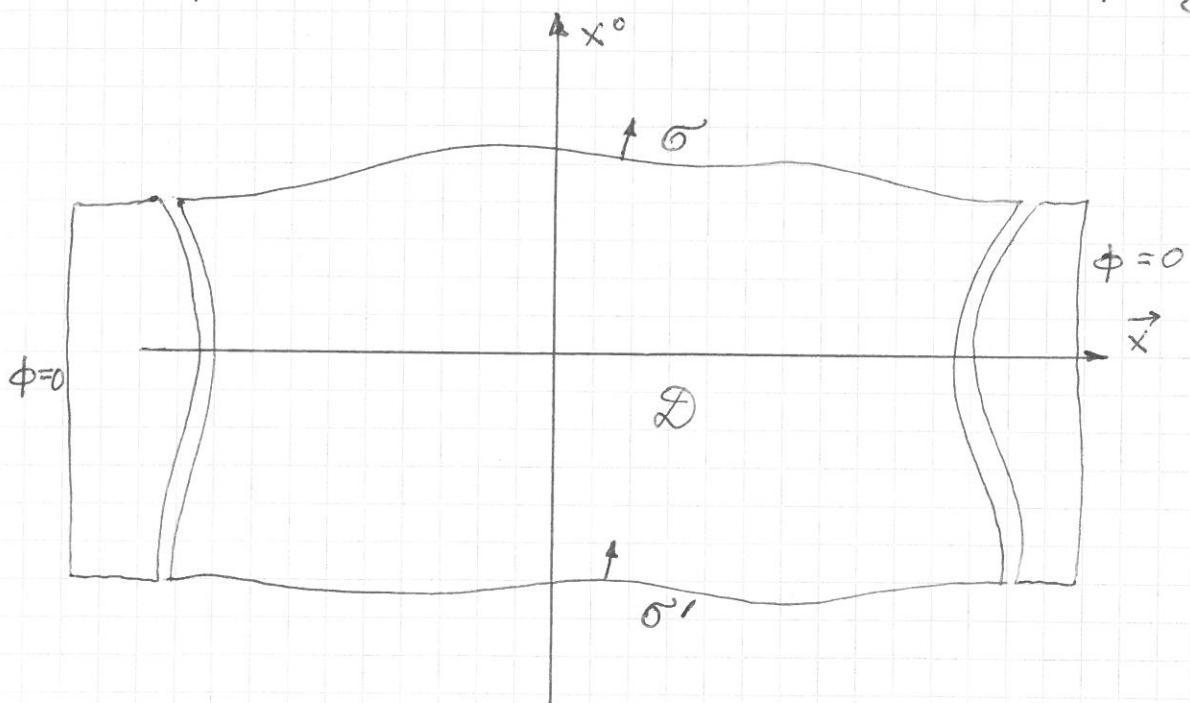
$$\text{Gauß} \Rightarrow \int \limits_D d^4 x j^\mu_{,\mu} = 0$$

$$\mathcal{D} \text{ arbitrary} \Rightarrow j^\mu_{,\mu} = 0$$

four-charge density  $j^\mu$  satisfies an equation of continuity

total charge  $Q = \int d\sigma_\mu j^\mu$  independent of the choice of the spacelike hyper-surface  $\sigma$

assumption:  $\phi(x) \rightarrow 0$  at spatial infinity



$$Q = \int_{\partial D} d\sigma_\mu j^\mu = \int_{\sigma} d\sigma_\mu j^\mu - \int_{\sigma'} d\sigma_\mu j^\mu$$

special case  $\sigma$  = spacelike hyperplane at  $x^0 = t = \text{const.}$   
parametrized by  $x^1, x^2, x^3$

$$\Rightarrow d\sigma_\mu = \epsilon_{\mu 123} dx^1 dx^2 dx^3 = \epsilon_{\mu 123} dx^3 = (dx^3, \vec{0})$$

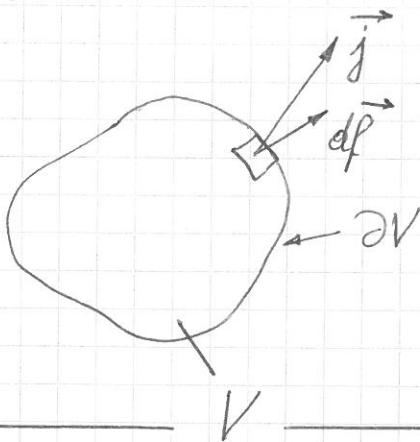
$$\Rightarrow Q = \int_{\sigma = \mathbb{R}^3} d^3x \underbrace{j^o(t, \vec{x})}_{\text{charge density}}$$

charge  $Q_V$  contained in a finite volume  $V \subset \mathbb{R}^3$

$$Q_V(t) = \int_V d^3x j^o(t, \vec{x})$$

$$\frac{d}{dt} Q_V(t) = \int_V d^3x \frac{\partial j^o}{\partial x^o} = - \int_V d^3x \vec{\nabla} \cdot \vec{j} = - \int_V d\vec{P} \cdot \vec{j}$$

$\vec{j}(x)$  = density of current



summary: Noether theorem in field theory

if action  $S$  invariant under one-parameter group of transformations  $x^\mu \rightarrow x'^\mu = X^\mu(x, \phi; \alpha)$ ,  $\phi_a \rightarrow \Phi_a(x, \phi; \alpha) \Rightarrow \exists$  conserved current

$$j^\mu = \frac{\partial \mathcal{L}}{\partial \dot{\phi}_a, \mu} \left. \frac{\partial \Phi_a}{\partial x} \right|_{x=0} - \Theta_\nu^\mu \left. \frac{\partial X^\nu}{\partial x} \right|_{x=0}$$

symmetry transformations usually not needed in full generality:

(i) (global) internal symmetries

$$x' = x, \quad \phi_a' = \phi_a'(\phi; \alpha)$$

(ii) spacetime symmetries

$$x'^\mu = x'^\mu(x; \alpha), \quad \phi_a' = \phi_a'(x, \phi; \alpha)$$

generalizations of the Noether theorem:

$$(a) \mathcal{L}' = \mathcal{L} + F_\alpha, \quad F = \frac{\partial f^\mu}{\partial x^\mu} + \frac{\partial f^\mu}{\partial \phi_a} \phi_{a,\mu} \quad \begin{matrix} \text{total} \\ \text{divergence} \end{matrix}$$

$$f^\mu = f^\mu(x, \phi)$$

$$\Rightarrow j^\mu = \frac{\partial \mathcal{L}}{\partial \phi_{a,\mu}} \left. \frac{\partial \Phi_a}{\partial \alpha} \right|_{\alpha=0} - \Theta^\nu \left. \frac{\partial X^\nu}{\partial \alpha} \right|_{\alpha=0} + f^\mu$$

(b) local gauge transformations  $\alpha \rightarrow \alpha(x)$

## Applications to Poincaré-covariant field theories

Poincaré transformation  $x' = \overset{\downarrow}{L} \underset{\downarrow}{x} + \overset{\downarrow}{a}$ ,  $L^T g L = g$ ;  $\overset{4 \times 4 \text{ matrix}}{\quad} \in \mathbb{R}^4$

$$(g_{\mu\nu}) = \text{diag}(1, -1, -1, -1) = (g^{1\nu})$$

leaves the line element  $ds^2 = (dx^\circ)^2 - d\vec{x}^2 = dx'^\mu dx'^\nu g_{\mu\nu}$  invariant.

Poincaré group  $P$  consist of elements  $g = (L, a)$

composition law:

$$\begin{aligned} x'' &= L' x' + a' = L' (L x + a) + a' \\ &= L' L x + L' a + a' \end{aligned}$$

$$g' g = (L', a') (L, a) = (L' L, L' a + a')$$

Lorentz group  $L$  = subgroup of  $P$  consisting of the elements  $(L, 0)$

$$x' = L x \quad (\underline{\text{no shift of the origin}})$$

- full Lorentz group  $L$  contains also space and time reversals  $\rightarrow L$  is not connected

$$T = \begin{pmatrix} 1 & O_3^T \\ 0 & \mathbb{1}_3 \end{pmatrix}, \quad P = \begin{pmatrix} 1 & O_3^T \\ 0 & \mathbb{1}_3 - \mathbb{1}_3 \end{pmatrix}$$

$$\underbrace{\det(L^T g L)}_{\det L^T \det g \det L} = \det g$$

$$\underbrace{\det L}_{\det L}$$

-proper Lorentz group  $L_+$

$$\Rightarrow (\det L)^2 = 1 \Rightarrow \det L = \pm 1$$

remark:  $\det P = \det T = -1$

$$(L^T g L)_{00} = g_{00} = 1$$

$$(L^T)_0^\mu g_{\mu\nu} L_0^\nu = 1$$

$$L_0^\mu g_{\mu\nu} L_0^\nu = 1$$

$$(L_0^\circ)^2 - \sum_{i=1}^3 (L_i^\circ)^2 = 1$$

$$(L_0^\circ)^2 = 1 + \sum_{i=1}^3 (L_i^\circ)^2 \Rightarrow (L_0^\circ)^2 \geq 1$$

$$\Rightarrow L_0^\circ \geq 1 \quad (\text{sense of time unchanged} \rightarrow)$$

$\swarrow$  orthochronous Lorentz group  $L^1$

orthochronous transformation)

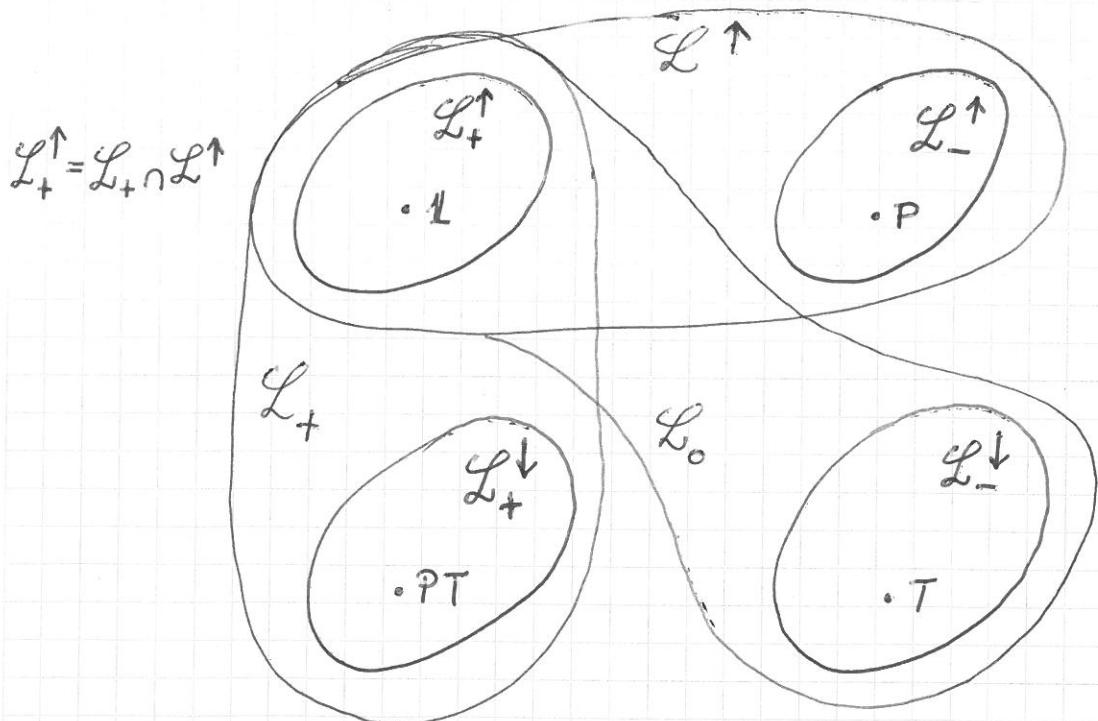
or

$L_0^\circ \leq -1$  (time reversal  $\rightarrow$  antichronous transf.)

decomposition of  $L$  into four components (each connected)

	$\det L$	$\operatorname{sgn} L^o$	
$L_+^\uparrow$	1	1	1
$L_-^\uparrow$	-1	1	P
$L_+^\downarrow$	+1	-1	PT
$L_-^\downarrow$	-1	-1	T

$$L_-^\uparrow = P L_+^\uparrow, \quad L_+^\downarrow = PT L_+^\uparrow, \quad L_-^\downarrow = T L_+^\uparrow$$



$L_+^\uparrow$  = proper orthochronous Lorentz group

$L_+ = L_+^\uparrow \cup L_+^\downarrow = L_+^\uparrow \cup PT L_+^\uparrow$  = proper Lorentz group

$L^\uparrow = L_+^\uparrow \cup L_-^\uparrow = L_+^\uparrow \cup PL_+^\uparrow$  = orthochronous Lorentz group

$L^o = L_+^\uparrow \cup L_-^\downarrow = L_+^\uparrow \cup TL_+^\uparrow$  = orthochronous -II- -II-

$L_-^\uparrow, L_+^\downarrow, L_-^\downarrow$  do not form groups!

fundamental interactions (strong, weak, electromagnetic, gravity) respect  $P_+^\uparrow$ ; weak interaction violates P, T

remark:  $P_+^\uparrow$  describes transformations from one inertial frame to another one (elements can be continuously connected with 1)

$(L, a) \in P_+^\uparrow$  can be written in the form

$$\underbrace{(L(\vec{x})}_{\text{rotation}}, \underbrace{L(\vec{u})}_{\substack{\text{Boost} \\ (\text{velocity transf.})}}, a)$$

$L_+^\uparrow$  has 6 parameters  $(\vec{x}, \vec{u})$

$$P_+^\uparrow \dashv \dashv 10 \dashv (\vec{x}, \vec{u}, a)$$

Poincaré invariance  $\rightarrow$  10 constants of motion

$$S = \underbrace{\int d^4x}_{\text{inv.}} \mathcal{L}(\phi, \phi_\mu)$$

$S$   $P_+^\uparrow$ -invariant if  $L$  transforms as a scalar (no explicit dependence on x)

a) translational invariance

$$x'^\mu = x^\mu + \underbrace{n^\mu}_{\Delta x^\mu} x$$

$$\Delta\varphi = \varphi'(x') - \varphi(x) = 0$$

$\Rightarrow$  conserved current  $j^\mu = -\Theta^\mu_\nu, n^\nu$

$$n^\nu \text{ arbitrary} \Rightarrow \partial_\mu \Theta^\mu_\nu = 0$$

$$P_\nu = \int_V d\sigma_\mu \Theta^\mu_\nu = \int_{t=\text{const.}}^3 d^3x \Theta^0_\nu = \text{const.}$$

$\Theta^\mu_\nu = \underline{\text{canonical energy-momentum tensor}}$

$P_\nu = \underline{\text{energy-momentum four-vector}}$

$$\frac{d}{dt} \int_V d^3x \Theta^{0\nu} = \int_V d^3x \underbrace{\frac{\partial}{\partial x^0} \Theta^{0\nu}}_{-\frac{\partial}{\partial x^i} \Theta^i_\nu} = - \int_V df_i \Theta^{i\nu}$$

finite domain  $\subset \mathbb{R}^3$

→ interpretation?

$\Theta^{00}$  = energy density-

$\Theta^{0i}$  = momentum density-

$\Theta^{i0}$  = energy current density-

$\Theta^{ij}$  = stress tensor density-

objection: conclusion integral → integrand

arbitrariness in the localization of energy, momentum,  
etc.

take arbitrary tensor field  $f^{\alpha\mu}_{\nu}$ , antisymmetric  
in  $\alpha, \mu$  and  $f^{\alpha\mu}_{\nu}(x) \xrightarrow[|\vec{x}| \rightarrow \infty]{} 0$

$$\Rightarrow \partial_\mu (\partial_\alpha f^{\alpha\mu}_{\nu}) = 0$$

$$\text{and } \int_{\Sigma} d\sigma_\mu \partial_\alpha f^{\alpha\mu}_{\nu} = \int_{\mathbb{R}^3} dx^\alpha \partial_\alpha f^{\alpha 0}_{\nu} = \int_{\mathbb{R}^3} dx^\alpha \partial_i f^{i0}_{\nu}$$

↑  
spacelike hypersurface

$$= \int_{\partial\mathbb{R}^3} d\Gamma_i f^{i0}_{\nu} = 0$$

$$\Rightarrow T^\mu{}_\nu := \Theta^\mu{}_\nu + \partial_\alpha f^{\alpha\mu}{}_\nu \quad \text{also divergence-free,}$$

same value for total energy-momentum  $P_\nu$

but: energy-momentum contained in a finite spatial volume depends on the choice of  $f^{\alpha\mu}{}_\nu$  (further arguments necessary to fix it)

### infinitesimal Lorentz transformation

$$\text{we know: } L^T g L = g \quad \underbrace{(L^T)^\alpha{}_\mu}_{\in \mathcal{L}_+^\uparrow} g_{\alpha\beta} L^\beta{}_\nu = g_{\mu\nu}$$

$$\text{infinitesimal transformation } L = 1 + \omega \quad L^\mu{}_\nu = \delta^\mu{}_\nu + \omega^\mu{}_\nu$$

$$(1 + \omega^\top) g (1 + \omega) = g + \omega^\top g + g \omega = g$$

$$\Rightarrow \omega^\top g + g \omega = 0$$

$$\omega^\alpha{}_\mu g_{\alpha\nu} + g_{\mu\alpha} \omega^\alpha{}_\nu = 0$$

$$\omega_{\nu\mu} + \omega_{\mu\nu} = 0 \quad (\omega_{\mu\nu} \text{ is } \underline{\text{antisymmetric}})$$

→ 6 real parameters  
(rotations and boosts)

$$x'^\mu = x^\mu + \omega^\mu{}_\nu x^\nu \Rightarrow \Delta x^\mu = \omega^\mu{}_\nu x^\nu$$

field transformation  $\phi_a'(x') = D_{ab}(L) \phi_b(x)$

$D(L)$  = (finite-dimensional) linear representation of  $L \in \mathcal{L}_+^\uparrow$   
(more precisely :  $SL(2, \mathbb{C})$ )  $D(L_1 L_2) = D(L_1) D(L_2)$

examples :

(i) scalar field  $\phi'(x') = \phi(x)$ ,  $D(L) = 1$   
(trivial repr.)

(ii) vector field  $A'^\mu(x') = L^\mu_\nu A^\nu(x)$ ,  $D(L) = L$   
(defining repr.)

infinitesimal transformation  $L = 1 + \omega$

$$\Rightarrow D(L) = 1 + \frac{1}{2} \omega_{\alpha\beta} \sum^{\alpha\beta} + O(\omega^2), \quad \sum^{\alpha\beta} = - \sum^{\beta\alpha}$$

$$D_{ab}(L) = \delta_{ab} + \frac{1}{2} \omega_{\alpha\beta} \sum^{\alpha\beta}_{ab} + O(\omega^2)$$

$$\phi_a'(x') = \phi_a(x) + \underbrace{\frac{1}{2} \omega_{\alpha\beta} \sum^{\alpha\beta}_{ab} \phi_b(x)}_{\text{skew variation } \Delta \phi_a}$$

for the previous examples :

(i)  $\sum = 0$

(ii)  $A'^\mu(x') = (\delta^\mu_\nu + \omega^\mu_\nu) A^\nu(x) =$   
 $= A^\mu(x) + \frac{1}{2} \omega_{\alpha\beta} (g^{\alpha\mu} g^{\beta\nu} - g^{\beta\mu} g^{\alpha\nu}) A^\nu(x)$

$$A'_\mu(x') = A_\mu(x) + \frac{1}{2} \omega_{\alpha\beta} \underbrace{\left( \delta_\mu^\alpha \delta_\nu^\beta - \delta_\mu^\beta \delta_\nu^\alpha \right)}_{\sum_{\mu\nu}^{\alpha\beta}} A^\nu(x)$$

divergence-free current:

$$j^\mu = \frac{\partial \mathcal{L}}{\partial \dot{\phi}_{\alpha\mu}} \frac{1}{2} \omega_{\alpha\beta} \sum_{ab}^{\alpha\beta} \dot{\phi}_b - \Theta^{\mu\alpha} \omega_{\alpha\beta} x^\beta$$

$$= \frac{1}{2} \left[ \frac{\partial \mathcal{L}}{\partial \dot{\phi}_{\alpha\mu}} \sum_{ab}^{\alpha\beta} \dot{\phi}_b - (\Theta^{\mu\alpha} x^\beta - \Theta^{\mu\beta} x^\alpha) \right] \omega_{\alpha\beta}$$

$\omega_{\alpha\beta}$  arbitrary

$$\Rightarrow \partial_\mu \underbrace{\left[ \frac{\partial \mathcal{L}}{\partial \dot{\phi}_{\alpha\mu}} \sum_{ab}^{\alpha\beta} \dot{\phi}_b - (\Theta^{\mu\alpha} x^\beta - \Theta^{\mu\beta} x^\alpha) \right]}_{m^{\mu\alpha\beta}} = 0$$

$$M^{\alpha\beta} = \int d\sigma_\mu m^{\mu\alpha\beta} = \int_{\mathbb{R}^3} d^3x m^{\alpha\beta} = \text{const.}$$

$$M^{\alpha\beta} = \int d^3x \left( x^\alpha \Theta^0\beta - x^\beta \Theta^0\alpha + \underbrace{\frac{\partial \mathcal{L}}{\partial \dot{\phi}_a}}_{Tl_a} \sum_{ab}^{\alpha\beta} \dot{\phi}_b \right)$$

↑      ↗  
energy-momentum density      spin contribution

scalar field:  $\sum^{\alpha\beta} = 0$

$$\partial_\mu (x^\alpha \Theta^{\mu\beta} - x^\beta \Theta^{\mu\alpha}) = 0$$

$$\Rightarrow \Theta^{\alpha\beta} - \Theta^{\beta\alpha} = 0$$

$\Rightarrow$  the canonical energy-momentum tensor of a scalar field is symmetric

general case:

$$\partial_\mu \mathcal{M}^{\mu\alpha\beta} = 0 \Rightarrow \Theta^{\alpha\beta} - \Theta^{\beta\alpha} + \partial_\mu \left( \frac{\partial L}{\partial \phi_{\alpha,\mu}} \sum_{ab}^{\alpha\beta} \phi_b \right) = 0$$

We want to construct a symmetric energy-momentum tensor  $T^{\mu\nu} = \Theta^{\mu\nu} + \partial_\alpha f^{\alpha\mu\nu}$  by an appropriate choice of  $f^{\alpha\mu\nu} = -f^{\mu\nu\alpha}$

$$\begin{aligned} 0 &= T^{\mu\nu} - T^{\nu\mu} = \Theta^{\mu\nu} + \partial_\alpha f^{\alpha\mu\nu} - \Theta^{\nu\mu} - \partial_\alpha f^{\alpha\nu\mu} \\ &= \partial_\alpha (f^{\alpha\mu\nu} - f^{\alpha\nu\mu}) - \partial_\alpha \left( \frac{\partial L}{\partial \phi_{\alpha,\mu}} \sum_{ab}^{\mu\nu} \phi_b \right) \end{aligned}$$

$$\Rightarrow \partial_\alpha (f^{\alpha\mu\nu} - f^{\alpha\nu\mu}) = \partial_\alpha \left( \frac{\partial \mathcal{L}}{\partial \phi_{\alpha,\alpha}} \sum_{ab}^{\mu\nu} \phi_b \right)$$

$$\Rightarrow f^{\alpha\mu\nu} - f^{\alpha\nu\mu} = \underbrace{\frac{\partial \mathcal{L}}{\partial \phi_{\alpha,\alpha}} \sum_{ab}^{\mu\nu} \phi_b}_{=: g^{\alpha\mu\nu}} + \partial_\alpha h^{\sigma\alpha\mu\nu}$$

with  $h^{\sigma\alpha\mu\nu} = -h^{\alpha\sigma\mu\nu} = -h^{\sigma\alpha\nu\mu}$

unique solution for  $f^{\alpha\mu\nu}$ :

$$f^{\alpha\mu\nu} - f^{\alpha\nu\mu} = g^{\alpha\mu\nu}$$

$$\underbrace{f^{\mu\nu\alpha}}_{f^{\alpha\mu\nu}} - \underbrace{f^{\mu\alpha\nu}}_{f^{\mu\nu\alpha}} = g^{\mu\nu\alpha}$$

$$- \underbrace{(f^{\nu\alpha\mu} - f^{\nu\mu\alpha})}_{-f^{\alpha\nu\mu}} = -g^{\nu\alpha\mu}$$

$$\Rightarrow 2f^{\alpha\mu\nu} = g^{\alpha\mu\nu} + g^{\mu\nu\alpha} - g^{\nu\alpha\mu}$$

$$f^{\alpha\mu\nu} = \frac{1}{2} (g^{\alpha\mu\nu} + g^{\mu\nu\alpha} - g^{\nu\alpha\mu})$$

$$f^{\alpha\mu\nu} = \frac{1}{2} \left\{ \frac{\partial \mathcal{L}}{\partial \phi_{a,\alpha}} \sum_{ab}^{\mu\nu} \phi_b + \frac{\partial \mathcal{L}}{\partial \phi_{a,\mu}} \sum_{ab}^{\nu\alpha} \phi_b - \right. \\ \left. - \frac{\partial \mathcal{L}}{\partial \phi_{a,\nu}} \sum_{ab}^{\alpha\mu} \phi_b + \partial_\sigma (h^{\sigma\alpha\mu\nu} + h^{\sigma\mu\nu\alpha} - h^{\sigma\nu\alpha\mu}) \right\}$$

possible choice:  $h^{\sigma\alpha\mu\nu} = 0$

$$\Rightarrow T^{\mu\nu} = \Theta^{\mu\nu} + \partial_\alpha f^{\alpha\mu\nu} = \\ = \Theta^{\mu\nu} + \frac{1}{2} \partial_\alpha \left\{ \frac{\partial \mathcal{L}}{\partial \phi_{a,\alpha}} \sum_{ab}^{\mu\nu} \phi_b + \frac{\partial \mathcal{L}}{\partial \phi_{a,\mu}} \sum_{ab}^{\nu\alpha} \phi_b \right. \\ \left. - \frac{\partial \mathcal{L}}{\partial \phi_{a,\nu}} \sum_{ab}^{\alpha\mu} \phi_b \right\} \quad (\text{with } h=0)$$

$$J^{\mu\alpha\beta} = x^\alpha T^{\mu\beta} - x^\beta T^{\mu\alpha} = \\ = x^\alpha \Theta^{\mu\beta} - x^\beta \Theta^{\mu\alpha} + x^\alpha \partial_\sigma f^{\sigma\mu\beta} - x^\beta \partial_\sigma f^{\sigma\mu\alpha} \\ = x^\alpha \Theta^{\mu\beta} - x^\beta \Theta^{\mu\alpha} + \partial_\sigma (x^\alpha f^{\sigma\mu\beta} - x^\beta f^{\sigma\mu\alpha}) \\ - \underbrace{f^{\alpha\mu\beta} + f^{\beta\mu\alpha}}_{f^{\mu\alpha\beta} - f^{\mu\beta\alpha}} = \frac{\partial \mathcal{L}}{\partial \phi_{a,\mu}} \sum_{ab}^{\alpha\beta} \phi_b + \partial_\sigma h^{\sigma\mu\alpha\beta}$$

$$J^{\mu\alpha\beta} = \overbrace{\frac{\partial L}{\partial \dot{\Phi}_{\alpha\mu}} \sum_{ab}^{\alpha\beta} + x^\alpha \Theta^{\mu\beta} - x^\beta \Theta^{\mu\alpha}}$$

$$+ \partial_\sigma (x^\alpha f^{\sigma\mu\beta} - x^\beta f^{\sigma\mu\alpha} + h^{\sigma\mu\beta})$$

$$\Rightarrow \int d\sigma_\mu J^{\mu\alpha\beta} = \int d\sigma_\mu m^{\mu\alpha\beta}$$

choice of  $h^{\sigma\mu\beta}$  → affects only localization of angular momentum

"correct" choice of  $h^{\sigma\mu\beta}$ ?

where does a localization of field energy and field momentum play a rôle? → general relativity: energy-momentum tensor acts as the source of the gravitational field → symmetric  $T^{\mu\nu}$  with  $h^{\sigma\mu\beta} = 0$  (Belinfante)

## What the graviton listens to

$$SS_{\text{matter}} = -\frac{1}{2} \int d^4x \sqrt{-g} T^{\mu\nu} \delta g_{\mu\nu} =$$

$$= \frac{1}{2} \int d^4x \sqrt{-g} T_{\mu\nu} \delta g^{\mu\nu}$$

$$g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}, \quad g^{\mu\nu} = \eta^{\mu\nu} - h^{\mu\nu}$$

$$-g = 1 + \eta^{\mu\nu} h_{\mu\nu}, \quad \sqrt{-g} = 1 + \frac{1}{2} \eta^{\mu\nu} h_{\mu\nu}$$

$$S_{\text{matter}} = \int d^4x \sqrt{-g} \underbrace{(A + g^{\mu\nu} B_{\mu\nu} + g^{\mu\nu} g^{\lambda\rho} C_{\mu\nu\lambda\rho} + \dots)}_{\mathcal{L}} \quad C_{\mu\nu\lambda\rho} = C_{\lambda\rho\mu\nu}$$

$$\Rightarrow T_{\mu\nu} = 2(B_{\mu\nu} + 2C_{\mu\nu\lambda\rho}\eta^{\lambda\rho} + \dots) - \eta_{\mu\nu} \mathcal{L}$$

remark :  $T^\mu_\mu = -(4A + 2\eta^{\mu\nu} B_{\mu\nu})$  no contribution from  $C_{\mu\nu\lambda\rho}$

example : massive spin 1 field

$$\mathcal{L} = -\frac{1}{4} g^{\mu\nu} g^{\lambda\rho} F_{\mu\lambda} F_{\nu\rho} + \frac{m^2}{2} g^{\mu\nu} A_\mu A_\nu$$

$$\Rightarrow T_{\mu\nu} = -F_{\mu\lambda} F_\nu{}^\lambda + m^2 A_\mu A_\nu - \eta_{\mu\nu} \mathcal{L}$$

symmetric energy-momentum tensor  $T^{\mu\nu} = T^{\nu\mu}$ ,  $\partial_\mu T^{\mu\nu} = 0$   
 correctly localizing energy and momentum

$$J^{\mu\alpha\beta} = x^\alpha T^{\mu\beta} - x^\beta T^{\mu\alpha}, \quad \partial_\mu J^{\mu\alpha\beta} = 0$$

$$\Rightarrow J^{\alpha\beta} = \int_{\sigma} d\sigma_\mu J^{\mu\alpha\beta} = \int_{\sigma} d\sigma_\mu (x^\alpha T^{\mu\beta} - x^\beta T^{\mu\alpha}) \\ = \int_{\mathbb{R}^3} d^3x (x^\alpha T^{0\beta} - x^\beta T^{0\alpha}) = \text{const.}$$

$\Rightarrow$  density of angular momentum:

$$J_k := \frac{1}{2} \varepsilon_{kij} J^{0ij} = \varepsilon_{kij} \underbrace{x^i T^{0j}}_{\text{momentum density}}$$

$\mathcal{E} := T^{00}$  energy density

$\mathcal{P}^i := T^{0i}$  momentum density

$J_k = \varepsilon_{kij} x^i \mathcal{P}^j$  angular momentum density

total angular momentum  $\vec{J} = \int_{\mathbb{R}^3} d^3x \vec{x} \times \vec{\mathcal{P}}$

$$\mathbb{J}^{i0} = \int_{\mathbb{R}^3} d^3x (x^i \mathcal{E} - x^0 p^i) = \text{const.}$$

dividing by the total energy  $E = P^0 = \int_{\mathbb{R}^3} d^3x \mathcal{E} \Rightarrow$

$$\frac{\int_{\mathbb{R}^3} d^3x \vec{x} \cdot \mathcal{E}}{\int_{\mathbb{R}^3} d^3x \mathcal{E}} = \frac{\vec{P}}{E} t + \underbrace{\vec{\alpha}_0}_{\text{const.}}$$

center of mass-energy (centroid) moves uniformly and rectilinearly with velocity  $\vec{P}/E$  with respect to the chosen reference frame

remark: position of the world line of ~~a~~ centroid depends on the inertial frame used for its definition; only its direction is uniquely given by the 4-momentum  $P^\mu$

if  $P^\mu P_\mu > 0$  ( $P^\mu$  timelike)  $\rightarrow$  restframe

with  $\vec{P}=0$  exists  $\rightarrow$  centroid with respect

to rest frame = relativistic center of mass

angular momentum with respect to rest frame = spin

assume:  $P^\mu$  timelike ( $P^2 > 0$ )

Behaviour of  $P^\mu$ ,  $J^{\mu\nu}$  under translations:  $P^\mu$  unchanged ( $P^\mu$  is a genuine 4-vector), but

$$J^{\mu\nu} \rightarrow \bar{J}^{\mu\nu} = \int d^3x (\bar{x}^\mu T^{0\nu} - \bar{x}^\nu T^{0\mu}) \\ = J^{\mu\nu} - a^\mu P^\nu + a^\nu P^\mu$$

under  $x \rightarrow \bar{x} = x - a$

$\mu^\mu$  = 4-velocity of the inertial frame used to define the centroid (space-time split)

world line of centroid = set of points  $a$  with

$$\bar{J}^{\mu\nu} \mu_\nu = 0$$

→ rest frame of  $\mu^\mu$ :  $\mu = (1, \vec{0})$

$$\Rightarrow \bar{J}^{i0} = 0 = J^{i0} - a^i P^0 + a^0 P^i$$

$$\Rightarrow a^i = \underbrace{\frac{P^i}{P^0} a^0}_{\frac{P^i}{E} t} + \underbrace{\frac{J^{i0}}{P^0}}_{a^i} \quad \checkmark$$

four-dimensional version :

$$O = \bar{J}^{\mu\nu} u_\nu = J^{\mu\nu} u_\nu - a^\mu P_\nu u + P^\mu a_\nu$$

$$\Rightarrow a^\mu = \frac{P^\mu}{P \cdot u} \underbrace{u \cdot u}_{\lambda} + \frac{J^{\mu\nu} u_\nu}{P \cdot u}$$

relativistic center of mass : choose reference frame

$$\text{with } u = P / \sqrt{P^2}$$

$\Rightarrow$  world line of relativistic center of mass :

$$a^\mu = \frac{P^\mu}{\sqrt{P^2}} \lambda + \frac{J^{\mu\nu} P_\nu}{P^2} \quad .$$

( $\lambda$  = proper time of relativistic center of mass )

spin tensor  $S^{\mu\nu}$  = tensor of angular momentum

with respect to the center of mass world line

$S^{\mu\nu}$  satisfies

$$S^{\mu\nu} P_\nu = 0 \quad (u^\mu \sim P^\mu)$$

$S^\mu$  contains the same information as the  
relativistic spin vector

$$S_\mu := \frac{1}{2} \epsilon_{\mu\alpha\beta\gamma} S^{\alpha\beta} P^\gamma = \frac{1}{2} \epsilon_{\mu\alpha\beta\gamma} J^{\alpha\beta} P^\gamma$$

the spin tensor  $S_{\alpha\beta}$  can be reconstructed from the spin 4-vector as -

$$\begin{aligned} S_{\alpha\beta} &= -\epsilon_{\alpha\beta\mu\nu} S^\mu P^\nu / P^2 = \\ &= J_{\alpha\beta} + (P_\alpha J_{\beta\gamma} - P_\beta J_{\alpha\gamma}) P^\gamma / P^2 \end{aligned}$$

remark: this formula can also be obtained from

$$S^{\mu\nu} = J^{\mu\nu} - \alpha^\mu P^\nu + \alpha^\nu P^\mu$$

$$\text{with } \alpha^\mu = \frac{P^\mu}{P^2} \gamma_0 + \frac{J^{\mu\nu} P_\nu}{P^2}$$

$S_\mu$  is orthogonal to  $P^\mu$ :

$$S_\mu P^\mu = 0$$

$$S_\mu \text{ is spacelike: } S_\mu S^\mu = -\frac{1}{2} S_{\mu\nu} S^{\mu\nu} \underbrace{P^2}_{>0} < 0$$