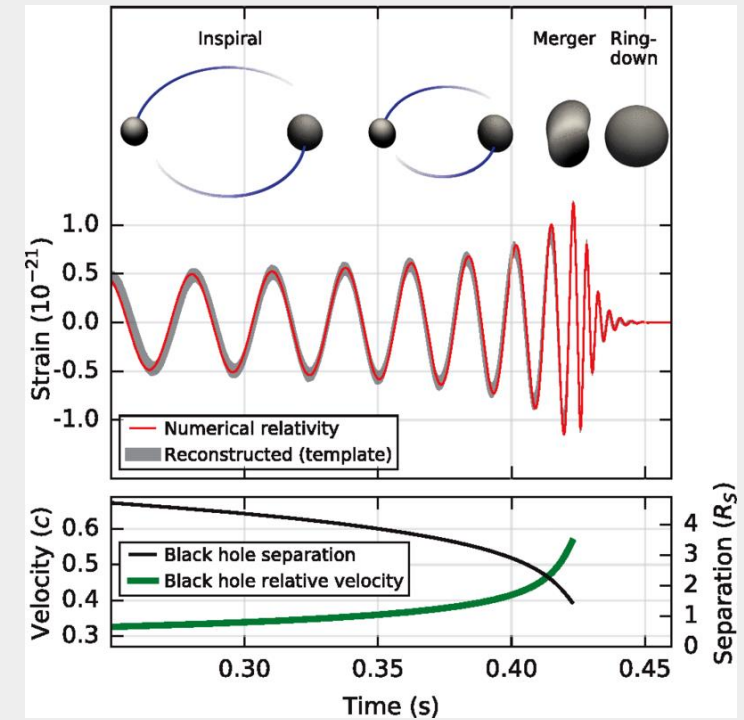
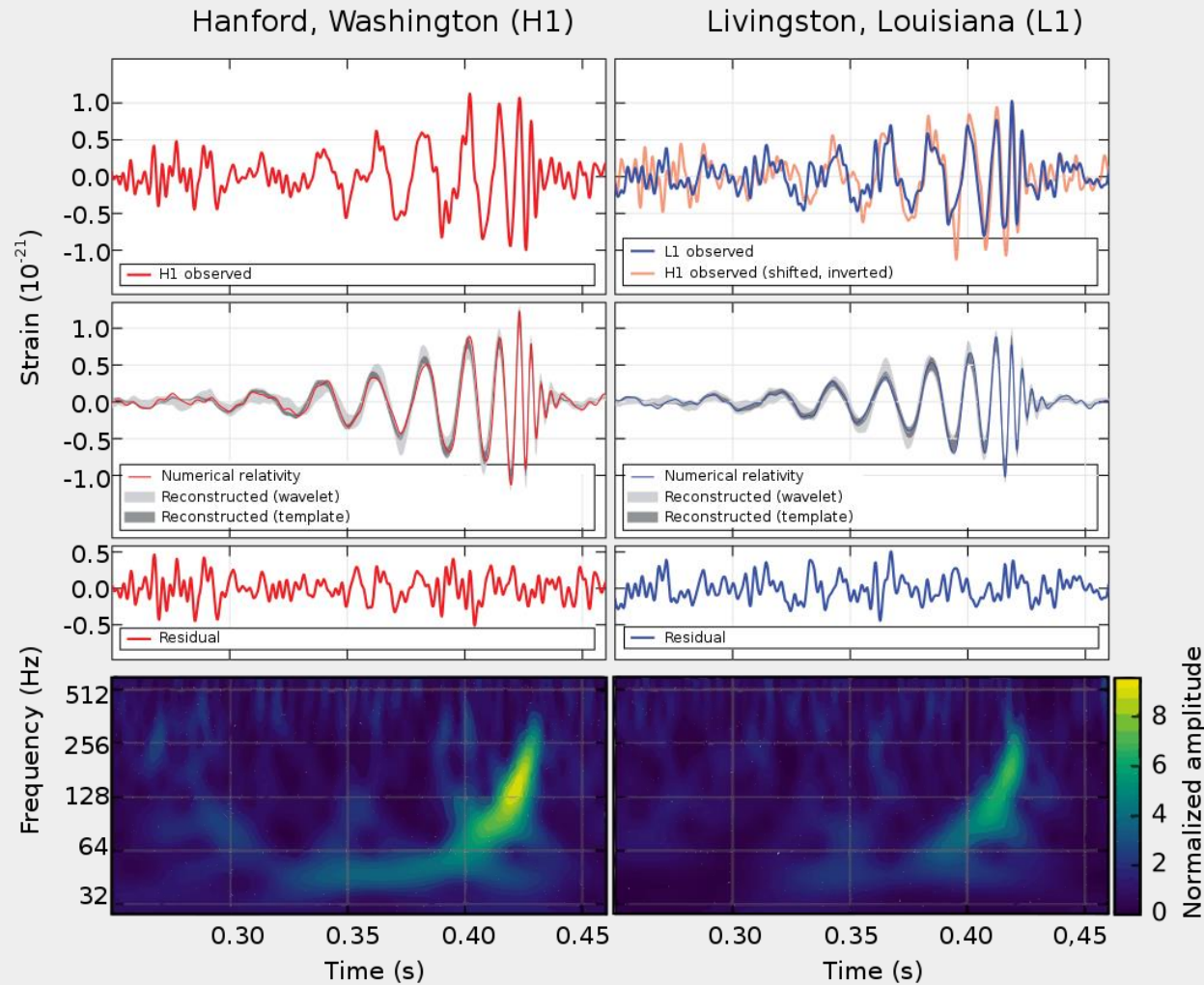


Advances in gravitational-wave predictions through quantum field theory methods

Outline

- Background
 - First gravitational wave detections
 - Detectors (2G and 3G)
 - Theory approaches
- Analytic methods
 - Terminology
 - Computation
 - Results

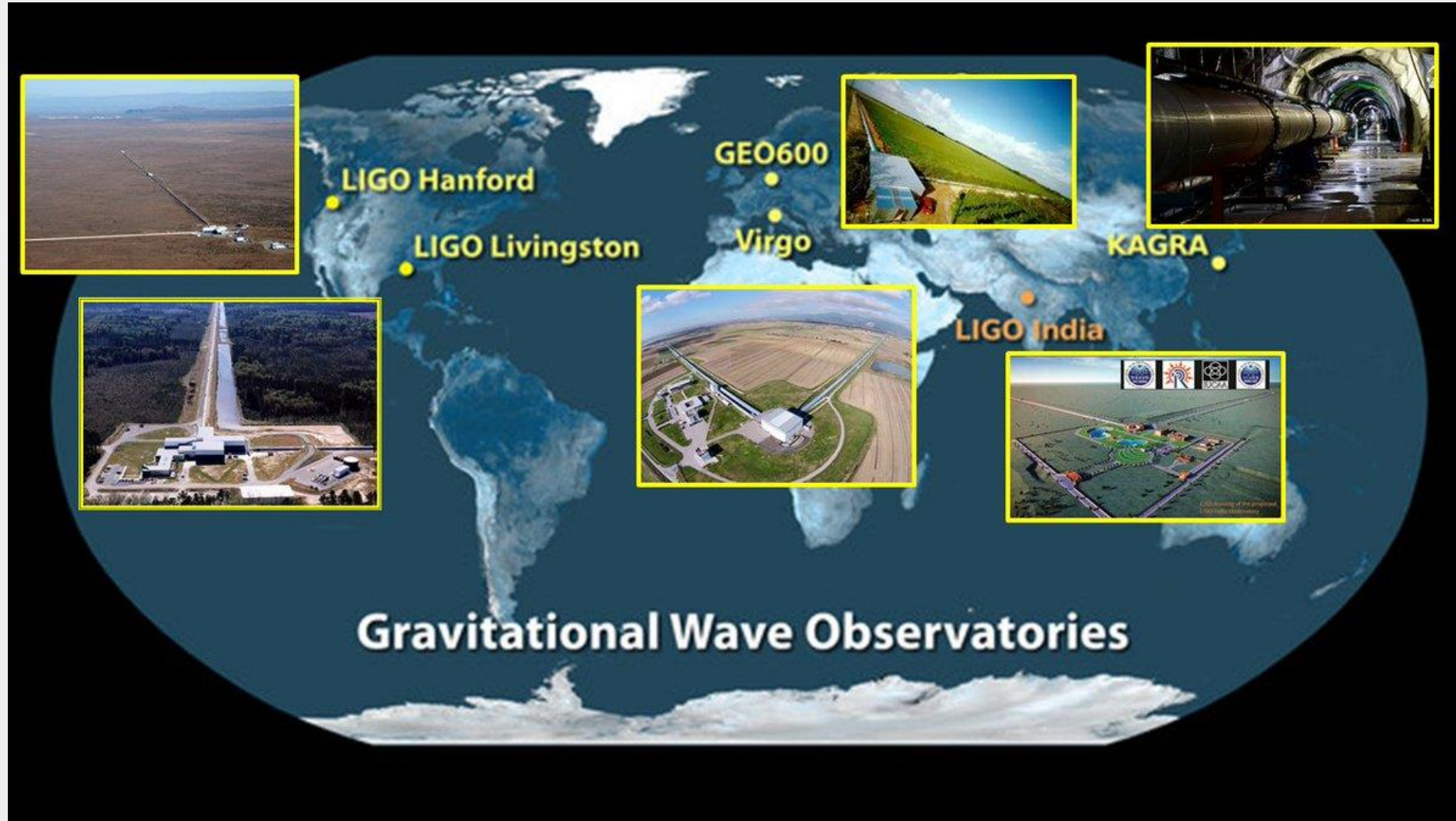
First detected gravitational wave



[LIGO]

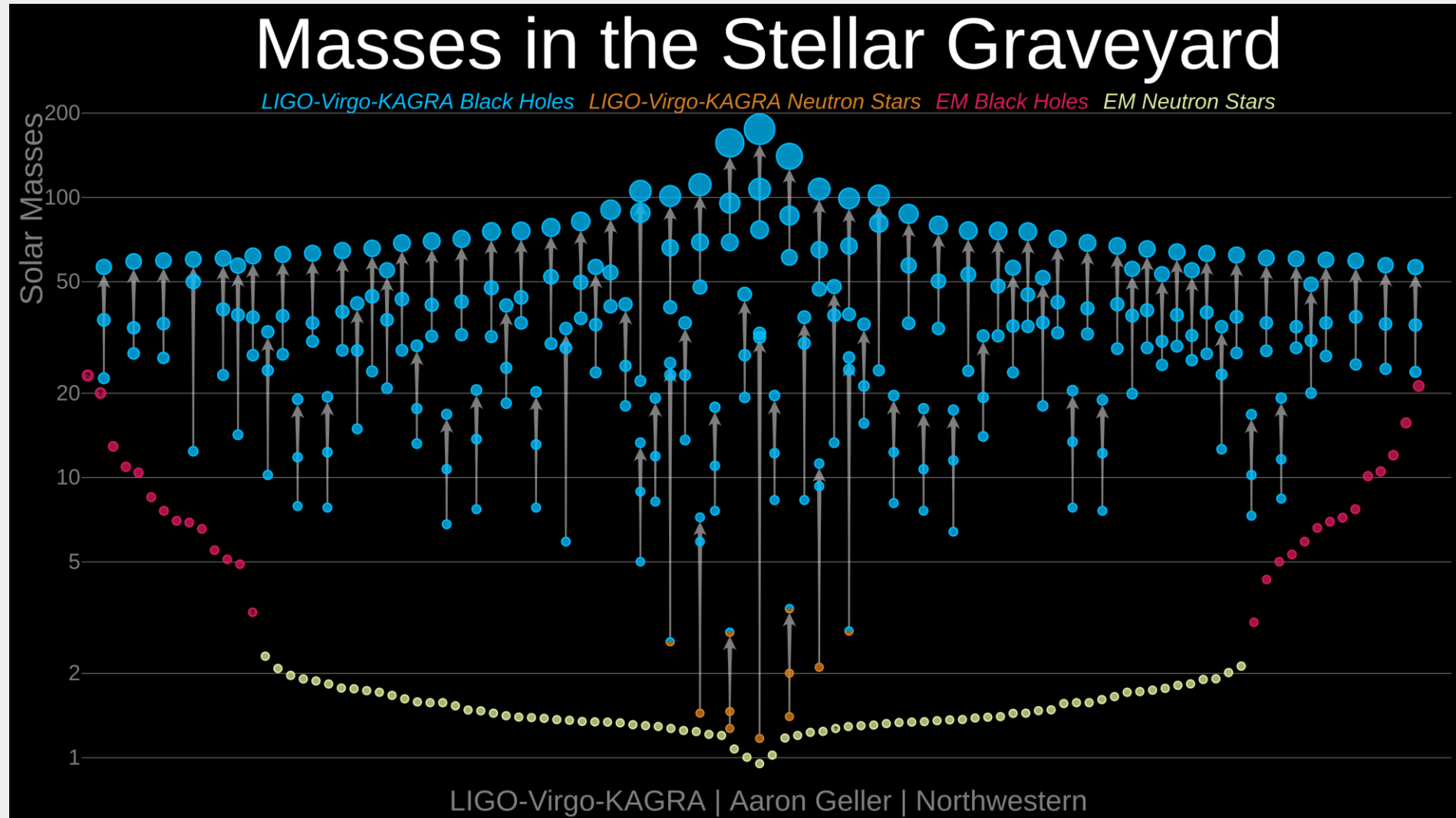
2017 Nobel Prize in Physics:
Weiss, Barish and Thorne

Global Network of 2G detectors



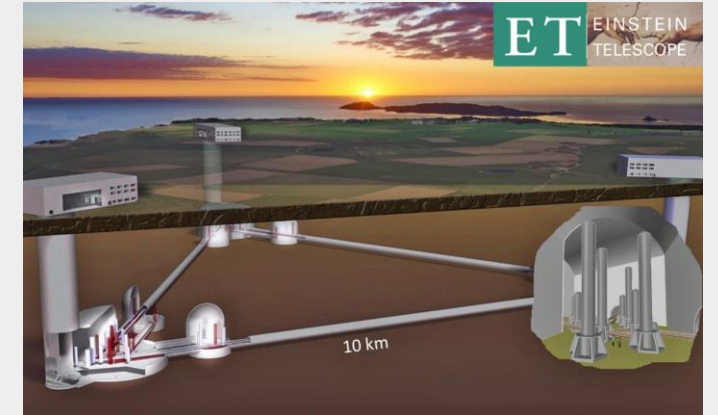
[LIGO]

Detections since then

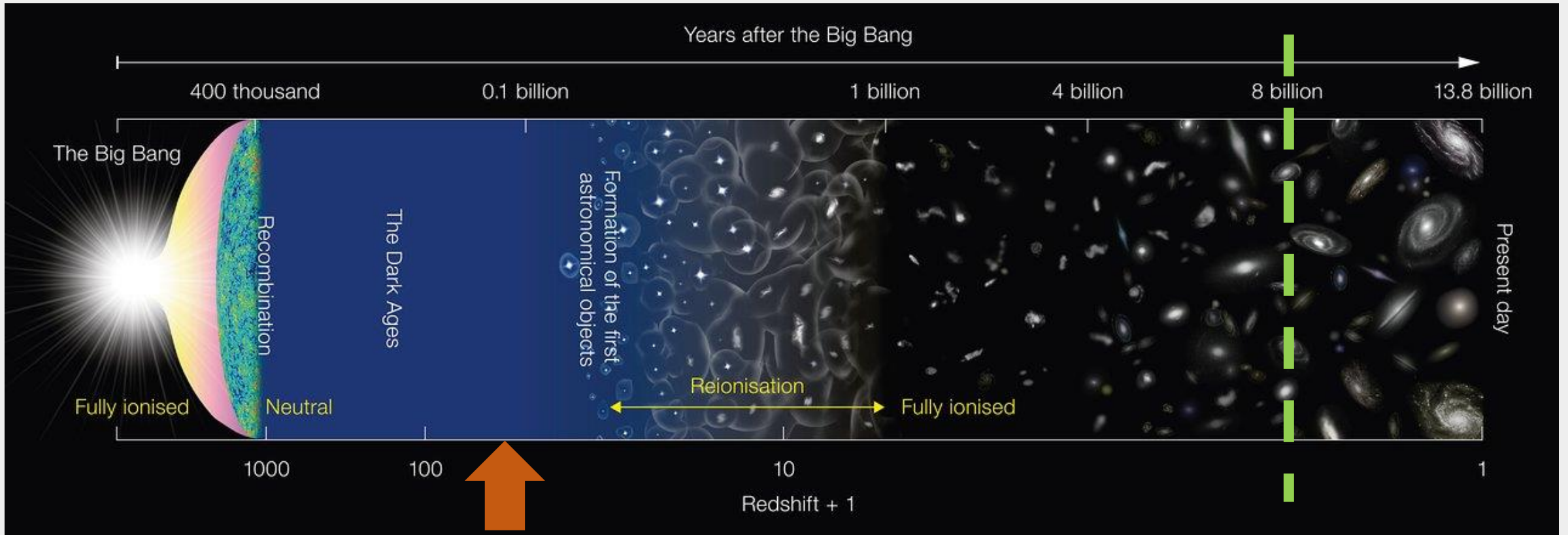


Proposed third-generation detectors

- Einstein Telescope (ET)
 - Belgium-Netherlands-Germany border
 - Sardinia (Italy)
 - underground, triangular shape, 10 km
- Cosmic Explorer (CE)
 - US
 - above-ground, L-shaped, 40 km + 20 km
- Laser Interferometer Space Antenna (LISA)
 - space-based (ESA), heliocentric orbit
 - triangular shape, 2.5 million km



Probing the Universe



3G Target

2G

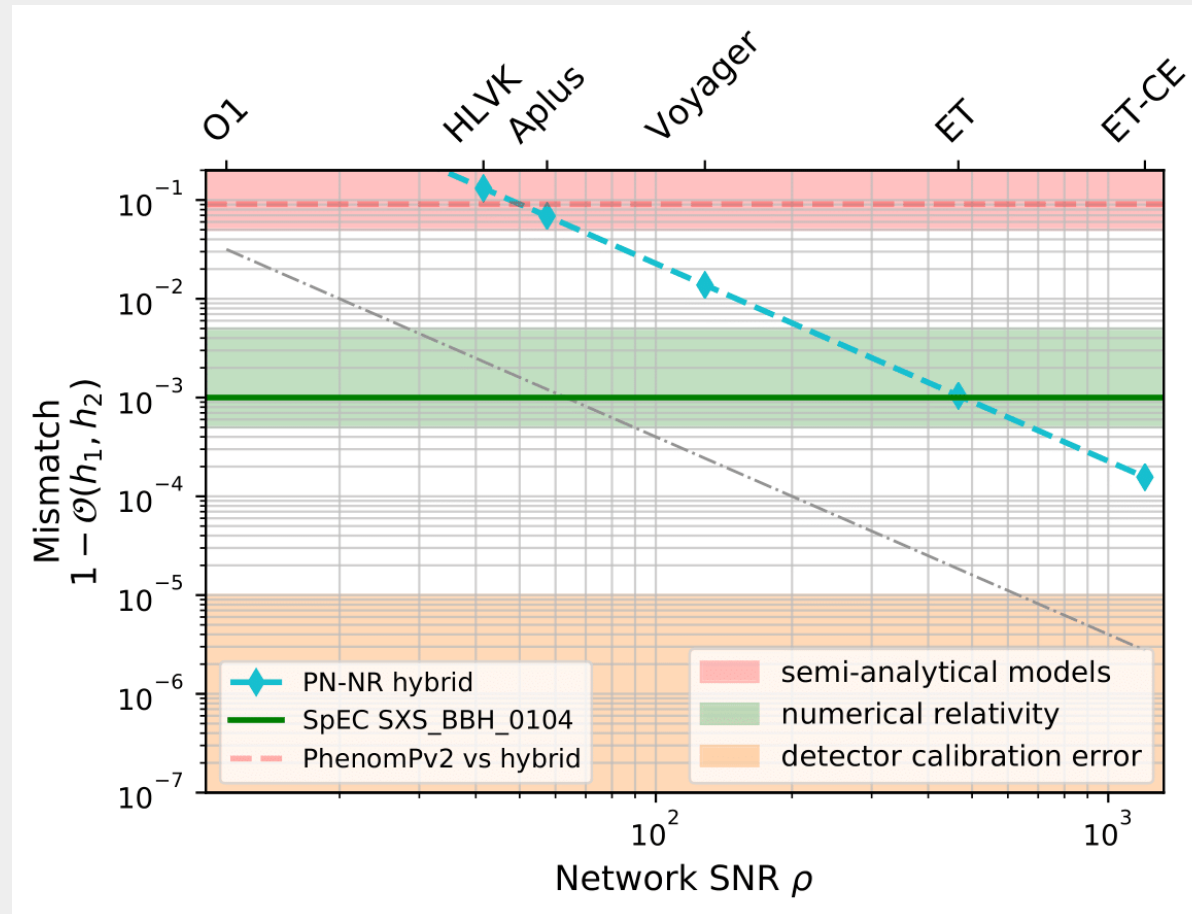
[ALMA collaboration]

Prospects

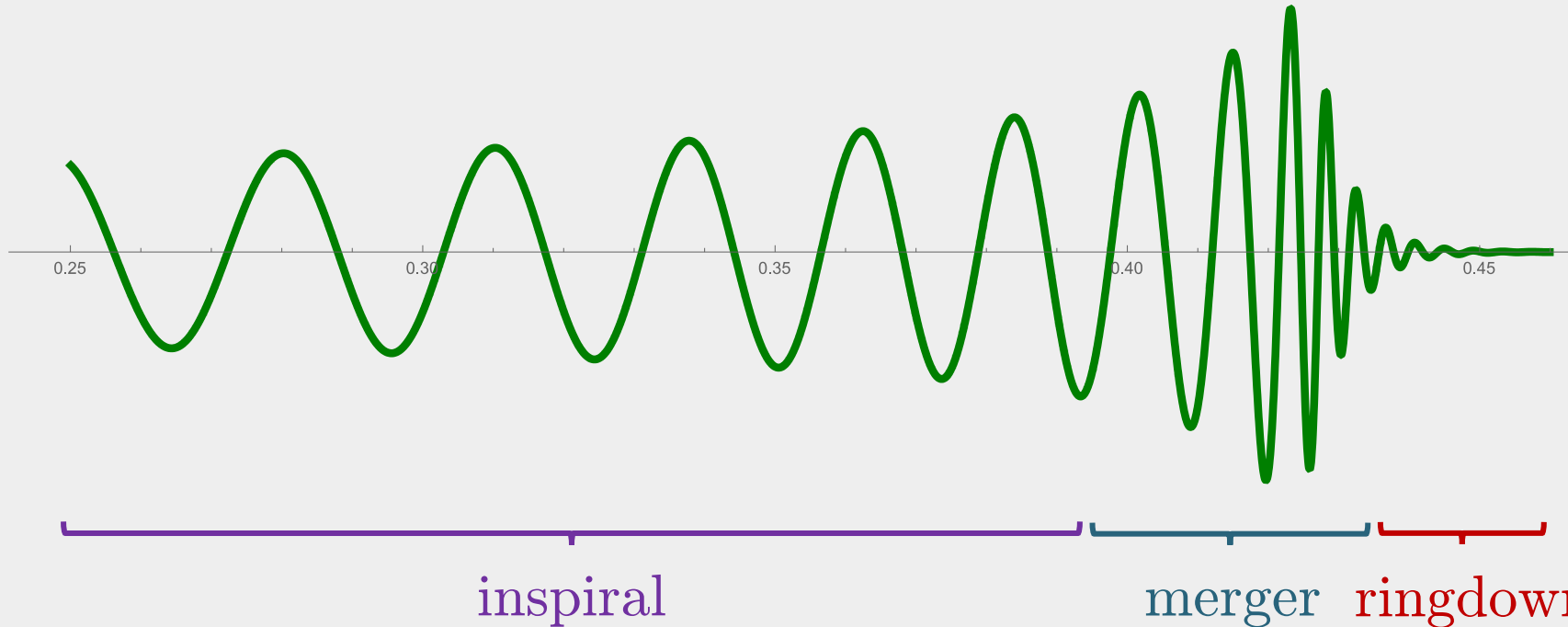
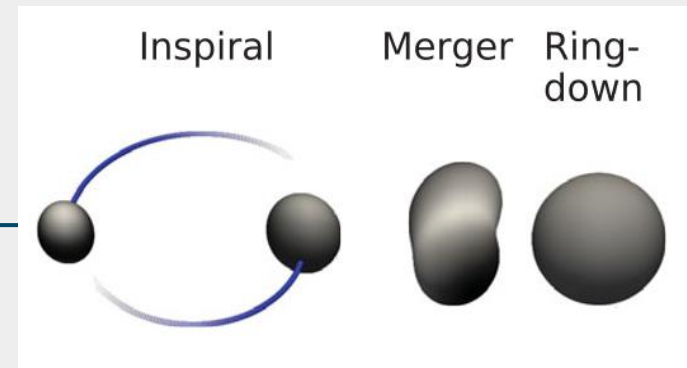
- Black hole properties
- Neutron star properties
- New astrophysical sources
- Dark matter (e.g. accreting on compact objects)
- Testing Einstein gravity (modified GW propagation)

3G Detectors: Higher accuracy needed

[Pürrer, Haster, '20]



Gravitational two-body problem



- **Inspiral:** analytic models for parameter estimation studies (PES)
- **Merger:** numerical relativity (NR)
- **Ringdown:** BH perturbation theory

Effective one-body (EOB) theory for inspiral-merger-ringdown waveforms

Terminology

non-spinning



spinning

black hole



neutron star

Post Newtonian (PN)



Post Minkowskian (PM)

bound (elliptic)



unbound (hyperbolic)

conservative



dissipative

potential modes



radiation modes

Terminology

this talk



non-spinning
black hole



spinning



neutron star



bookkeeping
in integrand

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PN vs PM

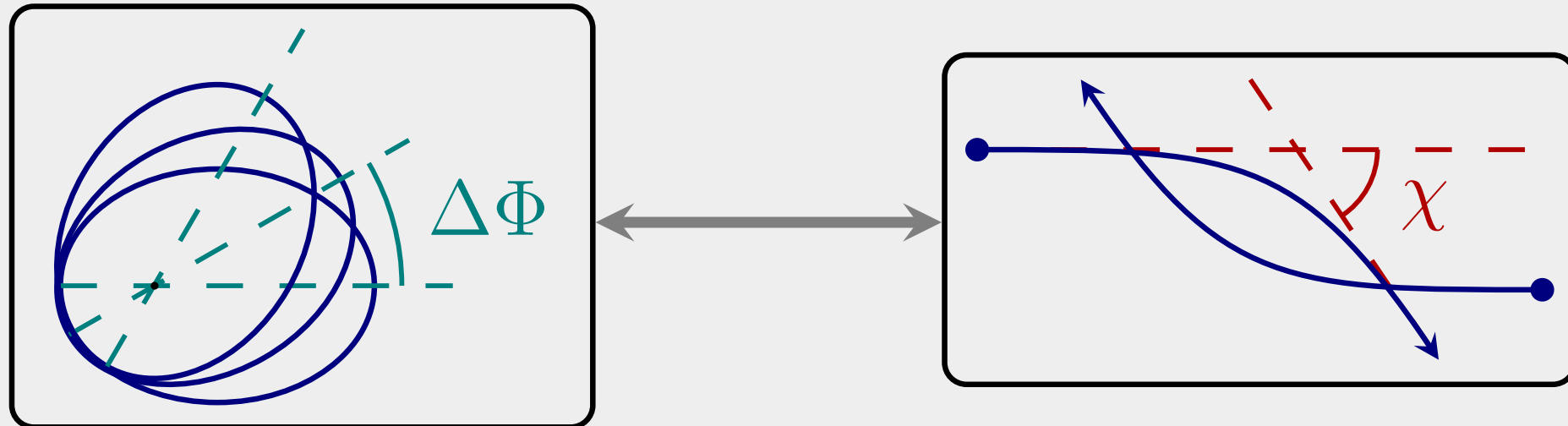
- Post-Newtonian (PN): $(v^2 \ll 1)$
 - non-relativistic expansion
 - integrals are non-relativistic (3d)
 - is simultaneously an expansion in the coupling

PN vs PM

- Post-Newtonian (PN): $(v^2 \ll 1)$
 - non-relativistic expansion
 - integrals are non-relativistic (3d)
 - is simultaneously an expansion in the coupling
- Post-Minkowskian (PM): $G \ll 1$
 - expansion in coupling
 - natural expansion for pQFT tools
 - incorporates all orders in velocities (more general)
 - results are naturally for scattering events

Bound vs Unbound motion

[Kälin, Porto, '20]



$\mathcal{E} > 0$

analytic continuation
in binding energy

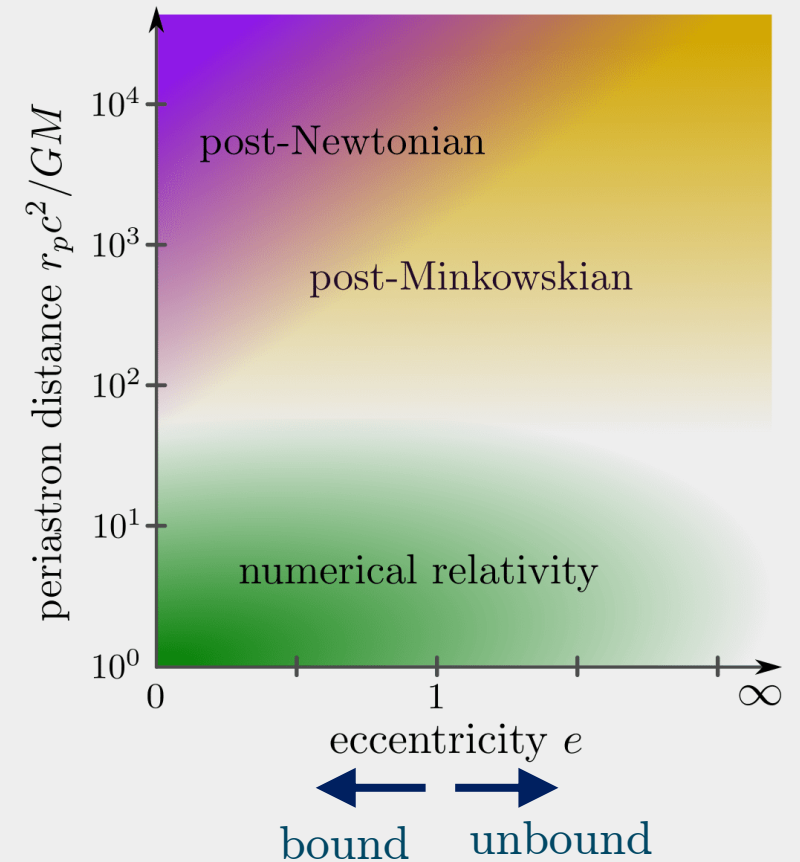
$\mathcal{E} < 0$

Eccentricity e



Post-Minkowskian (PM) expansion

- PM incorporates PN ($v^2 \ll 1$)
- Cleaner results: understand structure of observables, e.g. singularities, symmetries,...
- PM velocity resummation increases accuracy for scattering-like events
 - detection rate is expected to increase in 3G detectors
 - [Khalil, Buonanno, Steinhoff, Vines, '22; Damour, Rettegno, '23]



[Khalil, Buonanno, Steinhoff, Vines, '22]

PN vs PM



	0PN	1PN	2PN	3PN	4PN	5PN	6PN	
1PM	1	v^2	v^4	v^6	v^8	v^{10}	v^{12}	$\times G^1$
2PM		1	v^2	v^4	v^6	v^8	v^{10}	$\times G^2$
3PM			1	v^2	v^4	v^6	v^8	$\times G^3$
4PM				1	v^2	v^4	v^6	$\times G^4$
5PM					1	v^2	v^4	$\times G^5$
6PM						1	v^2	$\times G^6$

• State of the art for non-spinning BH:

- 4PM conservative [Bern, Parra-Martinez, Roiban, Ruf, Shen, Solon, Zeng, '21; CD, Kälin, Liu, Porto, '21, ...]
- 4PM dissipative [CD, Kälin, Liu, Neef, Porto, '22, ...]
- 5PM conservative [Driesse et al, '26, ...]
- 5PN & 6PN (conservative): [Foffa, Mastrolia, Sturani, Sturm Bobadilla, '19; Blümlein, Maier, Marquard, '19; Bini, Damour, Geralico, Laporta, Mastrolia, '20]

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Terminology

this talk



non-spinning
black hole



spinning



neutron star



bookkeeping
in integrand

Post Newtonian (PN)



Post Minkowskian (4PM)

bound (elliptic)



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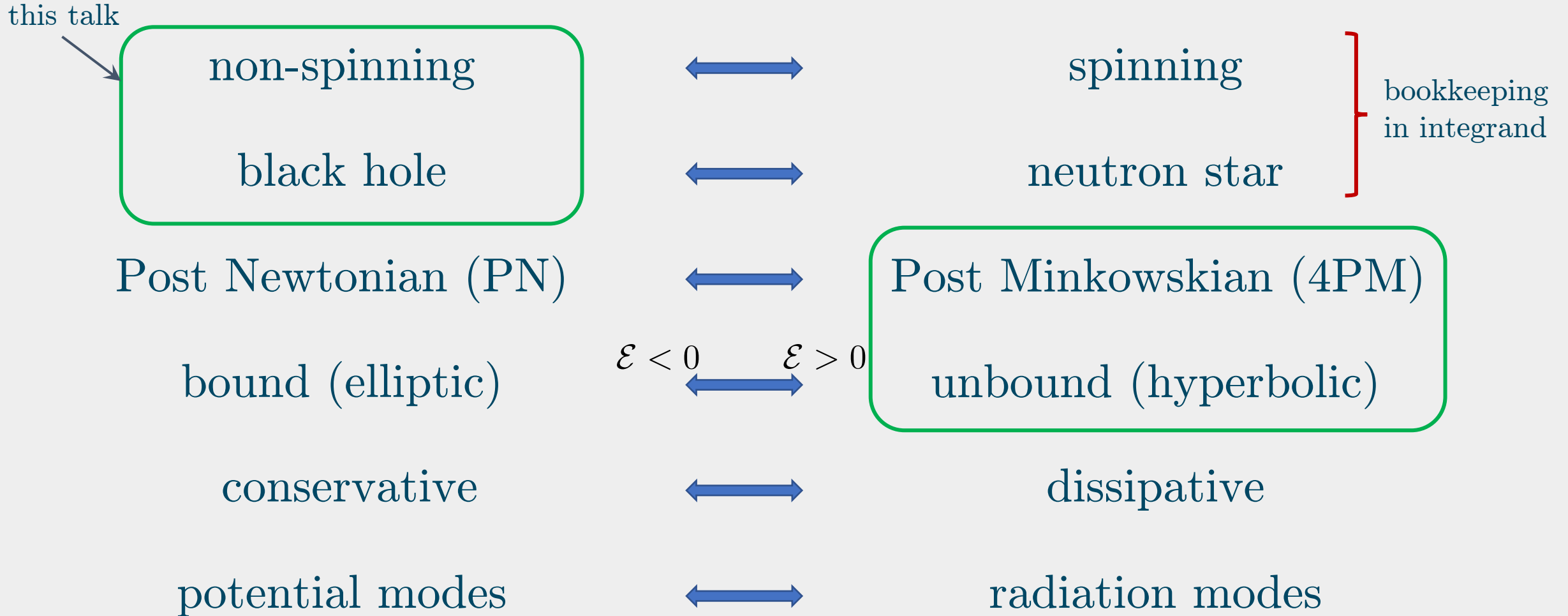
dissipative

potential modes



radiation modes

Terminology



Conservative vs. dissipative

- Conservative: standard Feynman propagators

$$\Delta_{\text{F}}(\ell) = \frac{1}{\ell^2 + i0} \quad \text{related to in-out formalism} \Rightarrow \text{vacuum-to-vacuum amplitudes}$$

- Dissipative: retarded/advanced propagators

$$\Delta_{\text{ret/adv}}(\ell) = \frac{1}{\ell^2 \pm i0} \quad \text{related to in-in formalism} \Rightarrow \text{vacuum only for in-state}$$

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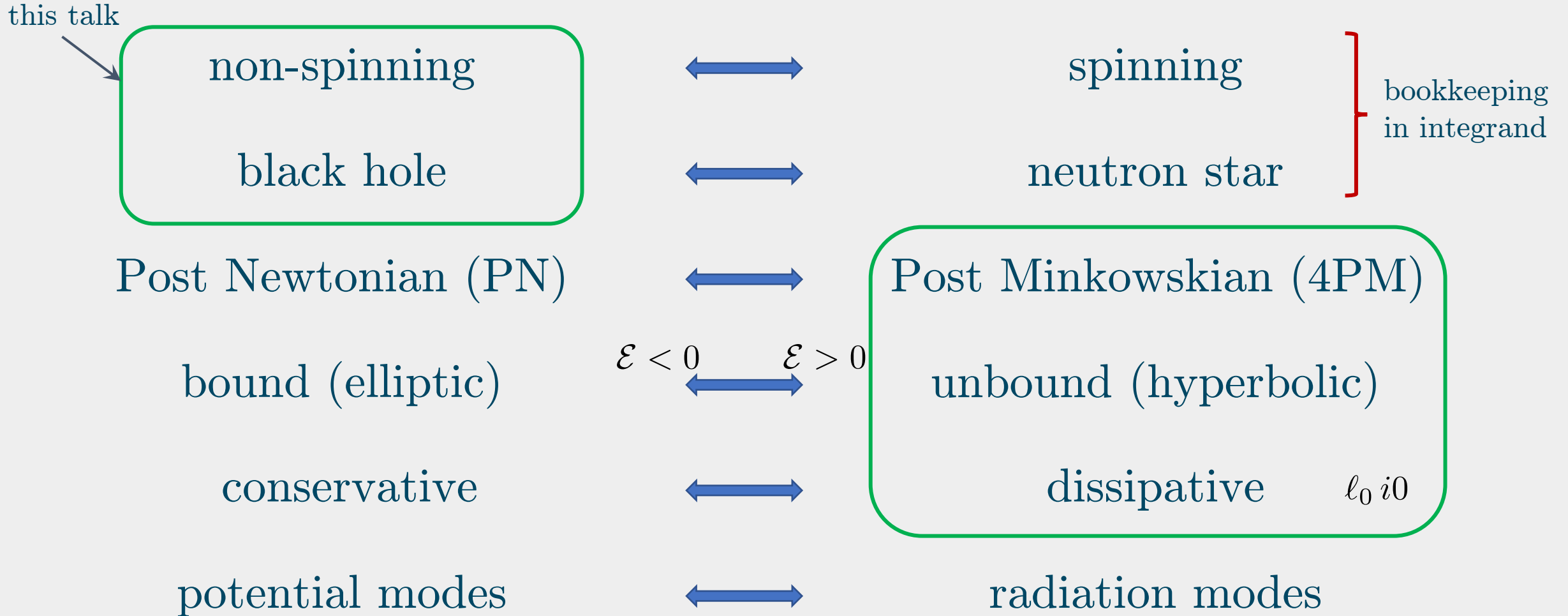
- Dissipative: retarded/advanced propagators

$$\Delta_{\text{ret/adv}}(\ell) = \frac{1}{\ell^2 \pm \ell_0 i0} \quad \text{related to in-in formalism} \Rightarrow \text{vacuum only for in-state}$$

- alternatively: view as virtual + real

$$\Delta_{\text{ret/adv}}(\ell) = \Delta_{\text{F}}(\ell) + 2\pi i \Theta(\mp \ell_0) \delta(\ell^2)$$

Terminology

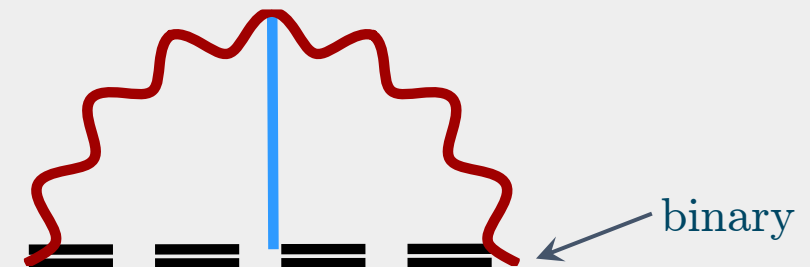


Potential vs. radiation modes

- Potential modes:
 - off-shell integration region
 - $i0$ prescription irrelevant
- Radiation modes:
 - on-shell integration region
 - $i0$ prescription important

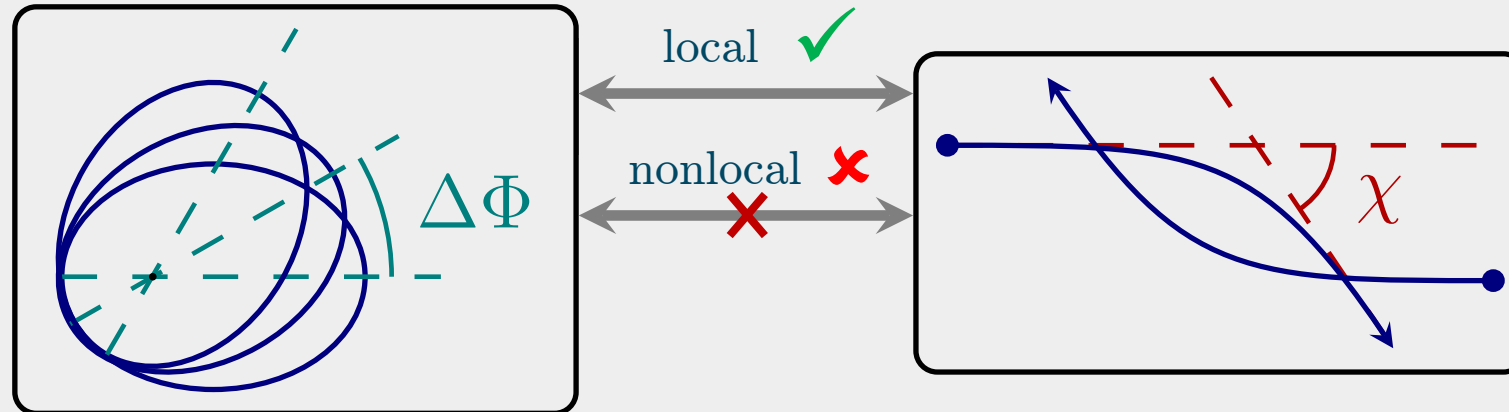
Potential vs. radiation modes

- Potential modes:
 - off-shell integration region
 - $i0$ prescription irrelevant
- Radiation modes:
 - on-shell integration region
 - $i0$ prescription important
 - Radiation modes possible in conservative computation!
 - complicates continuation to elliptic

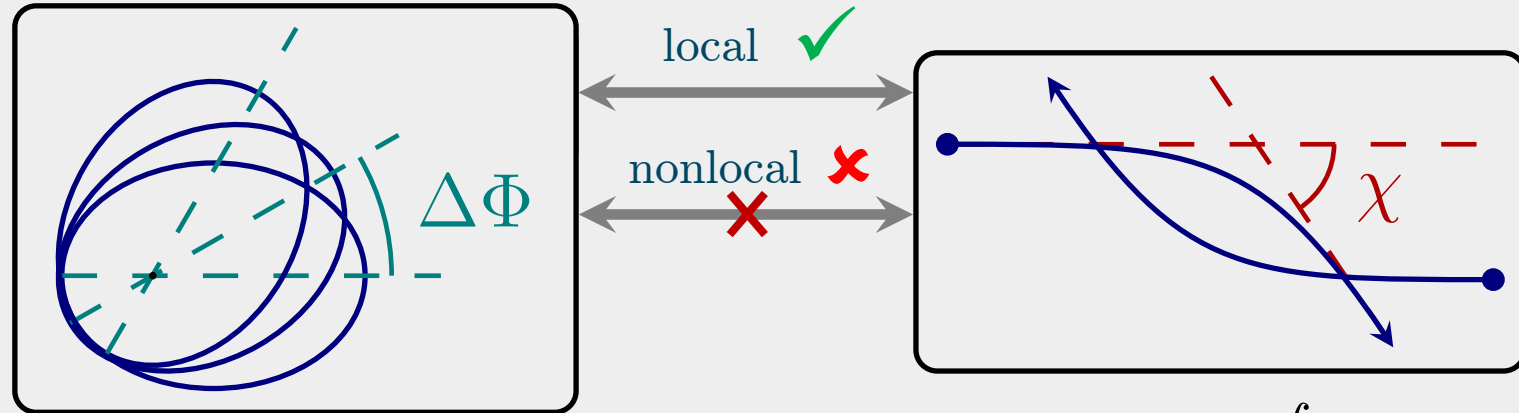


tail effect
[Galley, Leibovich, Porto, Ross, '15]

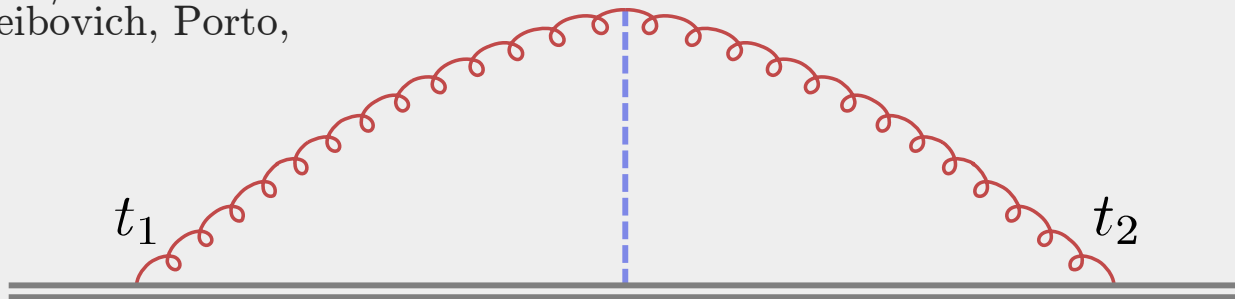
Nonlocal-in-time contributions



Nonlocal-in-time contributions



[Cho, Kälin, Porto, '22 /
Damour, Jaranowski,
Schäfer, '14 /
Galley, Leibovich, Porto,
Ross, '15]

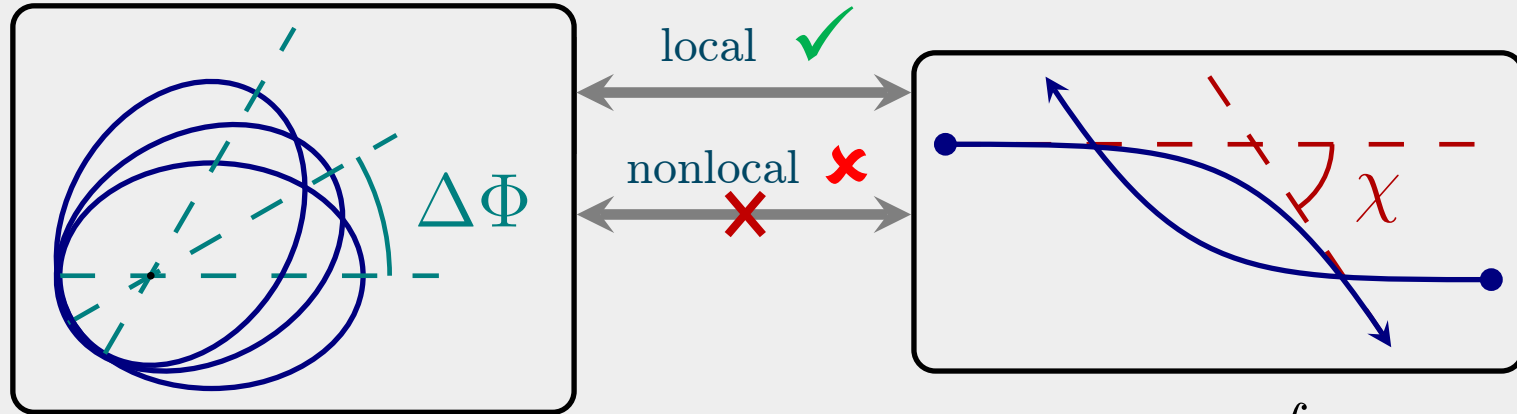


$$\mathcal{S}_r^{\text{nlloc}} \sim GE \int \mathcal{F}(t_1, t_2) \frac{dt_1 dt_2}{|t_1 - t_2|},$$

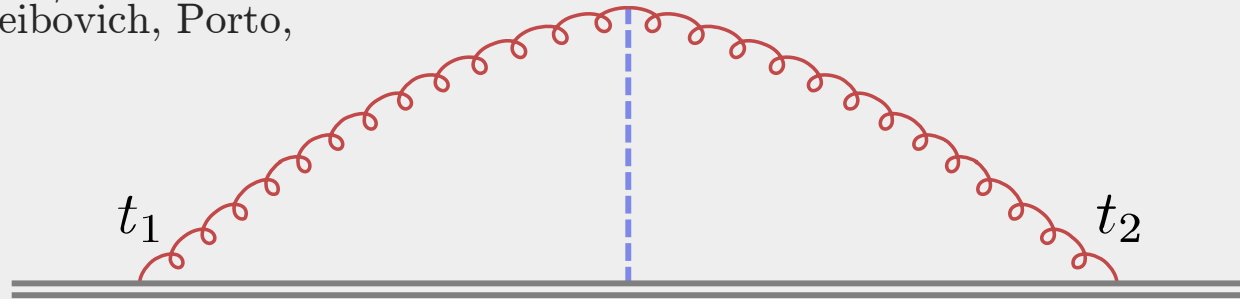
$$\sim GE \int \frac{dE}{d\omega} \log(\omega^2) d\omega$$

$$\chi \sim \partial_j \mathcal{S}_r$$

Nonlocal-in-time contributions



[Cho, Kälin, Porto, '22 /
Damour, Jaranowski,
Schäfer, '14 /
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$$\mathcal{S}_r^{\text{nlloc}} \sim GE \int \mathcal{F}(t_1, t_2) \frac{dt_1 dt_2}{|t_1 - t_2|},$$

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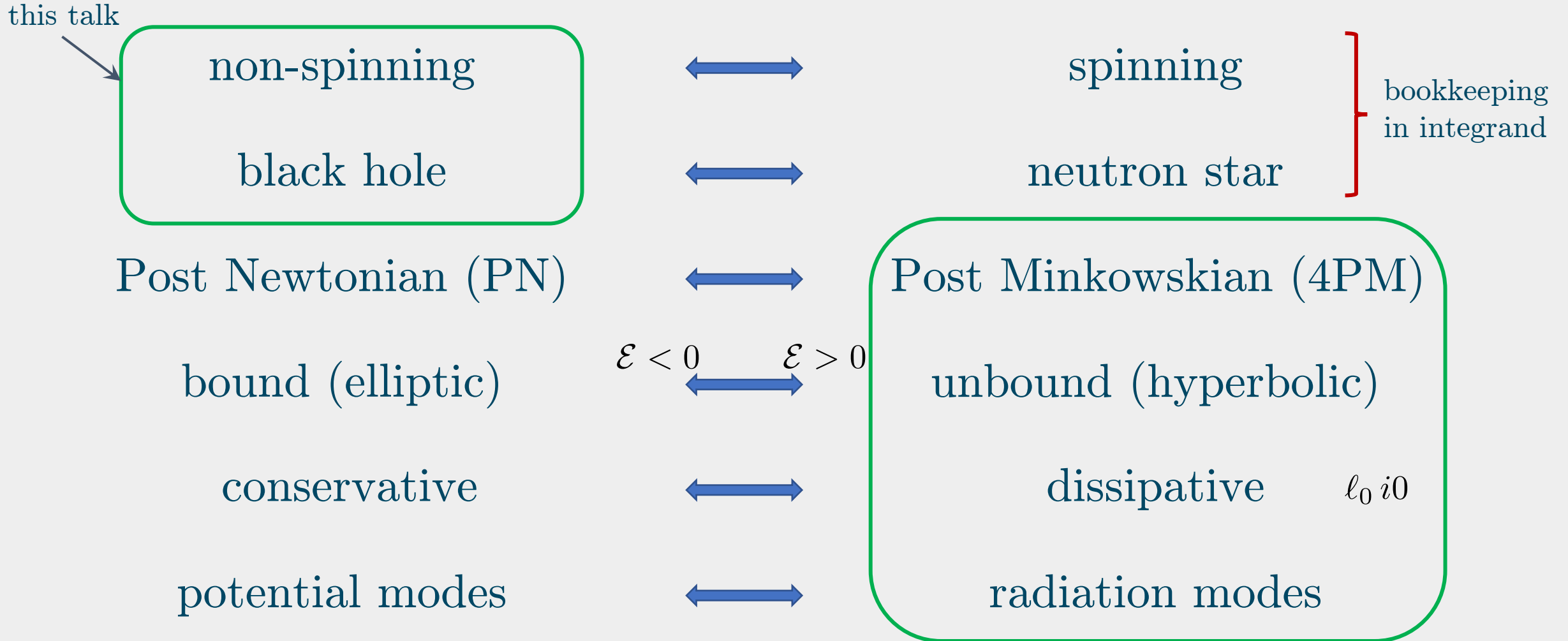
$$\Delta p_1 \rightarrow \Delta E$$

$$\Delta E = \int \frac{dE}{d\omega} d\omega,$$

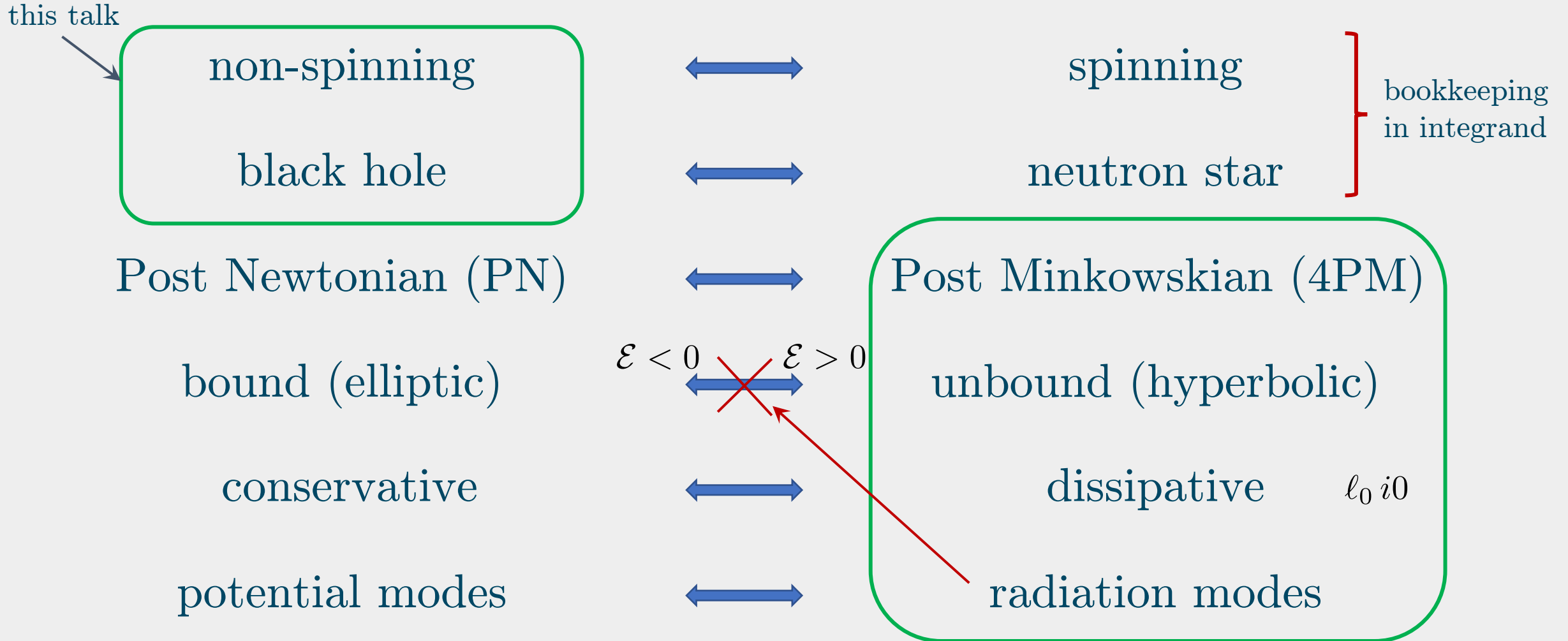
⇒ add log to integrand

- Related to previous computation:

Terminology



Terminology



Analytic computation methods

- Setup
 - Effective theory of point particles
- Integral computation
 - Example diagram
 - Method of differential equations (+ elliptic)
 - Boundary conditions

Other approaches:

- Scattering amplitudes [Cheung, Rothstein, Solon; Bern, Parra-Martinez, Roiban, Ruf, Shen, Zeng, ...]
- WL-QFT [Jakobsen, Mogull, Plefka, Steinhoff, '21]
- Eikonal [Di Vecchia, Heissenberg, Russo, Veneziano]

Gravitational two-body problem

$$S_{\text{pp}} = - \sum_{a=1,2} \frac{m_a}{2} \int d\tau g_{\mu\nu} v_a^\mu v_a^\nu = - \sum_{a=1,2} \frac{m_a}{2} \int d\tau [\eta_{\mu\nu} v_a^\mu v_a^\nu + h_{\mu\nu} v_a^\mu v_a^\nu]$$

$v_a^\mu = \frac{dx_a^\mu}{d\tau}$, weak field: $g_{\mu\nu} = \eta_{\mu\nu} + M_{\text{Pl}}^{-2} h_{\mu\nu}$

$$S_{\text{EH}} = -2M_{\text{Pl}}^2 \int d^4x \sqrt{-g} R,$$

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- EFT approach: $e^{iS_{\text{eff}}[x_a]} = \int \mathcal{D}h_{\mu\nu} e^{iS_{\text{EH}}[h] + iS_{\text{GF}}[h] + iS_{\text{pp}}[x_a, h]}$
- Post-Minkowskian expansion: $S_{\text{eff}} = \sum_{n=0}^{\infty} G^n S_n,$

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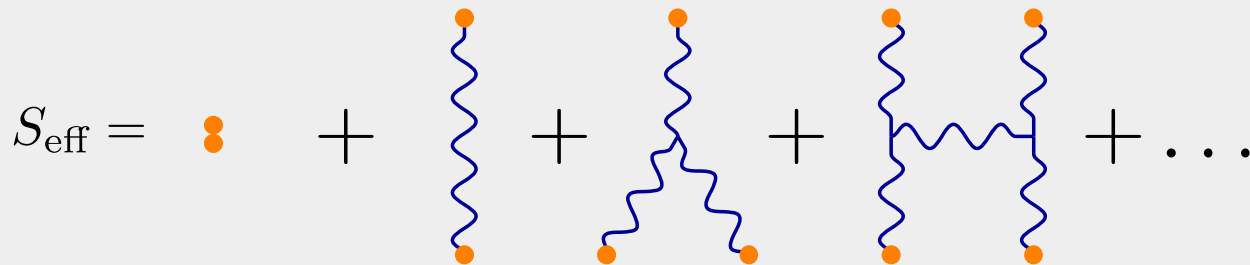
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$$S_{\text{eff}} = \text{orange dots} + \text{wavy line} + \text{wavy line with vertex} + \text{wavy line with loop} + \dots = - \sum_{a=1,2} \frac{m_a}{2} \int d\tau \eta_{\mu\nu} v_a^\mu v_a^\nu$$

Gravitational two-body problem

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$$S_{\text{eff}} = \text{orange dots} + \text{wavy line} + \text{two wavy lines} + \text{three wavy lines} + \dots = - \sum_{a=1,2} \frac{m_a}{2} \int d\tau \eta_{\mu\nu} v_a^\mu v_a^\nu + G \frac{i}{4} \int d\tau_1 d\tau_2 v_1^\alpha v_1^\beta v_2^\mu v_2^\nu \int d^4k \frac{P_{\alpha\beta\mu\nu}}{\ell^2} e^{i\ell \cdot (x_1 - x_2)}$$

Change in momentum

$$S_{\text{eff}} = \sum_{n=0}^{\infty} G^n S_n = - \sum_{a=1,2} \frac{m_a}{2} \int d\tau \eta_{\mu\nu} v_a^\mu v_a^\nu + \dots,$$

$$S_n = \int d\tau_1 \mathcal{L}_n[x_1, x_2]$$

• Equations of motion:

$$0 = \sum_{n=0}^{\infty} \left(\frac{\partial \mathcal{L}_n}{\partial x_1^\nu} - \frac{d}{d\tau_1} \frac{\partial \mathcal{L}_n}{\partial v_1^\nu} \right)$$
$$m_1 \frac{dv_1^\mu}{d\tau_1} = -\eta^{\mu\nu} \sum_{n=1}^{\infty} \left(\frac{\partial \mathcal{L}_n}{\partial x_1^\nu} - \frac{d}{d\tau_1} \frac{\partial \mathcal{L}_n}{\partial v_1^\nu} \right)$$

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• Corrections to trajectories:

$$\int_{-\infty}^{\tau_1} d\tilde{\tau}_1 \longrightarrow v_a^\mu = u_a^\mu + \sum_n G^n \delta^{(n)} v_a^\mu$$

• Total momentum change:

$$\int_{-\infty}^{\infty} d\tau_1 \longrightarrow \Delta p_1^\mu = m_1 \Delta v_1^\mu$$

Change in momentum

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• Total momentum change: $\int_{-\infty}^{\infty} d\tau_1 \longrightarrow \Delta p_1^\mu = m_1 \Delta v_1^\mu$ cut propagator

• PP propagators: $\int_{-\infty}^{\tau_1} d\tilde{\tau}_1 e^{i\tilde{\tau}_1(\ell \cdot u_1)} = \frac{-i e^{i\tau_1(\ell \cdot u_1)}}{\ell \cdot u_1}, \quad \int_{-\infty}^{\infty} d\tau_1 e^{i\tau_1(\ell \cdot u_1)} = 2\pi \delta(\ell \cdot u_1)$

Loop integrals at 4PM

- Example integral family:

$$D = 4 - 2\epsilon$$

$$I(\gamma) = \int d^D l_1 d^D l_2 \frac{\delta(l_1 \cdot u_1) \delta(l_2 \cdot u_2)}{(l_1 \cdot u_2)(l_2 \cdot u_1)} \frac{1}{l_1^2 l_2^2 (l_1 + l_2 - q)^2}$$

$$u_1^2 = u_2^2 = 1,$$

$$u_1 \cdot q = u_2 \cdot q = 0$$

$$u_1 \cdot u_2 = \gamma$$

- Treat cut propagators just as linear propagators

Loop integrals at 4PM

- Example integral family:

$$D = 4 - 2\epsilon$$

The diagram shows a two-loop integral with two wavy lines and two cut propagators. The top loop is formed by a solid line and a dashed line. The bottom loop is formed by a solid line and a dashed line. The two wavy lines connect the top and bottom lines. The cut propagators are the dashed lines. The diagram is labeled with $I(\gamma) =$ and the integral expression $= \int d^D l_1 d^D l_2 \frac{\delta(l_1 \cdot u_1) \delta(l_2 \cdot u_2)}{(l_1 \cdot u_2)(l_2 \cdot u_1)} \frac{1}{l_1^2 l_2^2 (l_1 + l_2 - q)^2}$. Arrows point from the labels $\frac{1}{l_1^2}$, $\frac{1}{l_1 \cdot u_2}$, and $\delta(l_2 \cdot u_2)$ to the corresponding parts of the diagram.

$$u_1^2 = u_2^2 = 1,$$

$$u_1 \cdot q = u_2 \cdot q = 0$$

$$u_1 \cdot u_2 = \gamma$$

- Treat cut propagators just as linear propagators
- Rationalize square-root:

$$\gamma = \frac{1}{2}(1/x + x),$$

$$v \equiv \sqrt{\gamma^2 - 1} = \frac{1}{2}(1/x - x)$$

only one variable

Differential equations method

- Integration-by-parts reduction to basis [FIRE6, Smirnov, Chukharev, '19 / LiteRed, Lee, '13]

$$I_{a_1 \dots a_{15}} = \sum_i c_i(x, \epsilon) f_i$$

ret/adv \Rightarrow turn off symmetry detection

Differential equations method

- Integration-by-parts reduction to basis [FIRE6, Smirnov, Chukharev, '19 / LiteRed, Lee, '13]

$$I_{a_1 \dots a_{15}} = \sum_i c_i(x, \epsilon) f_i \quad \text{ret/adv} \Rightarrow \text{turn off symmetry detection}$$

- Solve by transforming to canonical form: [Henn, '13]

$$\partial_x \vec{f} = A(x, \epsilon) \vec{f} \quad \vec{g} = T \vec{f} \quad \partial_x \vec{g} = \epsilon \tilde{A}(x) \vec{g} \quad D = 4 - 2\epsilon$$

rational \nearrow [Lee, '15]
 \longrightarrow [Libra, epsilon, Fuchsia, CANONICA]

[INITIAL:
 CD, Henn, Wagner, '22;
 CD, Henn, Yan, '20]

Dyson series

$$\vec{g} = \text{P} e^{\epsilon \int_{x_0}^x \tilde{A}(\tilde{x}) d\tilde{x}} \vec{g}_0(\epsilon) = \left[1 + \epsilon \int_{x_0}^x dx_1 \tilde{A}(x_1) + \epsilon^2 \int_{x_0}^x dx_1 \tilde{A}(x_1) \int_{x_0}^{x_1} dx_2 \tilde{A}(x_2) + \mathcal{O}(\epsilon^3) \right] \vec{g}_0(\epsilon)$$

Elliptic integrals

- Polylogarithms for most sectors

$$\partial_x \vec{g} = \epsilon \tilde{A}(x) \vec{g}, \quad \tilde{A}(x) = \sum_i M_i a_i(x)$$

\uparrow
 constant matrices

$$a_i(x) \in \left\{ \frac{1}{x}, \frac{1}{x-1}, \frac{1}{x+1}, \frac{x}{1+x^2} \right\}$$

new at 4PM

- 4PM: Elliptics integrals: $K(z) = \int_0^1 \frac{dt}{\sqrt{(1-t^2)(1-zt^2)}}$,

- 5PM: Calabi-Yau periods [Driesse et al, Nature 264]

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input from
mathematics,
computer
science and
particle physics

output to
particle physics

Boundary conditions

- From PN-like limit

$$v^2 \ll 1$$

- method of regions:

[Beneke, Smirnov, '97]

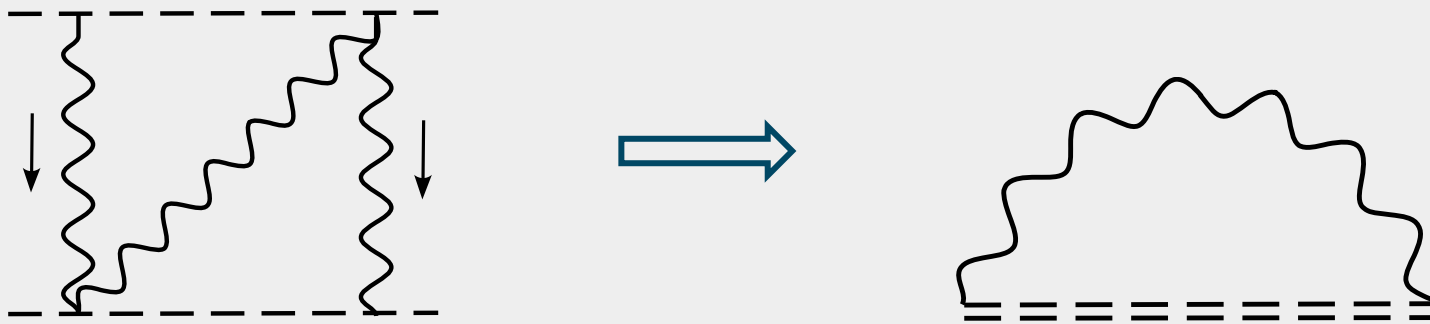
[asy2.m, Jantzen, Smirnov, Smirnov, '12; Pak, Smirnov, '10]

potential graviton: $(\ell^0, \ell) \sim (1, v)$

radiation graviton: $(\ell^0, \ell) \sim (v, v)$

[Goldberger, Rothstein, '06]

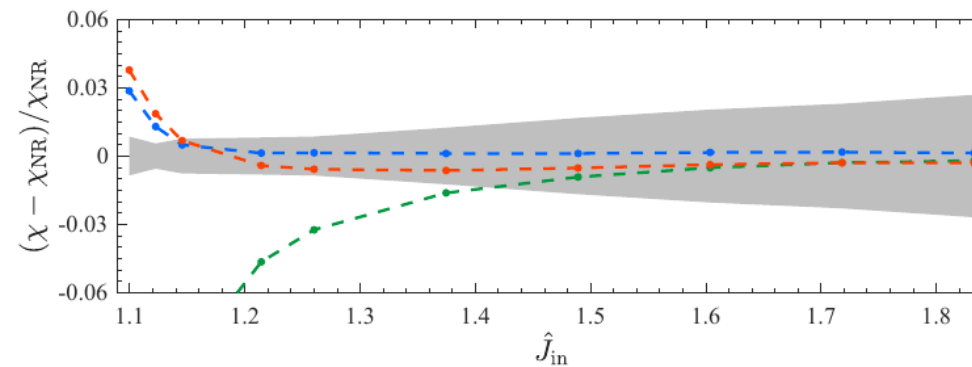
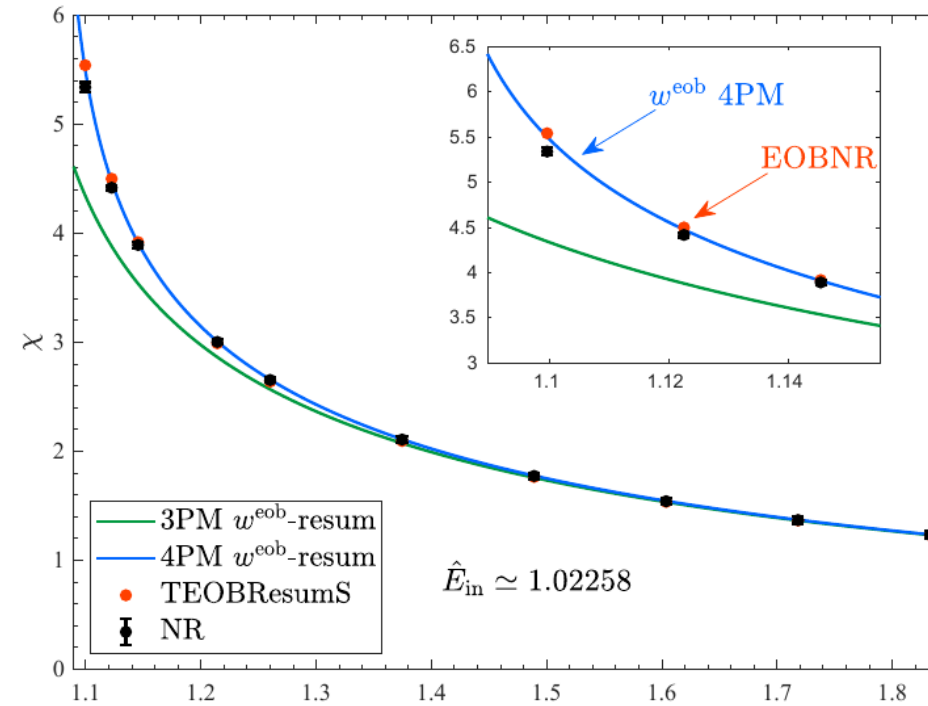
- correspond nicely to PN regions and (3d) integrals



[Galley, Leibovich, Porto, Ross, '15]

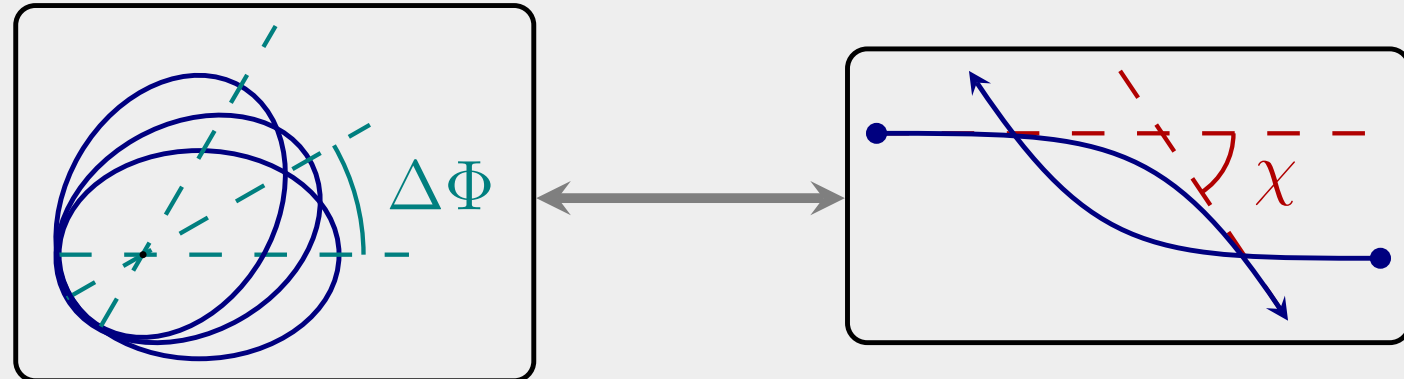
PM radiative state-of-the-art accuracy

[Damour, Rettegno, '22]



Summary

- Theory is not ready for 3G GW observatories
- Gravitational two-body problem
 - Post-Minkowskian approximation
 - Use **tools from particle physics**
 - Result for scattering shows remarkable agreement with NR
- B2B map:
 - Bound from scattering
 - Nonlocality problematic
- Outlook: **Waveform**



Challenges

- Detector accuracy
 - Higher PM, spin, internal structure, ...
 - Adjust QFT tools more to our needs
 - Continue scattering to bound state data
 - Incorporate PM more directly into waveform generation (EOB)

[Damour, '16; Damgaard, Vanhove, '21; Khalil, Buonanno, Steinhoff, Vines, '22; Damour, Rettegno, '23; ...]

Boundary conditions

$$\vec{f} = T^{-1} \mathbf{P} e^{\epsilon \int \tilde{A}(x) dx} \vec{g}_0(\epsilon)$$

- Compare series expansions around singular point

- Small velocity expansion: $v \equiv \sqrt{\gamma^2 - 1}$

$$v \rightarrow 0 \iff \gamma \rightarrow 1 \iff x \rightarrow 1$$

- 1) Solution of differential equations:

[Lee, Smirnov, Smirnov, Steinhauser, '19]
[Libra: Lee, '20]

- 2) Explicit integral expansions:

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$$\partial_x \vec{f} = A(x, \epsilon) \vec{f}, \quad \vec{f}(v, \epsilon) \simeq \sum_{n_1, n_2, k} v^{n_1 + n_2 \epsilon} \log^k(v) H_{n_1, n_2, k}(\epsilon) \vec{g}_0(\epsilon), \quad n_1, n_2, k \in \mathbb{Z}$$

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2) Explicit integral expansions:

method of regions:

$$I = \int dk \frac{\delta(k \cdot u_1) \dots}{k^2 (k \cdot u_2) \dots}, \quad \vec{f}(v, \epsilon) \simeq \sum_{n_1, n_2, k} v^{n_1 + n_2 \epsilon} \log^k(v) \vec{h}_{n_1, n_2, k}(\epsilon) \longleftarrow \text{PN-like integrals}$$

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relations between infinite set of PN-like integrals

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relations between infinite set of PN-like integrals

- Identify an independent set:

$$\vec{g}_0(\epsilon) = H_{\text{indep}}(\epsilon)^{-1} \vec{h}_{\text{indep}}(\epsilon)$$

The Scattering Amplitudes Approach

$$S_{\text{GR}} = \int d^D x \sqrt{-g} \left[-\frac{1}{16\pi G} R + \frac{1}{2} \sum_{a=1,2} (D^\mu \phi_a D_\mu \phi_a - m_a^2 \phi_a^2) \right]$$

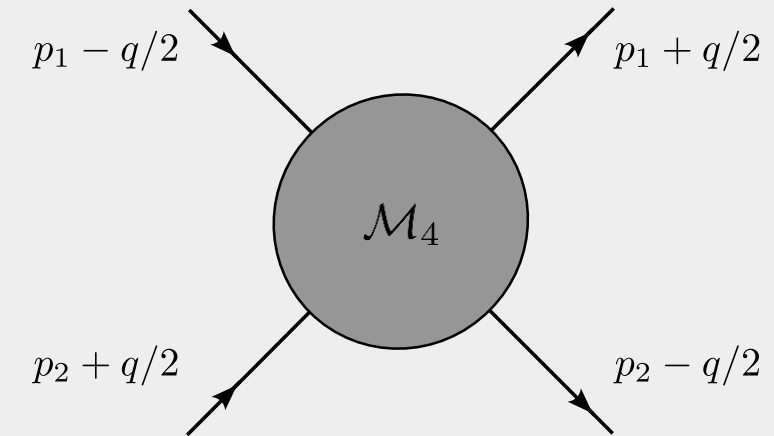
- Integrand construction:

- Generalized unitarity $C = \sum M_{(1)}^{\text{tree}} M_{(2)}^{\text{tree}} \dots M_{(m)}^{\text{tree}}$

- Double copy

$$M_{\text{tree}} \sim (A_{\text{tree}})^2$$

Gravity \nearrow \nwarrow Yang-Mills



$$p_1 \cdot q = p_2 \cdot q = 0$$

$$p_1^\mu = m_1 u_1^\mu, \quad p_2^\mu = m_2 u_2^\mu$$

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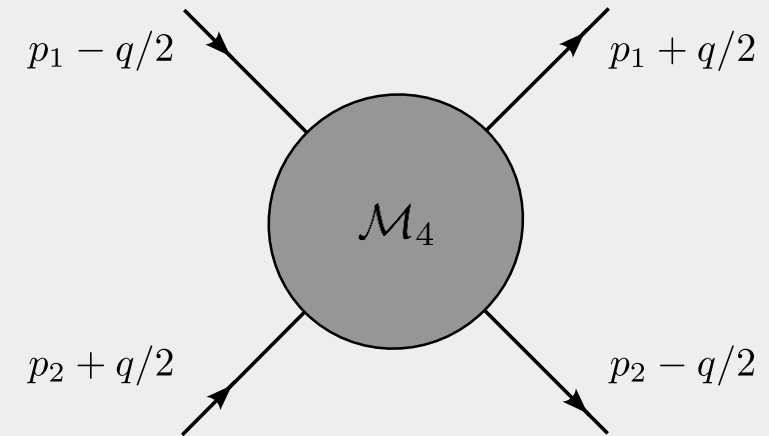
Gravity
Yang-Mills

- Classical limit:

$$G \longleftrightarrow \text{number of loops}$$

$$\hbar \longleftrightarrow |q| \quad \text{leading term in soft expansion!}$$

$$\frac{1}{(\ell + p_i - q/2)^2 - m_1^2} \simeq \frac{1}{2p_1 \cdot \ell} = \frac{1}{m_1 2u_1 \cdot \ell}$$



$$p_1 \cdot q = p_2 \cdot q = 0$$

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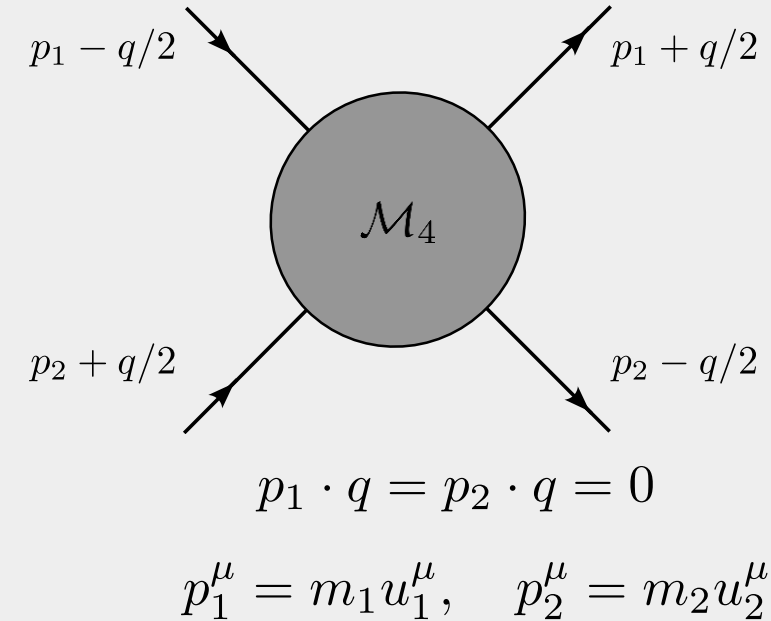
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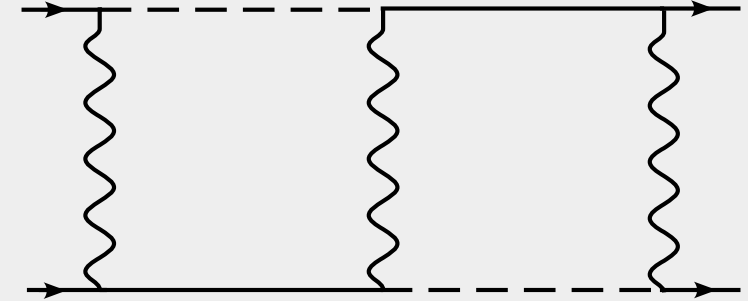
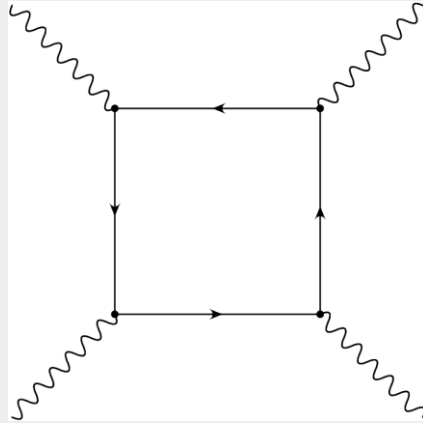
$$\frac{1}{(\ell + p_i - q/2)^2 - m_1^2} \simeq \frac{1}{2p_1 \cdot \ell} = \frac{1}{m_1 2u_1 \cdot \ell}$$

- Potential from EFT matching



The classical limit

[Damgaard, Haddad, Helset, '21,
Holstein, Donoghue, '13]



$$\int \mathcal{D}\phi e^{\frac{i}{\hbar}(S_0 + gS_3 + g^2S_4)} \sim \int \mathcal{D}\phi e^{\frac{i}{\hbar g^2}S}$$

$$\alpha = \frac{e^2}{4\pi\epsilon_0\hbar c}$$

$$p = \hbar k$$

$$m = \frac{\hbar}{\lambda}$$

$$b \gg \lambda$$

$$q \ll m$$

Elliptic integrals

- Polylogarithms for most sectors

$$a_i(x) \in \left\{ \frac{1}{x}, \frac{1}{x-1}, \frac{1}{x+1}, \frac{x}{1+x^2} \right\} \leftarrow \text{new at 4PM}$$

$$\partial_x \vec{g} = \epsilon \tilde{A}(x) \vec{g}, \quad \tilde{A}(x) = \sum_i M_i a_i(x)$$

↑
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constant matrices

- Algorithms do not work for one sector $\partial_x \vec{f} = A(x, \epsilon) \vec{f}$
 - analyze ϵ^0 -part:

$$A(x, \epsilon) = A^{(0)}(x) + \epsilon A^{(1)}(x) + \mathcal{O}(\epsilon^2), \quad \partial_x T^{(0)}(x) = A^{(0)}(x) T^{(0)}(x)$$

↑
remove with transformation

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\uparrow
 constant matrices

- Algorithms do not work for one sector

➤ analyze ϵ^0 -part:

$$\begin{aligned} \partial_x \vec{f} &= A(x, \epsilon) \vec{f} \\ \partial_\gamma \vec{f} &= A(\gamma, \epsilon) \vec{f} \longrightarrow \sqrt{\gamma^2 - 1} \\ &T^{(0)}(\gamma) \end{aligned}$$

$$A(x, \epsilon) = A^{(0)}(x) + \epsilon A^{(1)}(x) + \mathcal{O}(\epsilon^2),$$

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\uparrow
remove with transformation

Elliptic integrals

- Polylogarithms for most sectors $\partial_x \vec{g} = \epsilon \tilde{A}(x) \vec{g}$, $\tilde{A}(x) = \sum_i M_i a_i(x)$
constant matrices
- Algorithms do not work for one sector

- Third-order DE: $\left[\partial_x^3 - \frac{6x}{1-x^2} \partial_x^2 - \frac{1-4x^2+7x^4}{x^2(1-x^2)^2} \partial_x - \frac{1+x^2}{x^3(1-x^2)} \right] \Psi_{1,2,3} = 0$

- Solve with Mathematica: $\Psi_1 = xK^2(1-x^2)$, $\Psi_2 = xK(1-x^2)K(x^2)$, $\Psi_3 = xK^2(x^2)$

$$K(z) = \int_0^1 \frac{dt}{\sqrt{(1-t^2)(1-zt^2)}},$$

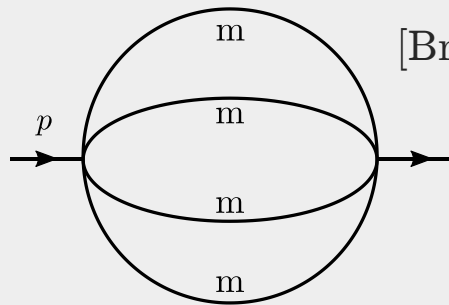
$$E(z) = \int_0^1 \frac{1-zt^2}{\sqrt{1-t^2}} dt$$

Elliptic integrals

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[Broedel, Duhr, Dulat, Marzucca, Penante, Tancredi, '19;
 Broedel, Duhr, Matthes, '21;
 Pögel, Wang, Weinzierl, '22]

$$K(z) = \int_0^1 \frac{dt}{\sqrt{(1-t^2)(1-zt^2)}}$$


$$E(z) = \int_0^1 \frac{1-zt^2}{\sqrt{1-t^2}} dt$$

ε-form using INITIAL

- Elliptic differential equations

68 master integrals:
 ⇒ constant 68 × 68 matrices

$$\partial_x \vec{g} = \epsilon \tilde{A}(x) \vec{g},$$

$$\tilde{A}(x) = \sum_i M_i a_i(x)$$


$$a_i(x) \in \left\{ \begin{array}{ccccc} \frac{\pi^2}{x(1-x^2)K^2(1-x^2)}, & \frac{1}{1-x}, & \frac{1}{x}, & \frac{1}{1+x}, & \frac{x}{1+x^2}, \\ \frac{K^2(1-x^2)}{\pi^2 x(1-x^2)}, & \frac{K^2(1-x^2)}{\pi^2(1-x^2)}, & \frac{K^2(1-x^2)}{\pi^2 x}, & \frac{K^2(1-x^2)}{\pi^2}, & \frac{(1-x^2)K^2(1-x^2)}{\pi^2 x}, \\ \frac{K^4(1-x^2)}{\pi^4 x(1-x^2)}, & \frac{K^4(1-x^2)}{\pi^4 x}, & \frac{(1-x^2)K^4(1-x^2)}{\pi^4 x}, & \frac{(1-x^2)^2 K^4(1-x^2)}{\pi^4 x} & \end{array} \right\}$$

- Differential equations not problematic for us, despite elliptics
- Canonical form just as in polylogarithmic 3PM case