Power corrections in collider processes

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ACCADEMIA NAZIONALE DEI LINCEI

- ▶ Power corrections in collider processes
- ▶ Renormalons and linear power corrections
- ▶ Massless partons
- \blacktriangleright Massive partons
- \triangleright e^+e^- annihilation: shape-variables in the 3-jet region.
- ▶ Fits to e^+e^- data.

Power corrections for collider processes

- ▶ Little is known about power corrections in QCD processes.
- ▶ Some simpler processes admit an OPE (the total cross section in e^+e^- annihilation and similar processes, DIS-like processes, B meson decays \ldots) so that power corrections can be parametrized.
- \triangleright For the complex collider processes one worries about the presence of linear power corrections, i.e. corrections of the order of Λ/Q , since these could be at the percent level, that is the accuracy one is aiming for at the HL LHC.
- ▶ One instrument for the investigation of linear power correction is the study of renormalons in the large b_0 approximation.

All-orders contributions to QCD amplitudes of the form

$$
\int_0^m \mathrm{d}k^p \,\alpha_s(k^2) = \int_0^m \mathrm{d}k^p \, \frac{1}{b_0 \log(k^2/\Lambda^2)}
$$

=
$$
\int_0^m \mathrm{d}k^p \frac{\alpha_s(m^2)}{1 + b_0 \alpha_s(m^2) \log \frac{k^2}{m^2}}
$$

=
$$
\alpha_s(m^2) \sum_{n=0}^\infty (2b_0 \alpha_s(m^2))^n \int_0^m \mathrm{d}k^p \log^n \frac{m}{k}
$$

.

Asymptotic expansion.

- ▶ Minimal term at $n_{\min} \approx \frac{1}{2pb_0\alpha_S(m^2)}$.
- Size of minimal term: $m^p \alpha_s(m^2) \sqrt{2\pi n_{\text{min}}} e^{-n_{\text{min}}} \approx \Lambda^p$.

Given an (IR safe) observable O, we introduce the notation

- \blacktriangleright Φ_B , phase space;
- $\blacktriangleright \Phi_{\epsilon}$, phase space for the emission of one massive gluon with mass λ .
- $\triangleright \Phi_{q\bar{q}}$, phase space for the emission of a $q\bar{q}$ pair
- the all-order result can be expressed in terms of
	- \triangleright $\sigma_B(\Phi_B)$, the differential cross section for the Born process;
	- $\blacktriangleright \sigma_{v}(\lambda, \Phi_{B})$, the virtual correction to the Born process due to the exchange of a gluon of mass λ ;
	- ► The real cross section $\sigma_{g^*}(\lambda, \Phi_{g^*})$, obtained by adding one massive gluon to the Born final state;
	- **►** The real cross section $\sigma_{q\bar{q}}(\Phi_{q\bar{q}})$, obtained by adding a $q\bar{q}$ pair, produced by a massless gluon, to the Born final state;

Large- n_f all-order result

Defining:

$$
T_O(\lambda) = \n\begin{array}{rcl}\n\text{result for a gluon with mass } \lambda & \text{Seymour, PN1995} \\
T_O(\lambda) & = & \sqrt{V_O(\lambda) + R_O(\lambda)} + \Delta_O(\lambda) \\
V_O(\lambda) & = & \int d\Phi_b \sigma_v^{(1)}(\lambda^2, \Phi_b) O(\Phi_b), \\
R_O(\lambda) & = & \int d\Phi_{g^*} \sigma_{g^*}^{(1)}(\lambda^2, \Phi_{g^*}) O(\Phi_{g^*}), \\
\Delta_O(\lambda) & = & \frac{3\pi\lambda^2}{\alpha_S T_F} \int d\Phi_{q\bar{q}} R_{q\bar{q}}(\Phi_{q\bar{q}}) \delta(m_{q\bar{q}}^2 - \lambda^2) \left[O(\Phi_{q\bar{q}}) - O(\Phi_{g^*}) \right]\n\end{array}
$$

The Δ term vanishes if the observable is totally inclusive in the radiated partons.

It turns out that a linear term in λ in the expansion of $T(\lambda)$ around zero is associated with linear renormalons.

Large- n_f all-order result

The all-order result is given by

$$
\langle O \rangle = B_O - \int d \lambda \frac{\mathrm{d} T_O(\lambda)}{\mathrm{d} \lambda} \underbrace{\frac{1}{\alpha_S} \overbrace{\left[\frac{1}{\pi b_0} \arctan \frac{\pi b_0 \alpha_S}{1 + b_0 \alpha_S \log \lambda^2 / \mu_C^2} \right]}}_{\alpha_{\mathrm{s,eff}}(\lambda)/\alpha_S}
$$

It is easy to show that a linear λ term in $T_O(\lambda)$ leads to a factorial growth related to a linear IR renormalon. In fact

$$
\int d\lambda \left[\frac{1}{\pi} \arctan \frac{\pi b_0 \alpha_S}{1 + b_0 \alpha_S \log \lambda^2 / \mu_C^2} \right] =
$$
\n
$$
\frac{1}{\pi} P \int_0^\infty dt \frac{\exp \left(-\frac{t}{2b_0 \alpha_S} \right)}{1 - t} - \exp \left(-\frac{1}{2b_0 \alpha_S} \right)
$$
\n+ terms analytic in α_S . (1)

Large- n_f all-order result

- ▶ We have a well-defined procedure for the computation of the T function..
- \triangleright Can be computed semi-numerically. This approach has been followed in
	- ▶ Ferrario Ravasio, Oleari, PN, 2019 for studies related to the top mass measurements.
	- ▶ Ferrario Ravasio, Limatola, PN, 2021 for showing the absence of linear corrections to the p_T spectrum of the Z in hadronic collisions.

Gavin Salam had often shown an argument in favour of the presence of linear power corrections to the inclusive p_T spectrum of the Z boson, based upon the fact that the soft radiation associated to this process is not azimuthally symmetric. Our attemt to actually compute such an effect in a model theory gave negative results.

It is however difficult, numerically, to show the absence of a correction, especially in this case where the cancellation of soft-collinear divergence between the virtual (computed analytically) and real (computed numerically) is involved. Analytic results were found:

- ▶ Analytic approach for massless partons: Caola,Ferrario Ravasio,Limatola,Melnikov, PN 2021,[2108.08897], same authors + Ozcelik 2022[2204.02247]
- ▶ Analytic approach for massive partons: Makarov, Melnikov, Ozcelik, PN, 2023, [2302.02729], 2023[2308.05526], 2024[2408.00632]

Cancellation of linear NP terms

Our findings can be summarized as follows:

- \triangleright Consider a process, described by a cross section (with no radiation) $B(p_1, \ldots, p_n)$ where p denotes a set of fixed external momenta, with $p_3 \ldots p_n$ colourless particles, and p_1, p_2 massless quark antiquark (final-final or initial-initial) or quark-quark (initial-final) pair.
- Assume that we emit a massive gluon of mass λ and momentum k , and we have a smooth (in a sense to be clarified afterwards), IR safe mapping from the full real emission configuration to the underlying Born one.

Then:

- \blacktriangleright No linear λ sensitivity arises from virtual corrections
- \blacktriangleright No linear λ sensitivity arises from the real contributions due to an unrestricted integration in k at fixed underlying Born kinematics.

The result is based upon two observations:

- ▶ Virtual corrections have no linear power corrections. One can show that the virtual integrals give rise to constants, logs and double logs of λ , but no linear terms in λ .
- \triangleright Writing the real emission term in a factorised form:

$$
d\Phi_g = J \times d\Phi_B \frac{d^3 k}{k_0} \tag{2}
$$

through the choice of a mapping to an underlying Born $\Phi_{g} \leftrightarrow {\Phi_{B}, k}$, (or choice of a recoil scheme), it can be shown that if the mapping is linear in k for small k , no linear renormalons are present after the k integration. So: in inclusive cross sections at fixed undelying Born no renormalons are present.

▶ Virtual corrections due to the exchange of a massive gluon emitted by massless partons never lead to linear terms in the mass λ . Besides verifying this in the practical case, this can be proven by considering that the Passarino-Veltman reduction procedure never leads to linear terms in λ , and by examining the IR divergent scalar integrals.

Hard collinear region

 \blacktriangleright Hard, collinear divergences do not lead to linear terms λ . In fact, defining Sudakov variables for the gluon momentum

$$
k = zp_1 + \beta p_2 + k_{\perp}, \qquad \beta = \frac{k_{\perp}^2 + \lambda^2}{z2p_1 \cdot p_2},
$$

collinear integrals have the form

$$
\int \mathrm{d}k_{\perp}^2 \frac{P(k_{\perp})}{\left(k_{\perp}^2 + (1+z)\lambda^2\right)^i}, \qquad i = 1, 2.
$$

 $P(k_{\perp})$, for small k_{\perp} , can start with a constant if $i = 1$, and must start with a term bilinear in k_\perp or proportional to λ^2 if $i=2.$ If the mapping near the collinear region is linear in k_+ , no linear terms in λ can arise, since subleading terms in k_{\perp} are odd, and vanish by azymuthal integration.

The soft region

The soft region leads to integrands of the form

$$
\int \frac{\mathrm{d}^3 \vec{k}}{\omega} P(k) \left[\frac{1}{(2p_1 \cdot k + \lambda^2)(2p_2 \cdot k + \lambda^2)}, \frac{\lambda^2}{(2p_{1/2} \cdot k + \lambda^2)^2} \right]
$$

It is easy to see that (in the p_1, p_2 dipole CM)

$$
p_1\cdot k = \frac{Q}{2}\omega(1-\beta cos\theta) \geq \frac{Q}{2}\omega(1-\beta) \geq \frac{Q\lambda^2}{4\omega} \geq \frac{Q\lambda}{4}\,,
$$

so, the denominators scale at worse like λ , so does ω and $|\vec{k}|$. By power counting the second integral scales like λ^2 , while the first one scales like 1.

In the first integral, subleading terms in ω , for example, may lead to terms linear in λ .

We now consider the k integral in the soft region at fixed underlying Born. We assume that the mapping from the underlying Born phase space to the full phase space is smooth for small k , in the sense that

$$
p_i^{\mu} = \tilde{p}_i^{\nu} + T_i^{\mu\nu} k_{\nu} + \mathcal{O}(\omega^2)
$$

where \tilde{p} are the underlying Born momenta. The "dangerous" soft integral gives rise to terms of the form

$$
\frac{1}{p_1 \cdot k + \lambda^2} = \frac{1}{\tilde{p}_1 \cdot k} \left[1 - \frac{kT_i k + \lambda^2}{\tilde{p}_1 \cdot k} + \dots \right]
$$

so that

$$
\frac{1}{(p_1 \cdot k + \lambda^2)(p_1 \cdot k + \lambda^2)} = \frac{1}{\tilde{p}_1 \cdot k \, \tilde{p}_2 \cdot k} \left[1 - \sum_{i=1,2} \frac{kT_i k + \lambda^2}{\tilde{p}_i \cdot k} + \dots \right]
$$

where ... indicate terms subleading by power counting.

But, for collinear safety, $kTk \propto \tilde{p}_1 \cdot k \tilde{p}_2 \cdot k$, since it must vanish in both collinear limits. For example, T_1k must vanish if k is collinear to p_2 , because in this case $p_1 = \tilde{p}_1$, and must be proportional to \tilde{p}_1 if k is collinear to p_1 . Thus we end up having to worry about the following integrals

$$
\int \frac{\mathrm{d}^3 \vec{k}}{\omega} \left[\frac{1}{\tilde{p}_1 \cdot k \, \tilde{p}_2 \cdot k}, \quad \frac{k \cdot v}{\tilde{p}_1 \cdot k \, \tilde{p}_2 \cdot k}, \quad \frac{\lambda^2}{(\tilde{p}_1 \cdot k)^2 \, \tilde{p}_2 \cdot k} \right],
$$

where ν is a generic vector. Notice that one should also worry about the Jacobian, when changing integration variables from p, k to \tilde{p} , k. However, if the mapping is smooth in k, such change contributes at most a linear term, i.e. can be lumped into the $k \cdot v$ term.

By direct calculation, it can be easily seen that the above integrals do not yield linear terms in λ .

Consequences

Old and new results can be derived: Linear corrections are absent in

- ▶ DIS (must be the case because of the OPE)
- ▶ Drell-Yan total cross section [Beneke and Braun]
- ▶ Drell-Yan rapidity distribution [Dasgupta]
- ▶ Dreal-Yan double differential cross section in transverse momentum and rapidity distribution of the pair (new)
- \triangleright In e^+e^- , shape variables power corrections can be computed also in the 3-jet regime! Before they had been computed only in the 2-jet limits, with

the only exception of the C-parameter in the 3-jet symmetric limit [Luisoni,Monni,Salam,2019]

The results on DIS and Drell-Yan follow because on can find an appropriate mapping that also maintains fixed $\,Q^2$ and x_{bj} for DIS, and the Drell-[Yan](#page-15-0) pair kinematics for Drell-Yan[.](#page-17-0) We investigating the structure of linear renormalons in the three jet region by computing the cross section for the process $\gamma^* \to q\bar{q}\gamma$

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including gluonic corrections in the large n_f limit.

Our new analytic findings can also be applied to shape variables. The large- $n_{f},\ T_{O}(\lambda)$ result can also be written as

$$
T_O(\lambda) = N \int d\Phi_3 \Bigg\{ \int d\Phi_k^{(\lambda)} M_{\mu\nu}(k) \Bigg[\int d\Phi_{\text{split}} \overbrace{P_{\text{split}}^{\mu\nu}(O_5 - O_4)}^{\Delta(\lambda)} + \overbrace{O_4 g^{\mu\nu}}^{R(\lambda)} \Bigg] + \overbrace{V_{\lambda} O_3}^{V(\lambda)} \Bigg\}.
$$

where $O_{3.5}$ is the observable in terms of 3, 4 or 5 particles, and:

- \blacktriangleright $M_{\mu\nu}g^{\mu\nu}$ is the cross section fo the production of the $q\bar{q}\gamma$ system plus a massive gluon with momentum k and mass λ
- \blacktriangleright $M_{\mu\nu}P^{\mu\nu}_{\rm split}$ is the square amplitude for the production of $q\bar{q}\gamma$ plus a $q\bar{q}$ pair by a (massless virtual) gluon of momentum k $(k^2 = \lambda^2)$ via splitting, but normalized so that

$$
\int {\rm d}\Phi^{\mu\nu}_{\rm split} P^{\mu\nu}_{\rm split} = g^{\mu\nu} - k^\mu k^\nu/\lambda^2
$$

▶ V_{λ} is the virtual corrections to the $\gamma^* \to q\bar{q}\gamma$ process, due to the exchange of a massive gluon. **KORK SERVER STARK**

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$$

We can rewrite it as

$$
T_O(\lambda) = N \int d\Phi_3 \left\{ \int d\Phi_k^{(\lambda)} M_{\mu\nu}(k) \left[\int d\Phi_{split} P_{split}^{\mu\nu}(O_5 - O_4) + (O_4 - O_3) g^{\mu\nu} \right] + \left[\int d\Phi_k^{(\lambda)} M_{\mu\nu} g^{\mu\nu} + V_{\lambda} \right] O_3 \right\}.
$$

Now, the second line contains the sum of the virtual plus an unrestricted integral in k of the massive real contribution. By our findings, if the mapping $\Phi_3, \Phi_k^{(\lambda)}$ $\kappa^{(\lambda)}$ is smooth, It does not have linear terms in $\lambda!$

Focus upon the first line:

$$
\mathcal{T}_{O}^{(\text{first})}(\lambda) = N \int \mathrm{d}\Phi_3 \Bigg\{ \int \mathrm{d}\Phi_k^{(\lambda)} M_{\mu\nu}(k) \Bigg[\int \mathrm{d}\Phi_{\text{split}} P^{\mu\nu}_{\text{split}}(O_5-O_4) + (O_4-O_3) g^{\mu\nu} \Bigg] \Bigg\}
$$

Since O is IR safe, there is one soft suppression from there. So, we can evaluate M neglecting $\mathcal{O}(1)$ terms in ω :

$$
T_O^{\text{(first)}}(\lambda) \approx N \int \mathrm{d}\Phi_3 B \int \mathrm{d}\Phi_k^{(\lambda)} P_{\mu\nu}^{\text{(soft)}} \Bigg[\int \mathrm{d}\Phi_{\text{split}} P_{\text{split}}^{\mu\nu} (O_5 - O_3) \Bigg]
$$

where we only need the Born cross section B , the soft emission tensor $P^{\mathrm{(soft)}}_{\mu\nu}$, and the splitting factor $P^{\mu\nu}_{\mathrm{split}}$:

$$
\begin{array}{ccl}\nP_{\mu\nu}^{\text{(soft)}} & = & \left(\frac{p_1^{\mu}}{(p_1+k)^2}-\frac{p_2^{\mu}}{(p_2+k)^2}\right)\left(\frac{p_1^{\nu}}{(p_1+k)^2}-\frac{p_2^{\nu}}{(p_2+k)^2}\right), \\
P_{\text{split}}^{\mu\nu} & = & N \text{Tr}[\cancel{q}\gamma^{\mu}(\cancel{k-q}\gamma\gamma^{\nu})].\n\end{array}
$$

We can evaluate numerically $\tau^{\rm (first)}_{O}$ $\sigma^{\text{(first)}}(\lambda) - \mathcal{T}_\mathcal{O}^{\text{(first)}}$ $O^{(\text{inst})}(0)$, (canceling the constant terms under the integral sign) to get the linear term.

Very early approaches [Dokshitzer,Webber,Marchesini 95] on non-perturbative corrections near the 2-jet limit suggested the formula

$$
\mathrm{d}\eta\,\mathrm{d}^2k_\perp\,\left(\frac{1}{k_\perp^2}\right)\,[O(P,k)-O(\rho)]\alpha(k_\perp)
$$

where the first term is the invariant phase space for soft emission, the term in the round bracket is the amplitude for soft emission in the eikonal approximation in the radiating dipole rest frame, and the term in square bracket is the contribution of the observable.

The ambiguity associated with the integration near the Landau pole for α_s corresponds to the linear power correction.

For example:

$$
1 - T: O(P, k) - O(p) = \frac{k_{\perp}}{Q} \exp(-|\eta|), \quad \int d\eta \exp(|\eta|) = 2,
$$

$$
C: O(P, k) - O(p) = \frac{k_{\perp}}{Q} \frac{3}{\cosh(\eta)}, \quad \int d\eta \frac{1}{\cosh(\eta)} = 3\pi
$$

These approaches ignored the gluon virtuality, that was set to zero in the formula, and the dependence of the shape variable upon the products of the gluon decaying into massless partons, that spoils the universality in the above formula [Seymour,PN,95].

Subsequently, Dokshitzer,Lucenti,Marchesini,Salam, 97-98 demonstrated that for a wide class of observables the inclusion of the gluon splitting process changed the original formula by a universal, constant factor, that was dubbed the Milan factor.

For this to hold, the observable must be additive under multiple soft emission, i.e.

$$
O(P, k_1, \ldots, k_n) \approx O(P_1, k_1) + \ldots + O(P_n, k_n)
$$

(among the early approaches, some advocated the use of a massive gluon [Akhoury,Zakarov,95]. This leads to different results, and their universality cannot be granted).

We found [Caola,Ferrario Ravasio,Limatola,Melnikov,Ozcelik,PN] that the also in the 3-jet limit the Milan factor formula can be derived (with the same Milan factor as in the 2-jet case)

$$
|O|_{\mathrm{NP}} = \mathcal{MI}_{\mathrm{NP}} \int \mathrm{d}\sigma_B(\Phi_B) \mathcal{T}^{\lambda} \left[\sum_{\mathrm{dip}} \int [\mathrm{d}k] \frac{M_S}{\alpha_S} \delta(|k_{\perp}| - \lambda) [O(\Phi_{B,k}) - O(\Phi_B)] \right]
$$

where $k^2=0$, and

$$
[\mathrm{d}k]\frac{M_s}{\alpha_s} = \frac{\mathrm{d}^3k}{2k^0(2\pi^3)}C_{\mathrm{dip}}\frac{g_s^2}{\alpha_s}\frac{p_1\cdot p_2}{p_1\cdot k\,p_2\cdot k} = \frac{2C_{\mathrm{dip}}\alpha_s}{\pi}\mathrm{d}\eta\,\frac{\mathrm{d}\phi}{2\pi}\frac{\mathrm{d}k_\perp}{k_\perp}
$$

- ▶ Notice the trade: $k^2 = \lambda^2 \rightarrow k_{\perp} = \lambda$.
- ▶ The shape variable is evaluated for the extra-emission of one massless parton.
- \blacktriangleright The proof is not simple. However it is clear how the addittivity of the shape variable makes this possible.
- \triangleright An extra condition emerges: the rapidity integral must converge.

For the cumulant of a shape variable, we obtain

$$
\Sigma(v)_{\mathrm{NP}} = \left\{\int \mathrm{d}\sigma_B(\Phi_B)\delta(v(\Phi_B) - v) \sum_{\mathrm{dip}} \left[-\mathcal{M} \frac{2C_{\mathrm{dip}}}{2\pi} \int \frac{\mathrm{d}\phi}{2\pi} \mathrm{d}\eta h_v(\eta, \phi) \right] \right\} \mathcal{I}_{\mathrm{NP}}
$$

where

$$
h_{\nu}(\eta,\phi)=\lim_{k_{\perp}\to 0}\left(\frac{Q}{|k_{\perp}|}\left[\nu(P,k)-\nu(p)\right]\right)
$$

and the function h is easily calculable for the shape variables of interest.

The calculation of the linear power corrections for the $\gamma^* \to q \bar{q} \gamma$ production process only involves the radiation from the $q\bar{q}$ dipole in the soft approximation.

This result suggests the generalization to the realistic $\gamma^* \to q \bar q g$, applying the soft approximation to this case.

Thus one can simply add the contributions arising from each one of the final state colour dipoles, i.e. $q\bar{q}$, $q\bar{q}$ and $\bar{q}g$.

Shape variables are IR stable functions of the final state kinematics. We consider:

$$
\blacktriangleright \text{ Thrust: } \tau = 1 - T, \ T = \max_{\hat{t}} \sum |\vec{p}_i \cdot \hat{t}| / \sum |\vec{p}_i|
$$

$$
\triangleright \ \mathsf{C} \colon \Theta^{\alpha\beta} \sum_{i} \frac{p_i^{\alpha} p_i^{\beta}}{|\vec{p}_i|} / \sum_{i} |\vec{p}_i|, \ \mathsf{C} = 3(\lambda_1 \lambda_2 + \lambda_1 \lambda_3 + \lambda_2 \lambda_3)
$$

- \triangleright y_3 : take Durham jet clustering, with distance measure $y_{ij} = 2min(E_i^2, E_j^2)(1 - \cos\theta_{ij})/Q^2$. Then y_3 is define as the value of y_{ii} at the clustering step that leads to the transition from 3 to 2 clusters.
- ▶ Mh2 (heavy jet mass): the heaviest of the squared masses of the two hemisperes defined by the plane orthogonal to the thrust axis, normalized to Q^2
- \triangleright Md2: the heaviest minus the lightes of the squared masses of the two hemisperes, normalized to Q^2 .

► Bw: max(
$$
B_1, B_2
$$
), $B_i = \sum_{p_k \in H_i} |\vec{p}_k \times \hat{t}| / (2 \sum_i |\vec{p}_i|)$.

Non-perturbative corrections can be parametrized as a shift in the perturbative cumulant distribution:

$$
\Sigma(s) \longrightarrow \Sigma(s+H_{\text{NP}}\zeta(s)), \quad \text{where} \quad \Sigma(s) = \int \mathrm{d}\sigma(\Phi)\theta(s-s(\Phi))
$$

and $H_{\text{NP}} \approx \Lambda/Q$ is a non-perturbative parameter that must be fitted to data.

The dot in the plots represents the constant value that was used in earlier studies. The value of $\zeta(c)$ at the symmetric point $c = 3/4$ was also computed by Luisoni,Monni,Salam [20](#page-27-0)[21.](#page-29-0)

Near $v = 0$, the Born amplitude is dominated by the soft-collinear region.

$$
radiation = \frac{C_A}{2} M_{\bar{q}_g} + \frac{C_A}{2} M_{qg} + \left(C_F - \frac{C_A}{2} \right) M_{q\bar{q}}
$$

but $M_{qg} \approx 0$, $M_{\bar{q}_g} \approx M_{q\bar{q}}$, so the total is $\approx C_F M_{q\bar{q}}$.

Our $\zeta(v)$ functions, for $v \to 0$ MUST approach the 2-jet limit value; but up to single logs!, i.e. terms of relative order $1/|\log(v)|$.

Singularity at the origin

Insist on $v \rightarrow 0$ (quadruple precision, log scale histogram). Two-jet limit reached, but subleading terms are extremely important! Singularity compatible with a form

$$
\frac{\log v + C}{\frac{v}{\log v + B}}
$$

for $B \neq C$.

- \blacktriangleright Historically the framework of choice to measure α_s directly from the $q\bar{q}g$ vertex.
- \blacktriangleright In practice: very convincing at the 10% level; affected by non-perturbative uncertainties if one wants higher precision
- \triangleright $\alpha_s(M_Z)$ from NNLO+NLL+Monte Carlo models:
	- ▶ 0.1224 \pm 0.0039 ALEPH 2009, [arXiv:0906.3436].)
	- ▶ 0.1189 \pm 0.0043 OPAL 2011, [arXiv:1101.1470])
	- ▶ 0.1172 ± 0.0051 JADE 2009, [arXiv:0810.1389]

The use of Monte Carlo models to correct for hadronization effects have long been criticized, since the interplay of perturbative and non-perturbative effects in Shower Monte Carlo is not fully clear.

 α_s from e^+e^- shape variables

As an alternative, another class of determinations is based upon analytic modeling of non-perturbative effects, using methods like SCET, dispersive models and low scale QCD effective couplings, and using $NNLO+N^3LL$ calculations:

- \triangleright 0.1135 \pm 0.0011 R. Abbate *et al*, 2011, [arXiv:0809.3326]
- \triangleright 0.1134 $^{+0.0031}_{-0.0025}$ Gehrmann,Luisoni,Monni, 2013,[arXiv:1210.6945]
- ▶ 0.1123 \pm 0.0015 Hoang et al, 2015 [arXiv:1501.04111]

They tend to result in a rather low value, not in good agreement with world data.

Results from Zanderighi, PN 2023

Simultaneous fit to C, t and y_3 , both for our newly computed $\zeta(v)$, and, for comparison, with $\zeta(v) \to \zeta_{2J}(v) = \zeta(0)$ (traditional method for the computation of power corrections).

The central value is at $\alpha_s(M_Z) = 0.1174$, $\alpha_0 = 0.64$. The "traditional" method leads to smaller values [of](#page-33-0) $\alpha_{S_{\rm t},\pi}$

Results from Zanderighi, PN 2023

Individual fits:

Only the combination of the three observables leads to a sensible determination of α_s

Inclusion of all data we could find at all energies

DELPHI 91.2 45 66 76 133 161 172 183 189 192 196 200 202 205 207 ALEPH 91.2 133 161 172 183 189 200 206 OPAL 91.2 133 177 197 L3 91.2 41.4 55.3 65.4 75.7 82.3 85.1 130.1 136.1 161.3 172.3 182.8 188.6 194.4 200 JADE 22 35 44 TRISTAN 58 JADEOPAL 91.2 35 44 133 161 172 183 189 SLD 91.2

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- \triangleright We perform the fits at the central scale, and then consider its variations by a factor of 2 below and above.
- \blacktriangleright In order to get a better fit of the very precise Z-peak data, we chose the central scale to be a function of the shape variable:

We first compute the average k_T as a function of the value of each shape variable (computed at the LO level), and then choose the k_T as central value of the scale. Fitting only ALEPH data on the Z peak we get:

(best μ_R leads to the lowest χ^2)

▶ The lower limit in the fit range is taken at twice the Sudakov peak position. Upper limit is 0.6 for C and 0.3 for thrust and y3.

PRELIMINARY PLOTS

Leading to $\alpha_{\mathcal{s}}(\mathcal{M}_Z)=0.1180$, and $\alpha_0=0.589$, with $\chi^2=1125.7$ over 895 degrees of freedom $(\chi^2/{\rm dof}=1.258)$.

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Fits to individual observables: C

 $\alpha_{\mathcal{\mathcal{S}}}(M_Z)=0.1172$, and $\alpha_0=0.6148$, with $\chi^2=269.0$ over 292 degrees of freedom $(\chi^2/{\rm dof}=0.921).$

Fits to individual observables: T

 $\alpha_{\mathcal{\mathcal{S}}}(M_Z)=0.1169$, and $\alpha_0=0.6245$, with $\chi^2=651.7$ over 151 degrees of freedom $(\chi^2/{\rm dof}=1.442).$

Fits to individual observables: y3

 $\alpha_{\mathcal{\mathcal{S}}}(M_Z)=0.1155$, and $\alpha_0=0.4151$, with $\chi^2=71.6$ over 292 degrees of freedom $(\chi^2/{\rm dof}=0.474).$

All together

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- \triangleright For C and T, individual fits are compatible with the CTv3 fit,
- \blacktriangleright For v_3 a much smaller α_0 is favoured.
- \blacktriangleright The χ^2 for the individual y_3 fit is very low, so that a larger value of α_0 (leading to a larger value of α_s) is also acceptable.
- \blacktriangleright The inclusion of y_3 in the CTy3 fit pulls α_0 to smaller values, and thus increases $\alpha_{\rm s}$ slightly.

Variations: global fit CTy3

We have considered:

- ▶ Scale variations, up and down by a factor of 2 from default
- \triangleright Mass scheme: how to solve the ambiguity in shape variables due to hadron masses [Salam,Wicke,2001]. We use as default the E scheme; variations: std. scheme, p scheme and D scheme.
- \blacktriangleright Range low limit: 2 (default), 1.7, 3 times the peak position.
- ▶ 4 different ways to implement NP corrections: shift in the full cumulant with or without adding an estimate of quadratic terms; shift in the LO cumulant; expand the correction around the perturbattive value. cumulant argument
- ▶ Subtract NP error
- ▶ Add NP error

In all cases we find the mass scheme issue very disturbing (slightly less than a 2% correction in both directions.

2-jet NP correction, fit CTy3

Everything else being equal, we found that using the two-jet limit NP correction lowers the value of α_s by nearly 0.002 in the CTy3 fit.

For all fits:

Stronger decrease for C, less for T, almost none for y_3 .

- ▶ These observables have $\zeta(v) < 0$ in most of the range, with an effective large jump near the origin. They are not easy to fit as the others, perhaps because of this jump.
- \blacktriangleright Unfortunately, the data is inconsistent for these observables, with DELPHI differing strongly from all other experiments.
- \triangleright An independent analysis of DELPHI data on the Z peak [Daniel Wicke, 1999 (thesis)] is instead consistent with the other experiments.

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Including M_h^2 and M_d^2 in the fit

Including M_h^2 and M_d^2 in the fit, 2-jets NP corrections

CONCLUSIONS

- ▶ Something new has been understood in the framework of power corrections for collider processes.
- ▶ Applications in e^+e^- shape variables seem to support these findings to some extent.
- ▶ Prospects for fits to α_s in e^+e^- framework are unclear because of:
	- ▶ Hadron mass effects are poorely understood.
	- ▶ Interplay of (new) NP corrections and resummation needs more work
	- ▶ Possible inconsistencies between experiments should be carefully assessed.
	- ▶ Correlations in data (and theory) needs a better treatment.

BACKUP

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2-jet NP correction, fit CT

2-jet NP correction, fit C

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2-jet NP correction, fit T

2-jet NP correction, fit y3

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The generic statement that can be made for massless partons cannot be generalized to the massive case. Nevertheless, with a reasoning inspired by the Low-Burnett-Kroll theorem, some results can be obtained also in this case. In particular:

- \triangleright The absence of linear renormalons can be derived for B meson decays, as long as the B mass is expressed in a short-distance scheme (like the $\overline{\text{MS}}$ one). This result was already obtained by Beneke, and it also follows from the existance of an OPE for includive B decays.
- \blacktriangleright The absence of linear renormalon in the *t*-channel, total single top cross section (if m_t is in a short distance scheme!), and the computation of linear corrections in the top differential distributions.
- ▶ The absence of linear renormalons in $q\bar{q} \rightarrow t\bar{t}$ total cross section (again with m_t in a short distance scheme), and the computation of linear corrections in the top differential distributions.

Single Top

The result for the differential distribution can be expressed as a shifts in the argument of the Born cross section. For the transverse momentum and rapidity of the top the shift are given by

$$
\frac{\delta_{\rm NP}[\rho_\perp]}{\rho_\perp} = \frac{\alpha_s C_F \pi \lambda}{2\pi m_t} \n\delta_{\rm NP}[y_t] = \frac{\alpha_s C_F \pi \lambda}{2\pi m_t} \times \frac{8m_t^2 \, s \, ch^2(y_t)}{(s + m_t^2)^2}
$$

Since we have

$$
\frac{\delta_{\rm NP} m_t}{m_t} = \frac{\alpha_s C_F}{2\pi} \frac{\pi \lambda}{m_t}
$$

we can use current determination of the top quark pole mass renormalon uncertainty $\delta_{\rm NP} m_t = 0.1 - 0.2$ GeV to estimate these effects.

 $q\bar{q} \rightarrow t\bar{t}$

The results have a more interesting structure

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