

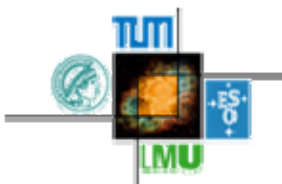
An Excursion into the Attouniverse and Beyond

Andrzej J. Buras

(Technical University Munich, TUM-IAS)

Schrödinger Gastprofessur

Vienna, 19.10.2010



Overture

1676

**A very important year for
the humanity !**

1676 : The Discovery of the Microuniverse (Animalcula) (The Empire of Bacteria)



Antoni van Leeuwenhoek
***24.10.1632 †27.08.1723**

10^{-6}m

~500 Microscopes

**(Magnification
by ~300)**

Microbe Hunters



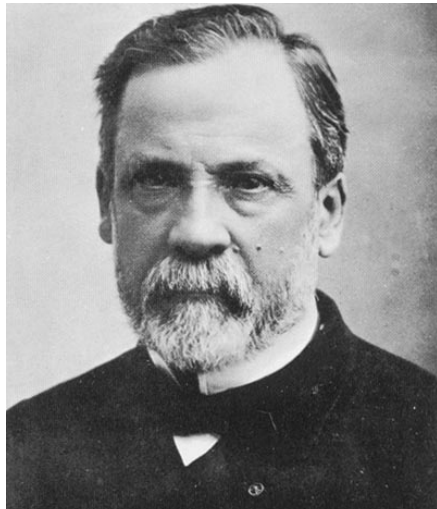
Antoni van Leeuwenhoek

***24.10.1632 †27.08.1723**



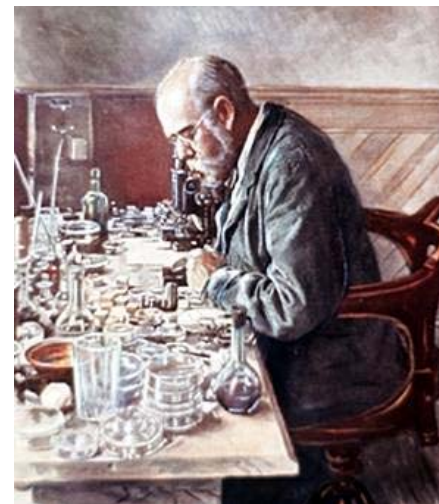
Lazzaro Spallanzani

***12.01.1729 †12.02.1799**



L. Pasteur

***27.12.1822 †28.09.1895**



Robert Koch

***11.12.1843 †27.05.1910**

An Excursion towards the Very Short Distance Scales:

1676 - 2020

Microuniverse

10^{-6}m

**Bacteriology
Microbiology**

Nanouniverse

10^{-9}m

Nanoscience

Femtouniverse

10^{-15}m

**Nuclear Physics
Low Energy Elementary
Particle Physics**

Attouniverse

10^{-18}m

**High Energy Particle
Physics (present)**

**High Energy Proton-Proton
Collisions at the LHC**

$5 \cdot 10^{-20}\text{m}$

**High Precision Measurements
of Rare Processes (Europe,
Japan, USA)**

10^{-22}m

**Frontiers of Elementary
Particle Physics in 2010's**

The Technology for the Microuniverse

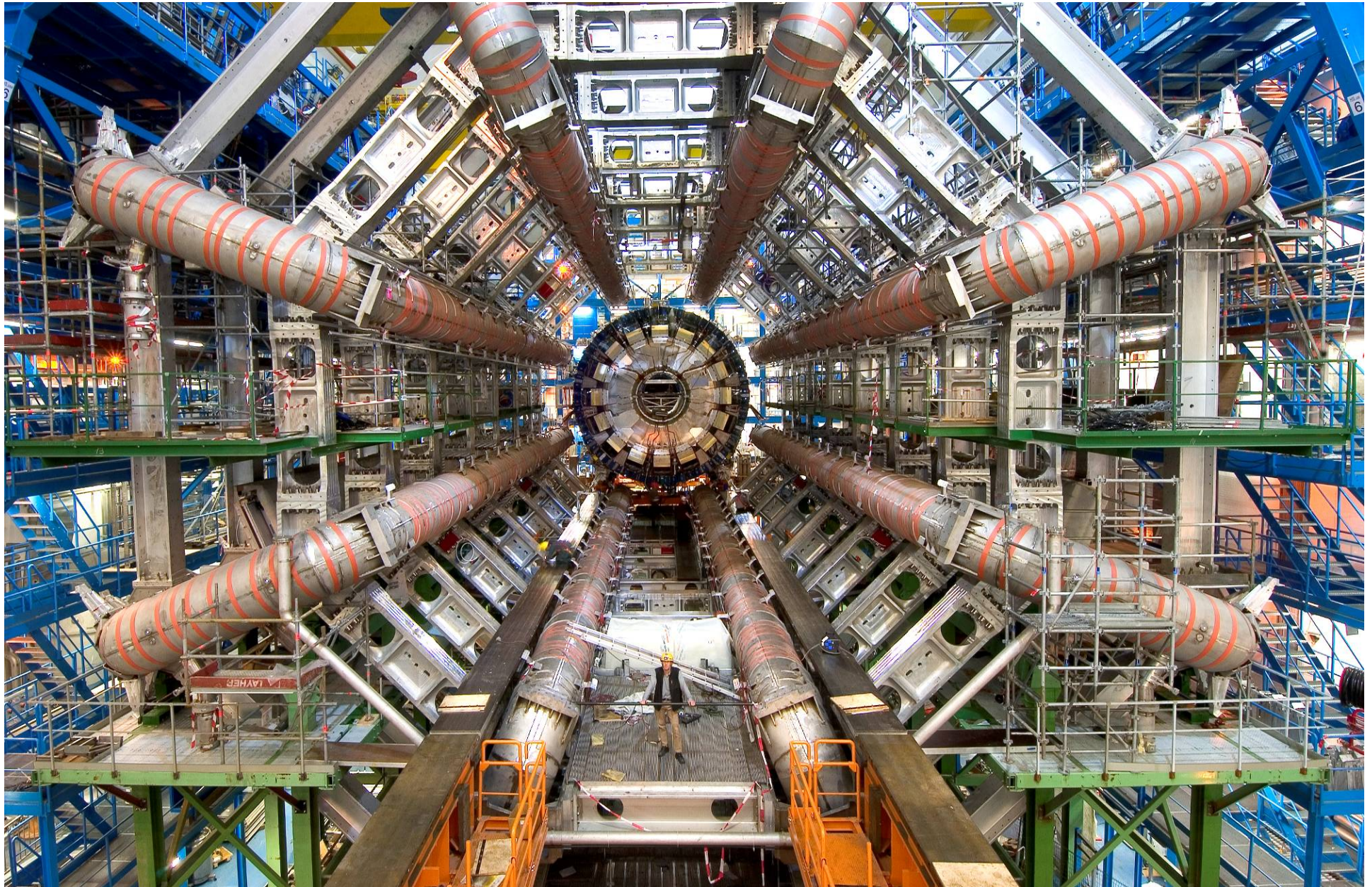


Robert Koch

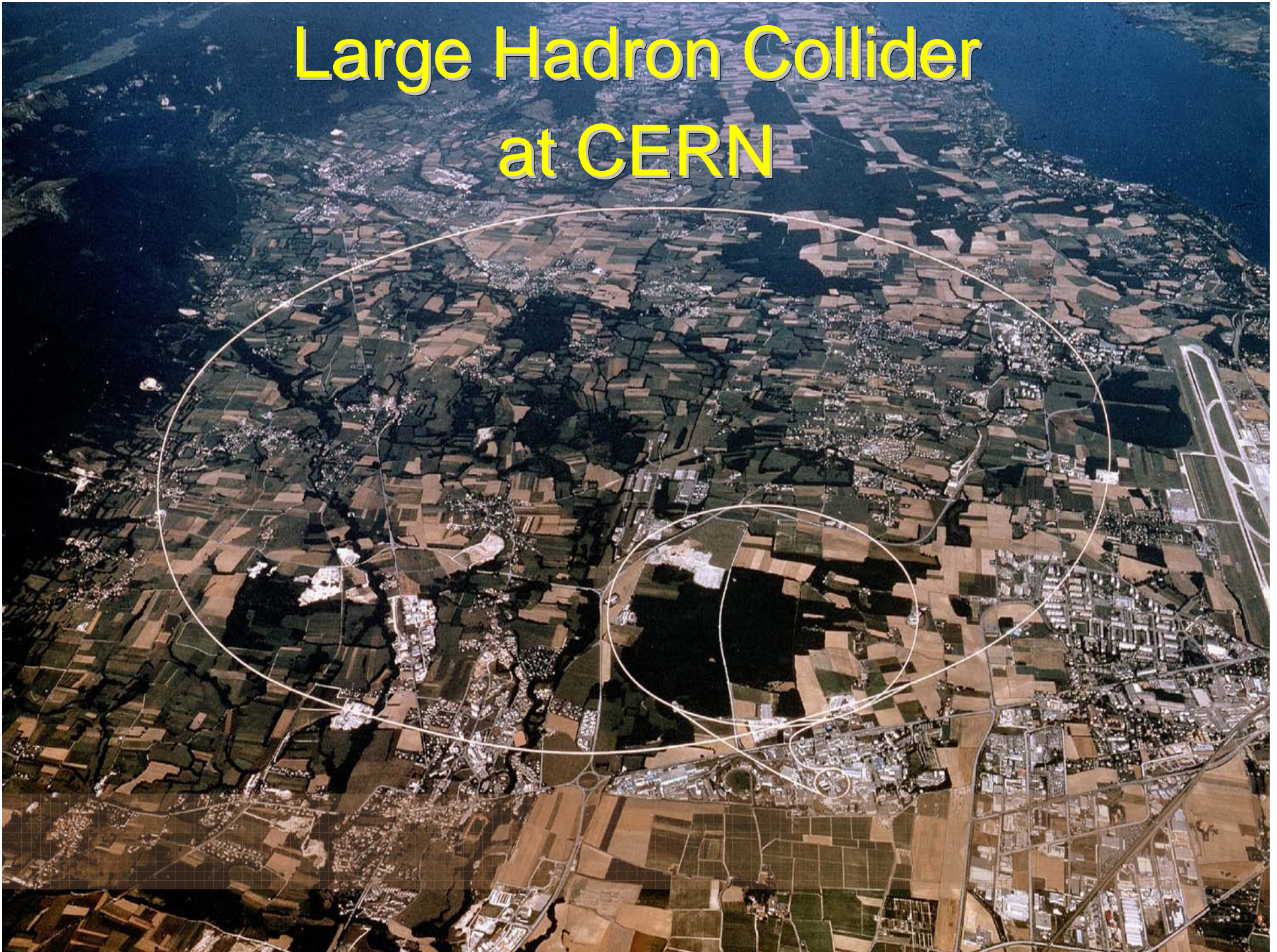
***11.12.1843 †27.05.1910**



Technology to go beyond the Attouniverse



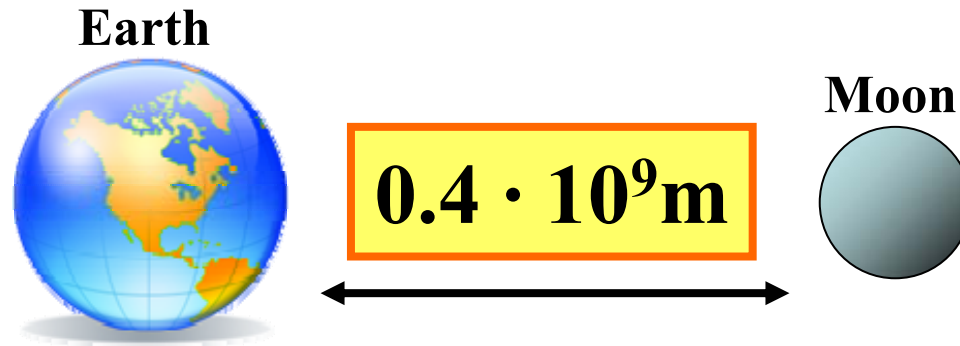
Large Hadron Collider at CERN



How small is really 10^{-18}m ?

How small is really 10^{-18}m ?

10^9m



3 light seconds

How small is really 10^{-18}m ?

10^9m



$0.4 \cdot 10^9\text{m}$



3 light seconds

10^{18}m

105 light years

How small is really 10^{-18}m ?

10^9m

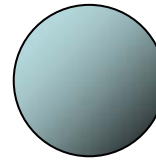
Earth



$0.4 \cdot 10^9\text{m}$



Moon



3 light seconds

10^{18}m

105 light years

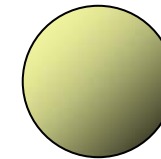
Earth



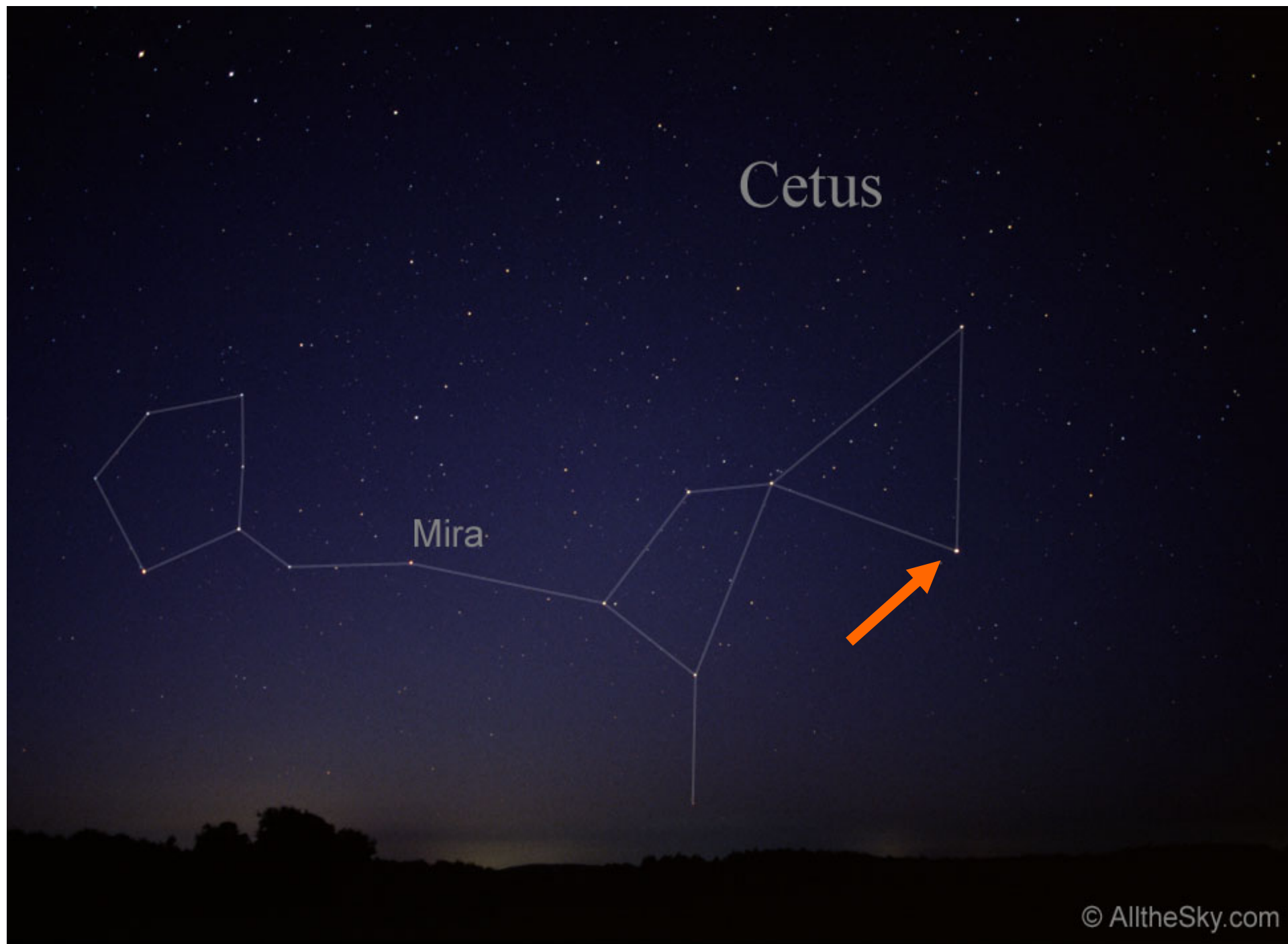
92 light years



Beta Ceti
(Deneb Kaitos
in the constellation
Cetus)





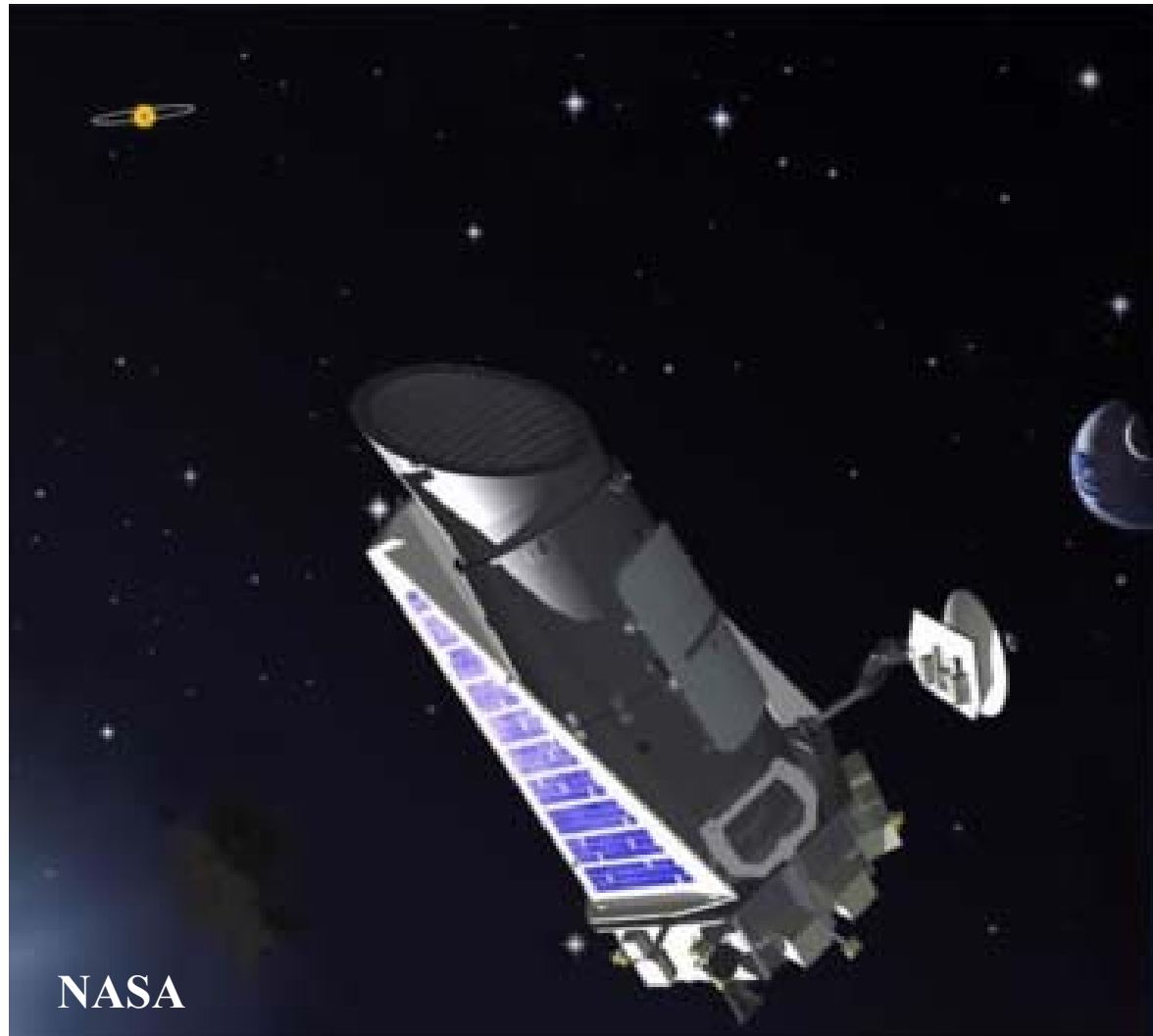


Cetus

Mira

© AlltheSky.com

Kepler Searching for new Planetary Systems and New Lives



March 2009

Most important Message from this Talk

Antoni van Leeuwenhook discovered in 1676

Das Reich des mikroskopisch Kleinen (Animalcula)

Elementary Particle Physicists expect to discover

New Animalcula in the coming years with the help

of **LHC** and **High Precision Experiments**

Next 57 min

**1st
Movement**

: Introducing the Attouniverse (33 min)

**2nd
Movement**

**: Expectations for New Animalcula
(10 min)**

**3rd
Movement**

**: First Messages from New Animalcula
(7 min)**

**4th
Movement**

: Final Messages (TUM, IAS, EC) (7 min)

1st Movement

Introducing the Attouniverse

How to explore the Attouniverse ?

Heisenberg's
Uncertainty
Principle

:

$$\underbrace{\Delta x}_{\text{Distance Interval}} \cdot \underbrace{\Delta p}_{\text{Momentum Interval}} \geq \frac{\hbar}{2}$$

$$\hbar = \frac{h}{2\pi}$$

h =Planck's
constant

Use $\hbar = c = 1$:

Speed
of light

$$E = mc^2$$

$$E = m$$

($p = 0$)

$$E = \sqrt{p^2 c^2 + m^2 c^4}$$

$$E = \sqrt{p^2 + m^2}$$

$$\approx \sqrt{p^2} \text{ for } p^2 \gg m^2$$

In order to explore:

Femtouniverse : $E \approx 200 \text{ MeV} \approx m_\pi$

Attouniverse : $E \approx 200 \text{ GeV} \approx m_{\text{top}}$
 10^{-19}m $E \approx 2000 \text{ GeV} = 2 \text{ TeV}$

$$\text{MeV} = 10^6 \text{eV}$$

$$\text{GeV} = 10^9 \text{eV}$$

(proton mass)

What do we know about the Attouniverse ?

Periodic Table of Elementary Particles

(Fermions : Spin $\frac{1}{2}$)

Charge

Families



1

2

ν_e

e^-

ν_μ

μ^-

ν_τ

τ^-

0

-1

Neutrinos

Charged
Leptons

3

4

u u u

d d d

c c c

s s s

t t t

b b b

$\frac{2}{3}$

$-\frac{1}{3}$

Quarks

Generations

1

2

3

6
Flavours

u = up
d = down
c = charm
s = strange
t = top
b = bottom

Masses of Elementary Particles

in units of $\text{GeV} = 10^9 \text{eV}$

$$\begin{pmatrix} m_\nu, \\ m_u, \\ \vdots \end{pmatrix}$$

Families



1

2

$\nu_e (\approx 0)$

$e^- (5 \cdot 10^{-4})$

$\nu_\mu (\approx 0)$

$\mu^- (0.105)$

$\nu_\tau (\approx 0)$

$\tau^- (1.78)$

$< 10^{-9}$

Very
hierarchical
structure

3

4

$u (3 \cdot 10^{-3})$

$d (6 \cdot 10^{-3})$

$c (1.3)$

$s (0.110)$





$t (170)$

$b (4.5)$

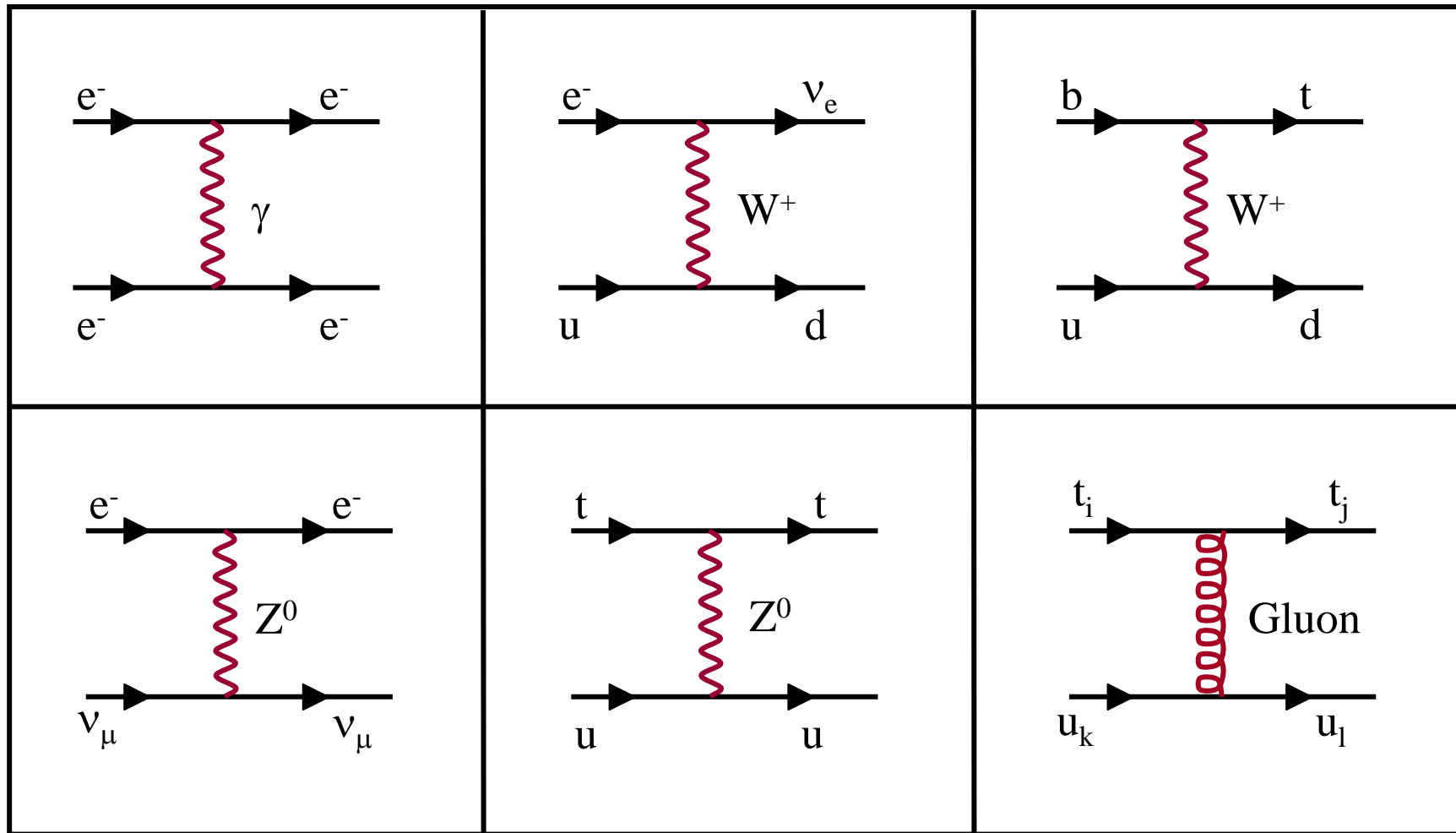
Particles in a given family distinguished only by the mass!

Interactions between Elementary Particles

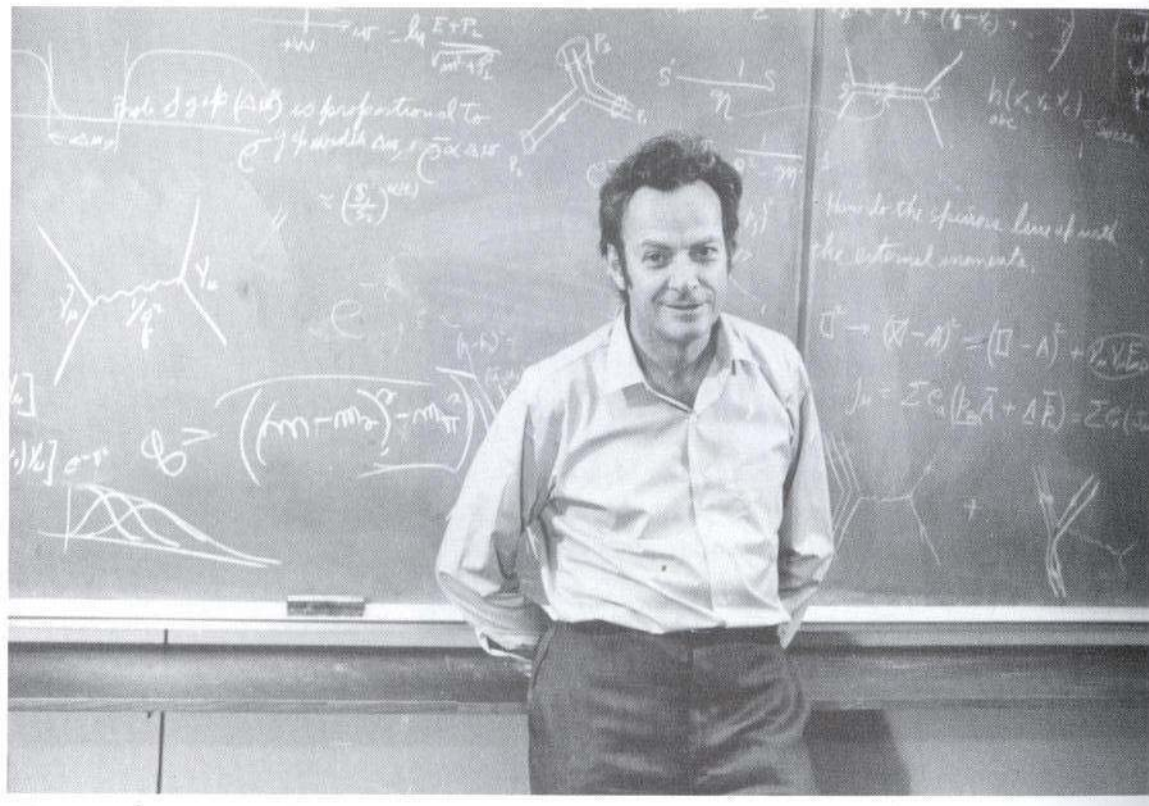
Mediated by Gauge Bosons (Spin 1)

	Photon	Charge = 0	(QED)	<div>(massless)</div> <div>Electroweak Interactions</div> <div>(massive) 80 GeV 91 GeV</div>
W^\pm 	W-Boson	Charge = ± 1	Weak Interactions	
Z^0 	Z-Boson	Charge = 0		
$8G$ 	G_a-Gluons ($a = 1, \dots, 8$)	Charge = 0 (massless)	(QCD)	<div>Strong Interactions</div> <div>Confinement of Quarks at “Large” (10^{-15}m) scales</div>
But carry new “Charge” - Colour				

Feynman Diagrams at Work



The Great Master at Work

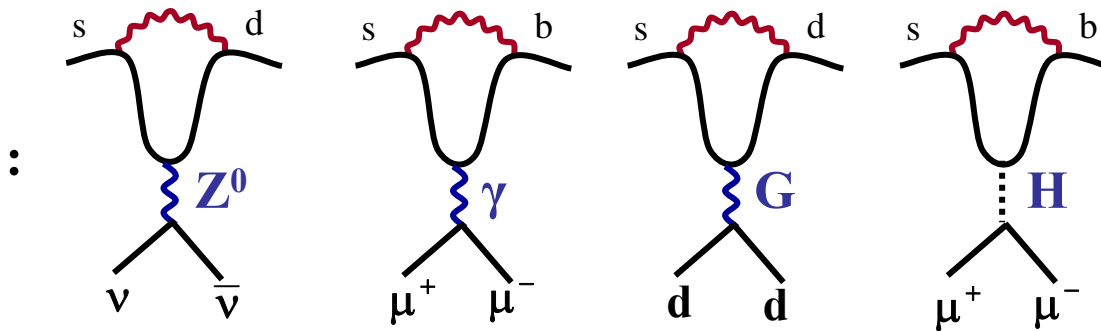


Richard Feynman

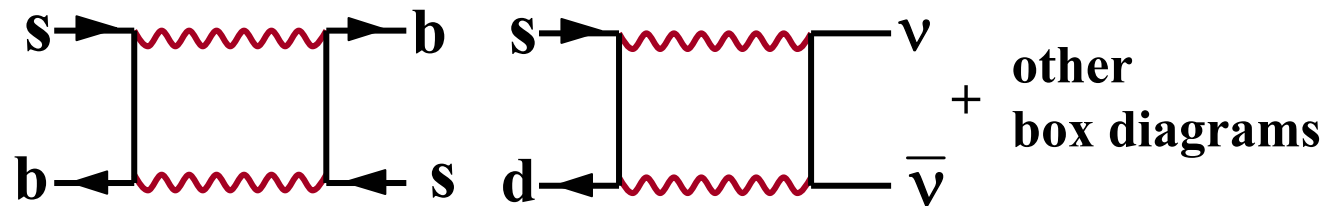
Phys. Rev. Bd 76 (1949) 769

More Complicated Feynman Diagrams

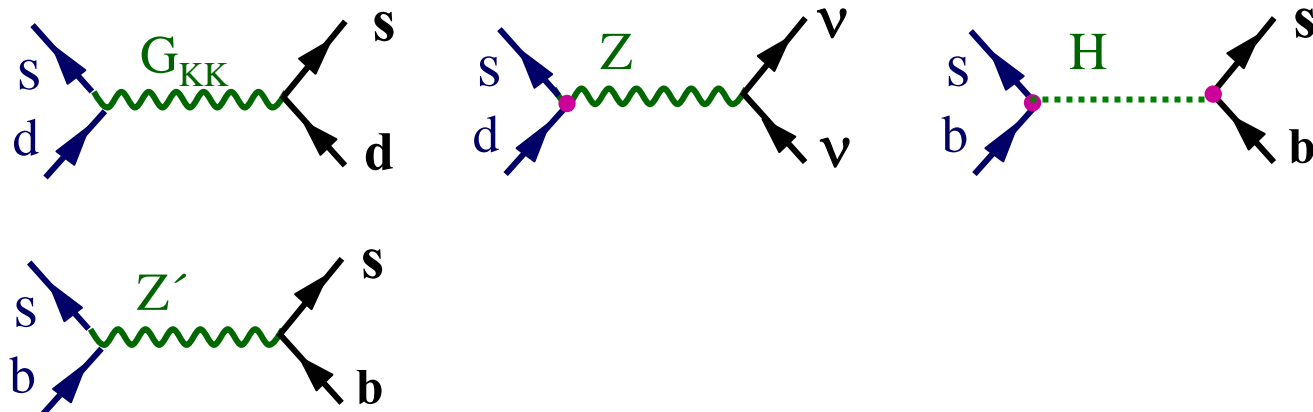
Penguin Family



Box Diagrams



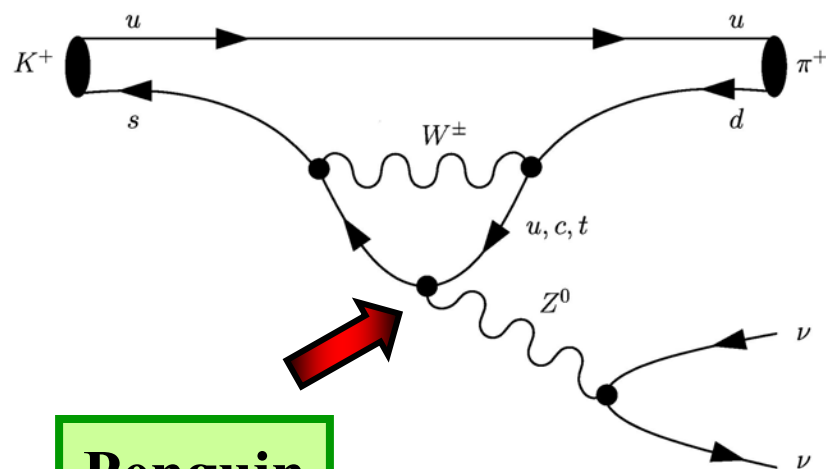
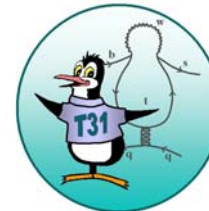
Tree Diagrams



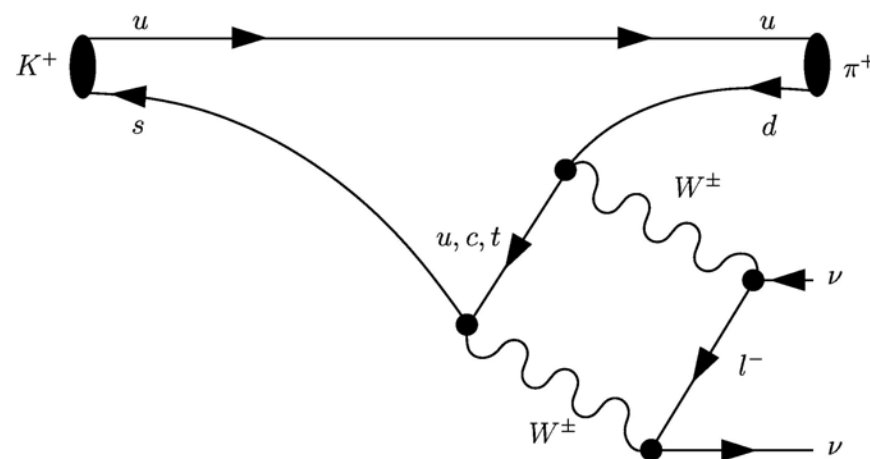
$$K^+ = (u\bar{s})$$

$$\pi^+ = (u\bar{d})$$

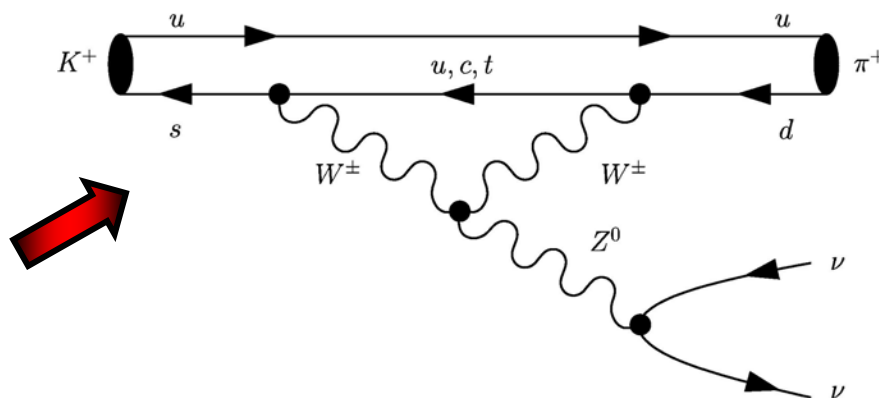
$$K^+ \rightarrow \pi^+ \nu \bar{\nu}$$



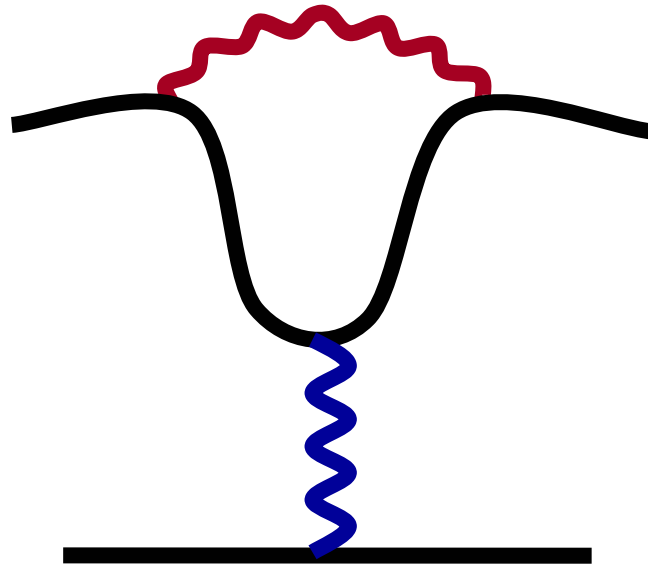
Penguin



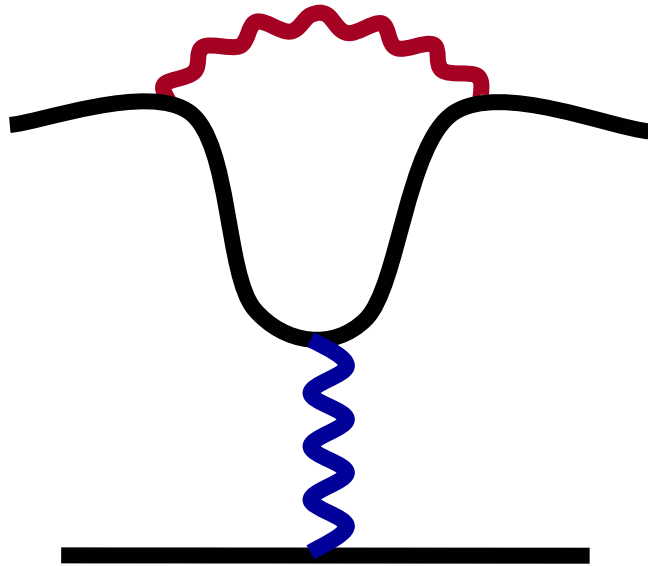
Box



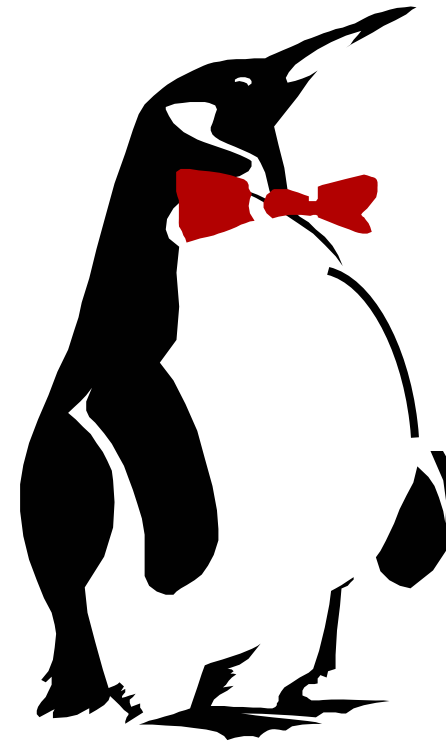
**Still
another
Penguin**



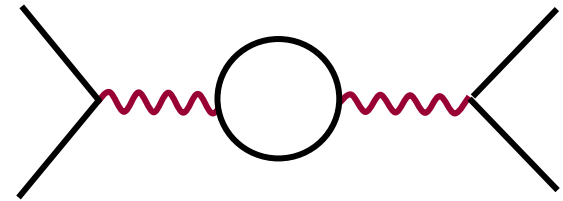
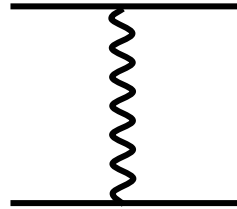
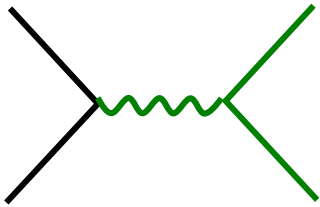
Penguin Diagram



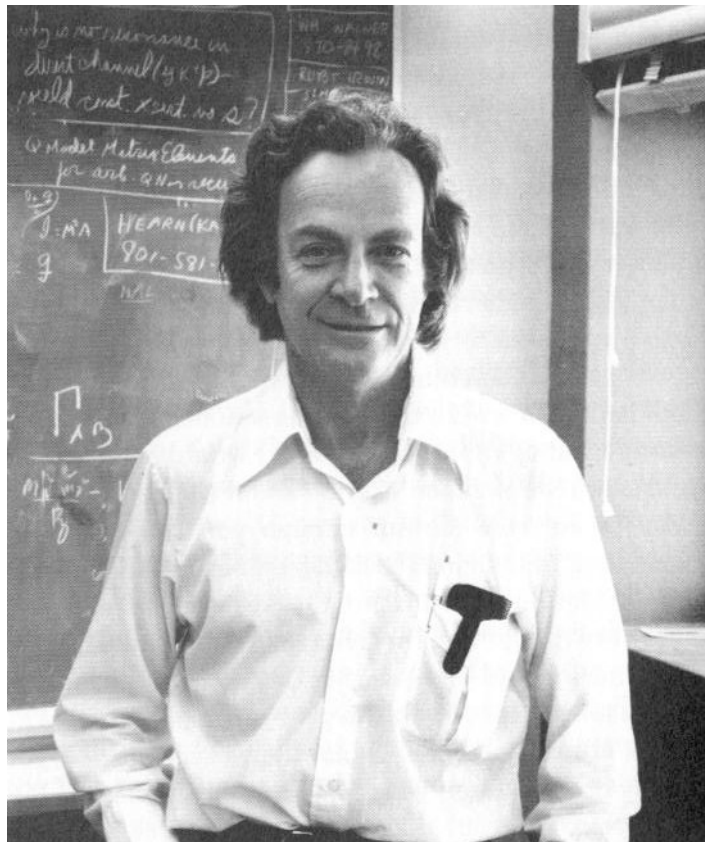
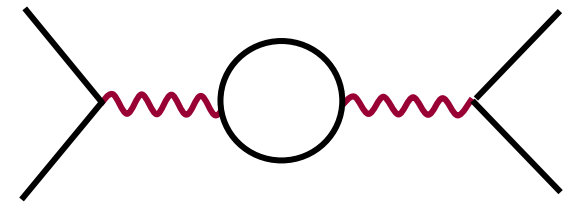
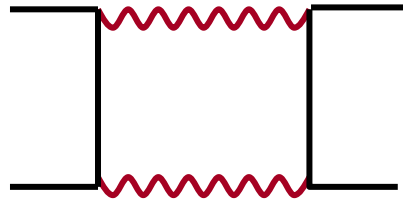
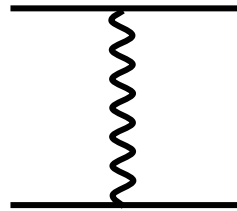
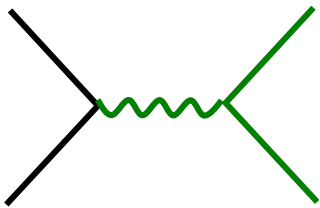
Penguin Diagram



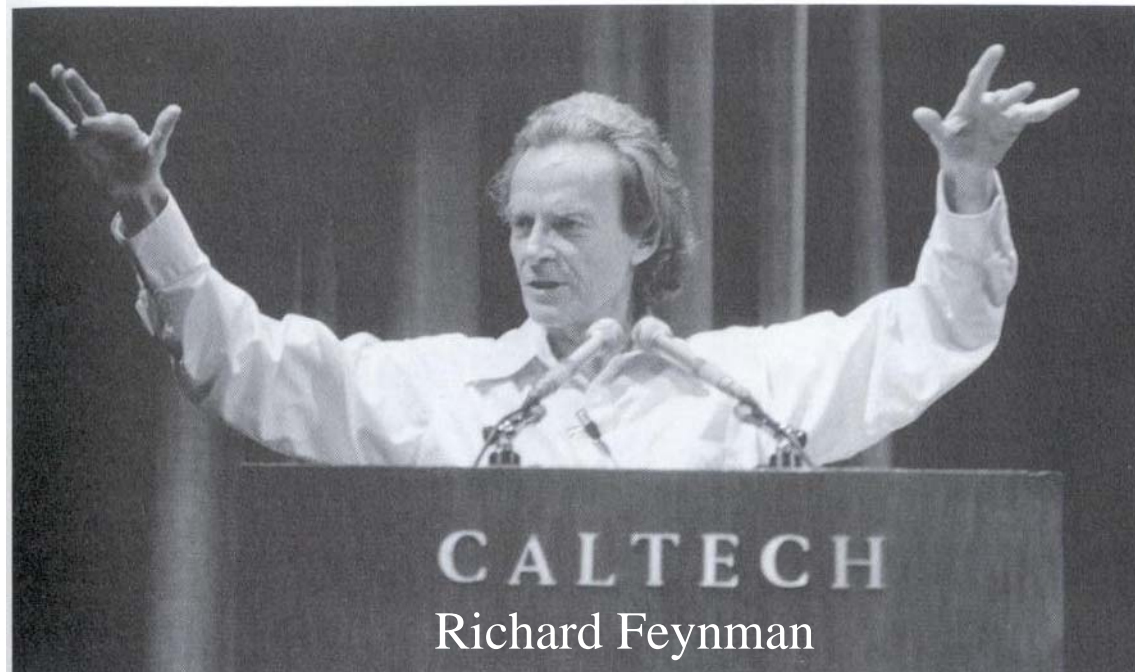
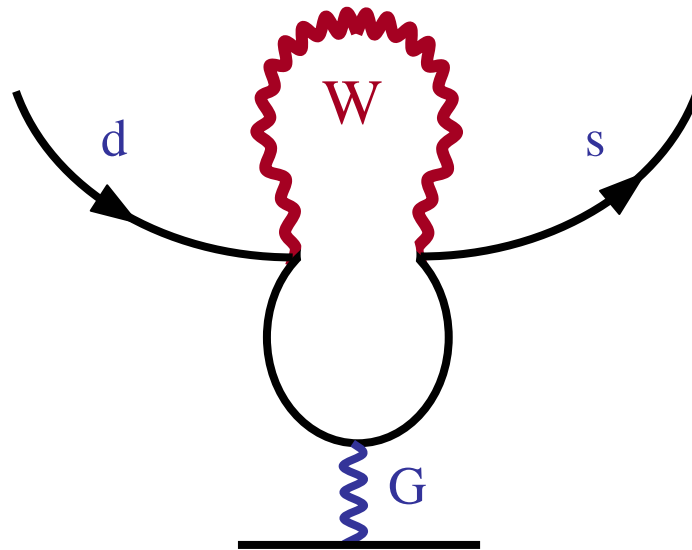
Feynman Diagrams



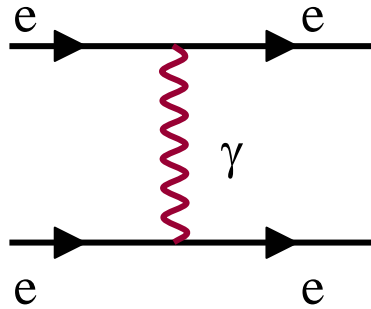
Feynman Diagrams



Richard Feynman

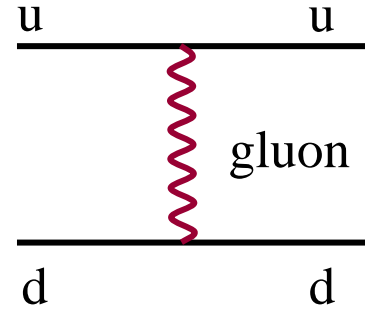


Strength of Interactions



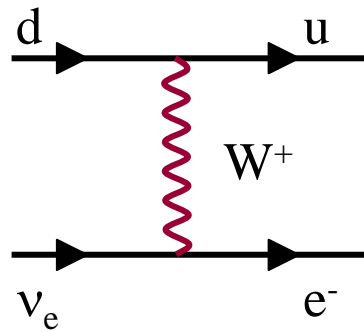
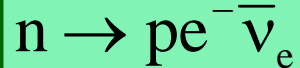
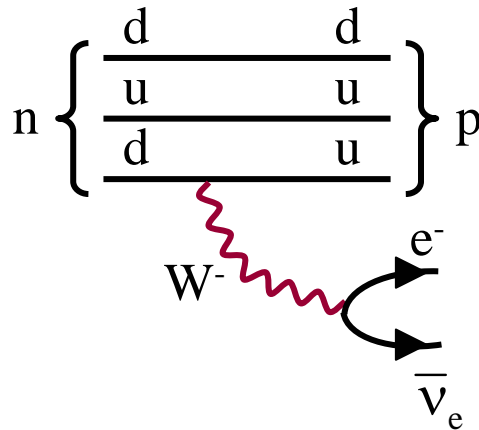
$$\alpha_{\text{QED}} \approx \frac{1}{137}$$

Sommerfeld's
Structure
Constant



$$\alpha_{\text{QCD}} \approx 0(1)$$

Strong Interactions
Structure Constant



$$\sim \frac{\alpha_{\text{weak}}}{M_W^2} \quad \text{at low Energies} \quad (E \leq 1\text{GeV})$$

$$\sim \underbrace{G_F}_{\text{Fermi Constant}} \cong 1 \cdot 10^{-5} \frac{1}{\text{GeV}^2}$$

$$\alpha_{\text{weak}} \approx 4\alpha_{\text{QED}}$$

Weak Structure
Constant

At low energies Weak
Interactions are weak
because of the large mass
of W^\pm (similarly for Z^0)

Basic
Framework

Gauge Theories:
Relativistic Quantum Field Theories with
elementary Forces following from Gauge Symmetries

Quantum
Electrodynamics
(QED)

Symmetry: $U(1)_Q$

Quantum
Flavouredynamics
(QFD)

Symmetry:
 $SU(2)_L \otimes U(1)_Y$

Quantum
Chromodynamics
(QCD)

Symmetry: $SU(3)_C$

Theory of
electromagnetic
interactions

Mediated by
Photon (γ)
($m_\gamma = 0$)

Unified theory of
weak and electromagnetic
interactions

Mediated by weak gauge bosons
Photon (γ), W^\pm Z^0
($m_\gamma = 0$) (80 GeV) (91 GeV)

Theory of
strong interactions
(also basic for
Nuclear Physics)

Mediated by
8 Gluons
($m_G = 0$)

More Messages from the Attouniverse

1.

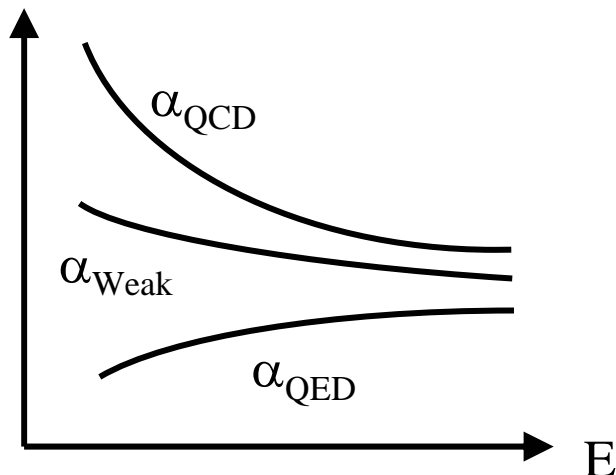
At very high energies ($E \approx M_W, M_Z, m_t$)
supressions $1/M_W^2, 1/M_Z^2$ not important



Weak Interactions as strong as electromagnetic ones.

2.

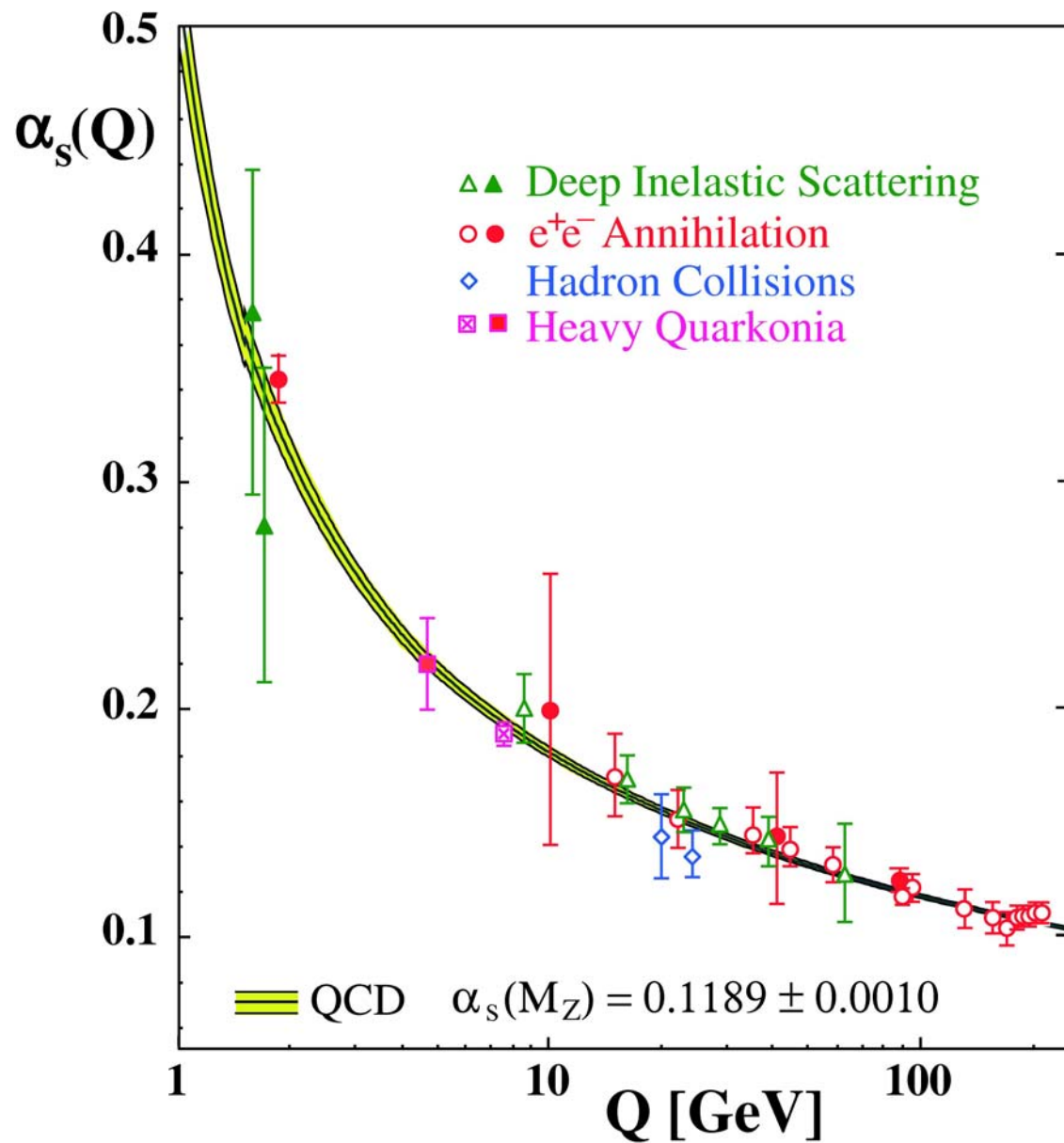
At super high energies ($E \approx 10^{16}$ GeV)



All interactions could
have the same strength
(Grand Unification of Forces)

In a Quantum Field Theory
structure constants are
energy dependent

Asymptotic Freedom in QCD

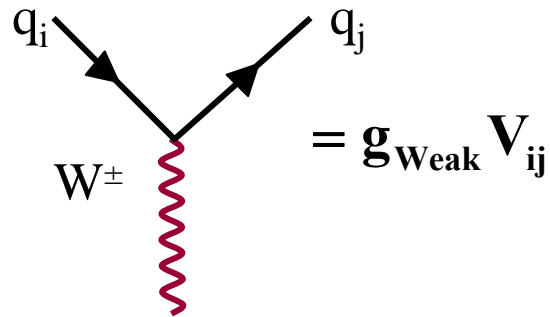


Bethke
hep-ex/0606035

Gross
Politzer
Wilczek (1973)

Nobel Prize
2004

Hierarchical Structure of Quark Flavour-Changing Interactions



$$\alpha_{\text{Weak}} = \frac{g_{\text{Weak}}^2}{4\pi}$$

Cabibbo-Kobayashi-Maskawa matrix

V_{ud}	V_{us}	V_{ub}
V_{cd}	V_{cs}	V_{cb}
V_{td}	V_{ts}	V_{tb}

**Nobel
Prize
2008**

(would be a unit matrix for $m_i=0$)

u	c	t
d	s	b
①	②	③

≈ 1	≈ 0.2	$0.004e^{-i\gamma}$
≈ -0.2	≈ 1	0.04
$0.008e^{-i\beta}$	-0.04	≈ 1

Complex Phases (β, γ)
responsible for
violation of CP Symmetry



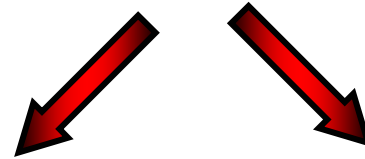
Matter – Antimatter
Symmetry

$$\beta \approx 23^\circ$$

$$\gamma \approx 70^\circ$$



(Nobel Prize 2008)



**Dirac Medal
(2010)**



**N. Cabibbo
(1935-2010)**



M. Kobayashi



T. Maskawa

Crucial Question

**What is the Origin of
Particle Masses and the Reason
for their Hierarchy and
Hierarchy of their
Flavour-Changing Interactions ?**

**Explored in the Research Area C of our
Universe Cluster**

**But the $SU(2)_L \otimes U(1)_Y$ Symmetry
of the Standard Model**

implies that

**W^\pm , Z^0 , quarks and leptons
all should have masses 0 !**

Total disaster !

Higgs Mechanism

Early 1960' → Now

1.

Add a system of scalar particles H that causes spontaneous breakdown

$$SU(2)_L \otimes U(1)_Y \rightarrow U(1)_{QED}$$

Breakdown
only in vacuum!
Not in \mathcal{L}_{SM}

2.

Interactions of H with W^\pm, Z^0
give $M_{Z^0} \neq 0, M_{W^\pm} \neq 0$ while $m_\gamma = 0$


{ Many properties
of TH remain }

3.

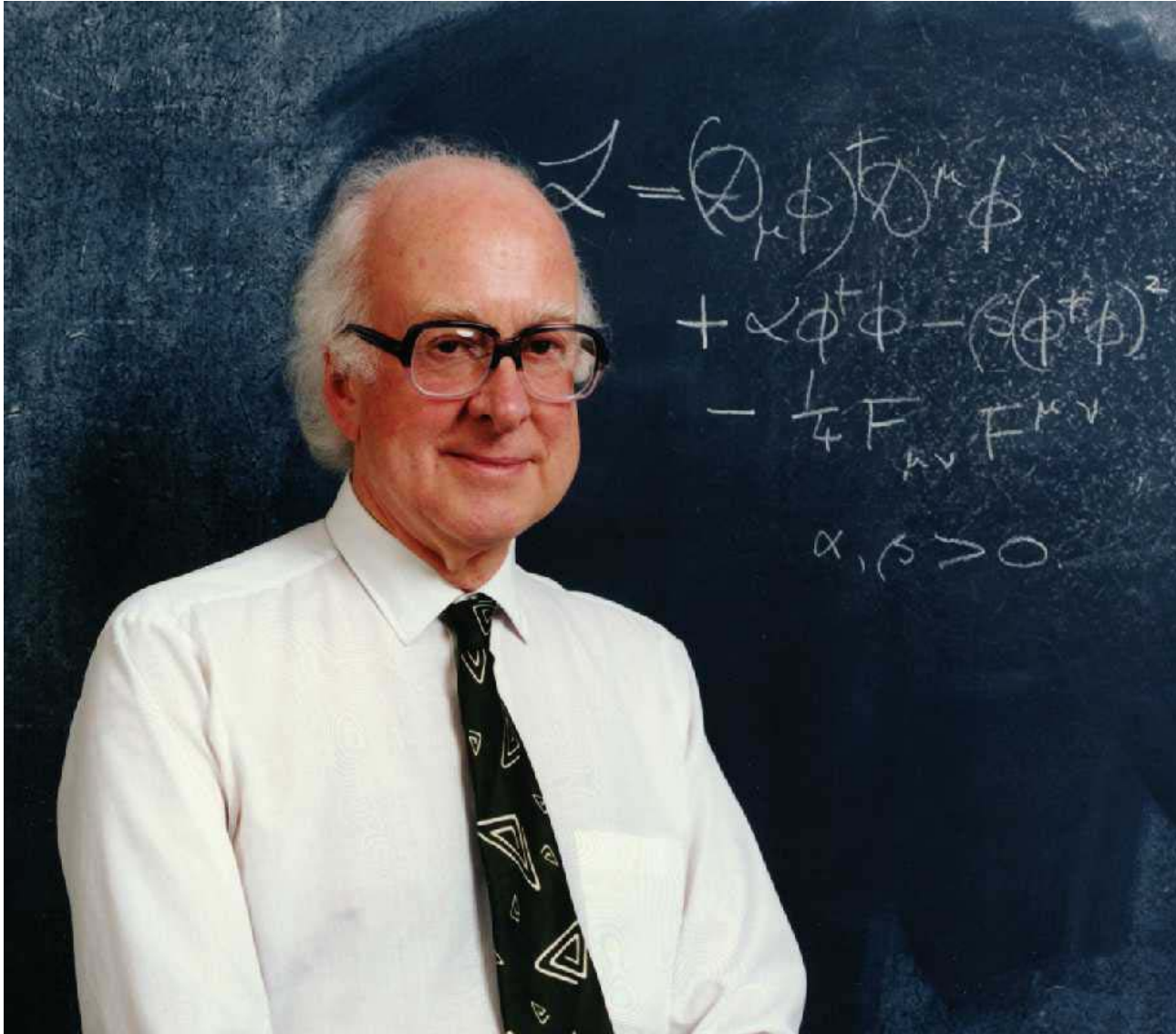
Interactions of H with quarks and leptons



$$\begin{array}{ll} m_q \neq 0 & V_{CKM} \neq 1 \\ m_l \neq 0 & V_{PMNS} \neq 1 \end{array}$$

Parametrization
of quark and lepton
mass and of
flavour-changing
interactions possible

The only Higgs found : Peter Higgs



his mass
(LEP)

$$m_H > 114 \text{ GeV}$$

but in SM
at most

$$190 \text{ GeV}$$

Standard Model of Strong and Electroweak Interactions

**Low Energy Effective Quantum Field Theory
based on (< 200 GeV)**

$$\text{SU}(3)_C \otimes \text{SU}(2)_L \otimes \text{U}(1)_Y \xrightarrow[\text{broken}]{\text{spontaneously}} \text{SU}(3)_C \otimes \text{U}(1)_{\text{QED}}$$

**which describes low energy phenomena in terms
of 28 Parameters that have to be determined from
experiment.**

**The agreement of the Standard Model with the existing
experimental data is very impressive**

But there are Questions !

Where is the Higgs ?

How could we reduce the number of free parameters ?

What is the origin of different generations ?
(Why only 3 ?)

Are there only
3 + 1 Dimensions ?

Matter-Antimatter
Asymmetry ?

How can Higgs mass be protected from becoming $M_{\text{Planck}} \approx 10^{19} \text{ GeV}$?

What is the origin of mass spectrum of Quarks and Leptons ?

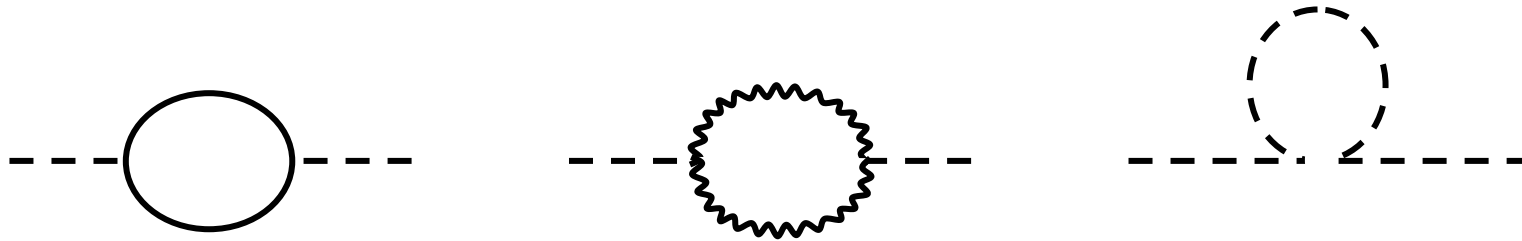
How could one unify all forces including Gravity ?

Dark Matter ?

Hierarchy (Naturalness) Problem

(Quadratic divergences in Higgs mass)

$$m_H^2 \sim \Lambda_{\text{cut-off}}^2 \quad \text{through radiative effects}$$



Disaster for $\Lambda_{\text{cut-off}} \gg 1 \text{ TeV}$

$$\Lambda_{\text{cut-off}} \approx \Lambda_{\text{Planck}}$$

Must fine tune parameters to 34 decimal places to keep $m_H \sim \text{few } 100\text{'s GeV}$

**or postulate New Physics at scales 0 (1 TeV) = 10^3 GeV
which would remove these divergences**

$$\sim 10^{-19} \text{ m}$$

We need

New Physics
(new particles and forces)

**in order to answer
all these questions and
solve all existing problems !**

New Animalcula

Complementary Methods to Search for New Physics

**Direct
Searches**

:

Limited by the available Energy

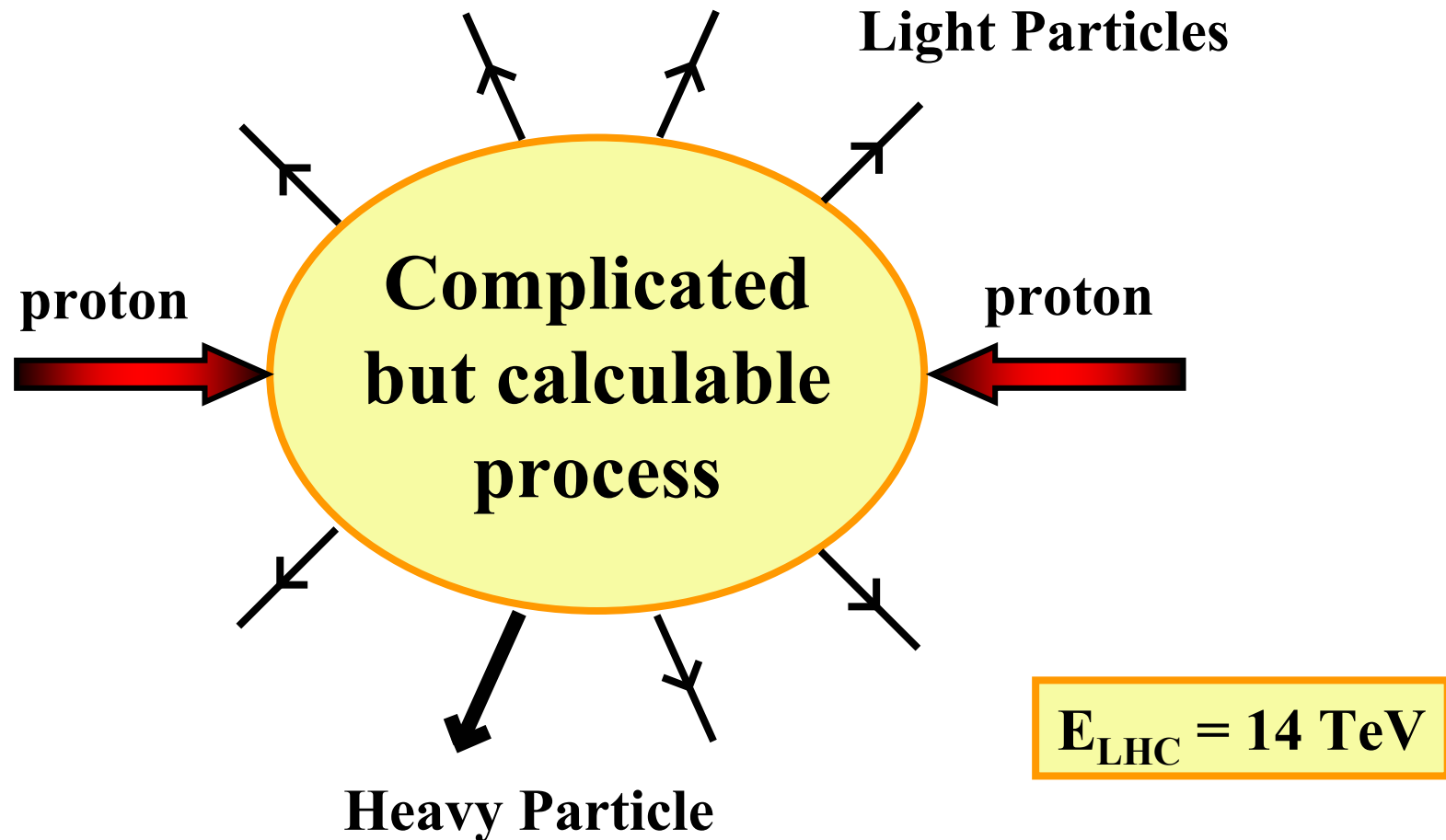
**Indirect
Searches**

:

Quantum Fluctuations

(Limited by precision)

Direct Search: Production of New Heavy Particles

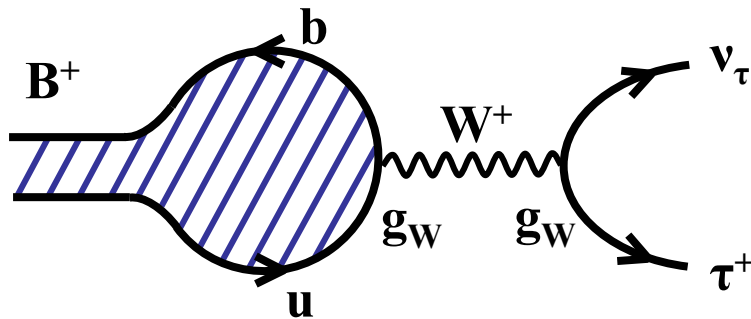


$$c = 1$$

For a Heavy Particle with mass M we need
at least $E_{pp} = M$ if single produced (t)
 $E_{pp} = 2M$ if pair produced ($t\bar{t}$)

Indirect Search: Precision Measurement of Decays of Mesons and Leptons

$$B^+ \rightarrow \tau^+ \nu_\tau$$

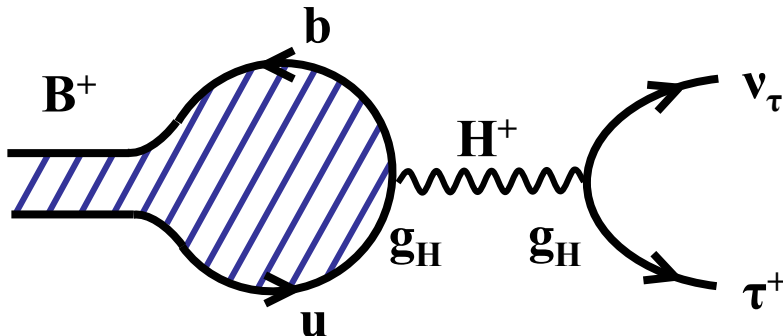


Standard Model

$$\text{Br}(B^+ \rightarrow \tau^+ \nu_\tau)_{\text{SM}} = \left| A \frac{g_W^2}{M_W^2} \right|^2$$

$$m_B \approx 5 \text{ GeV}$$

A, B – parameters of a given theory



Contribution of a new charged Heavy Particle

$$\text{Br}(B^+ \rightarrow \tau^+ \nu_\tau) = \left| A \frac{g_W^2}{M_W^2} + B \frac{g_H^2}{M_H^2} \right|^2$$

$$\Delta = \text{Br}(B^+ \rightarrow \tau^+ \nu_\tau) - \text{Br}(B^+ \rightarrow \tau^+ \nu_\tau)_{\text{SM}} \neq 0$$

Signal of a new particle

Still Large Room for **New Physics**

Standard Model

Exp Upper Bound

$$A_{\text{CP}}(B_s) \approx 0.04$$

$$< 1.0$$

$$\text{Br}(B_s \rightarrow \mu^+ \mu^-) \cong 3 \cdot 10^{-9}$$

$$\sim 6 \cdot 10^{-8}$$

$$\text{Br}(K_L \rightarrow \pi^0 \nu \bar{\nu}) \cong 3 \cdot 10^{-11}$$

$$\sim 6 \cdot 10^{-8}$$

$$\text{Br}(K_L \rightarrow \mu e) \cong 10^{-40}$$

$$\sim 10^{-12}$$

$$\text{Br}(\mu \rightarrow e \gamma) \approx 10^{-54}$$

$$\sim 10^{-11}$$

$$d_n \text{ electric dipole moment of the neutron} \approx 10^{-32} \text{ ecm}$$

$$\sim 10^{-26} \text{ ecm}$$

2nd Movement

Expectations for New Animalcula

Elementary Scalar Particles (Spin = 0)

(Perturbative Framework for Electroweak Symmetry Breaking)

1.

h^0 = Higgs of the Standard Model (Charge = 0)

2.

But many models predict more Higgs particles (Neutral) h^0, H^0, A^0 and H^\pm (charged)

3.

Super-Partners of Quarks and Leptons in Supersymmetry (+2/3, -1/3, -1)

But

**Higgs could be a bound state of
new very heavy fermions: $H = (\bar{F}F)$
bound by a **New strong force:** (Technicolour)**

**Dynamical Breakdown of Electroweak
Symmetry $SU(2)_L \otimes U(1)_Y \rightarrow U(1)_{QED}$**

**Similar to the BCS Theory
of Superconductivity**

4th Generation of Quarks and Leptons

$$\text{Heavy Quarks} : \begin{pmatrix} t' \\ b' \end{pmatrix} \begin{matrix} +2/3 \\ -1/3 \end{matrix}$$

$$\text{Heavy Leptons} : \begin{pmatrix} N \\ E \end{pmatrix} \begin{matrix} 0 \\ -1 \end{matrix}$$

$$300 \text{ GeV} \leq m_t, m_b, \leq 600 \text{ GeV}$$

$$100 \text{ GeV} \leq m_N, m_E, \leq 600 \text{ GeV}$$

If they exist, LHC will find them !

**Most
important
implication**

**New source of CP violation
in 4 x 4 CKM-like Matrix
required for matter-antimatter
asymmetry observed in the universe**

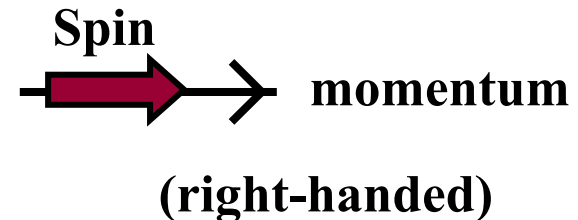
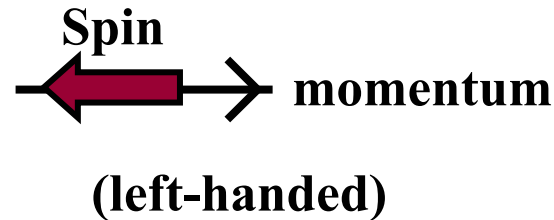
**Standard
Model
cannot
explain this
asymmetry!**

Two Peculiarities of Weak Interactions within the Standard Model

1.

Breakdown of Parity Symmetry (conserved in QED and QCD)

Quark or Lepton



W^\pm interact only with left-handed quarks and leptons

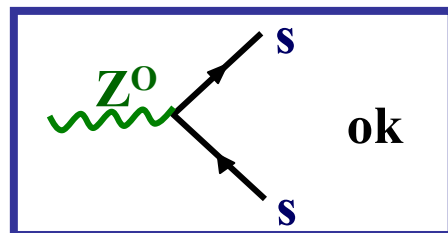
2.

No elementary Flavour-violating Couplings of Neutral Gauge Bosons

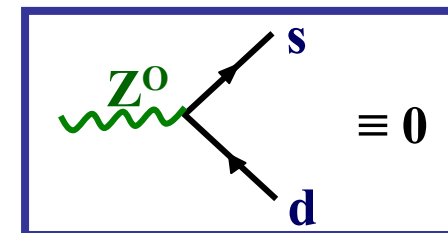
(Z^0 , Photon, Gluons)

Glashow-Maiani-Iliopoulos Mechanism

Example



but



New Electroweak Forces

Most interesting :

Left-right symmetric models
 $SU(2)_L \otimes SU(2)_R \otimes U(1)$

New Heavy Gauge Bosons

W_R^\pm, Z' masses \sim **1 TeV**
 Interacting with
 right-handed fermions

**Order of
 magnitude
 heavier than**

W^\pm, Z

Implication :

**Parity Conservation at
 Ultra short distances $\sim 10^{-20}m$!**

**Spontaneous
 Breakdown
 of Parity :**

$SU(2)_L \otimes SU(2)_R \otimes U(1) \rightarrow SU(2)_L \otimes U(1) \rightarrow U(1)$

**Parity Violation in weak interactions
 at distances $> 10^{-19}m$**

3rd Movement

First Messages from New Animalcula

Departures from Standard Model Expectations

(Most spectacular)

1. ν 's have masses! $(m_\nu)_{\text{SM}} = 0$

2. Anomalous Magnetic Moment of the muon $((g-2)_\mu)$

disagrees by 3σ with the SM

3. CP violation in $B_s \equiv (\bar{b}s)$ decays much larger than predicted (3σ)

hopes for the matter-antimatter asymmetry

DNA Tests of Flavour Models

O_i : *Observables*

M_i : *Models beyond SM*

	M_1	M_2	M_3	M_4	M_5
O_1	★★★	★	★	★	★★
O_2	★	★★	★★★	★★	★
O_3	★★	★★★	★★	★	★
O_4	★★★	★★	★	★★★	★★
O_5	★	★★★	★	★★	★★★



Very large New Physics effect



Moderate New Physics effect



Very small New Physics effect

Models investigated by T31-Teams

(Last decade)

SM

MFV

MSSM+MFV

Z'-Models

**General
MSSM**

**Universal
Extra
Dimensions**

**RS with
custodial
protection**

**Right-
Handed
Currents**

**Littlest
Higgs**

**Littlest
Higgs with
T-Parity**

**SUSY+Flavour
Abelian
Symmetry
(Agashe+Carone)**

**2 Higgs
Doublet
Models**

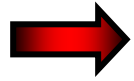
**SUSY with
SU(3) Flavour
(Ross et al)
(RVV2)**

**SUSY with
SU(2) Flavour
(LH-currents)**

**Flavour Blind
MSSM**

4G

"DNA" Patterns of New Physics Models



	AC	RVV2	AKM	δ LL	FBMSSM	LHT	RS	4G
$D^0 - \bar{D}^0$	★★★★	★	★	★	★	★★★★	?	★★
ϵ_K	★	★★★★	★★★★	★	★	★★	★★★★	★★
$S_{\psi\phi}$	★★★★	★★★★	★★★★	★	★	★★★★	★★★★	★★★★
$S_{\phi K_S}$	★★★★	★★	★	★★★★	★★★★	★	?	★★
$A_{CP}(B \rightarrow X_s \gamma)$	★	★	★	★★★★	★★★★	★	?	★
$A_{7,8}(B \rightarrow K^* \mu^+ \mu^-)$	★	★	★	★★★★	★★★★	★★	?	★★
$A_9(B \rightarrow K^* \mu^+ \mu^-)$	★	★	★	★	★	★	?	★★
$B \rightarrow K^{(*)} \nu \bar{\nu}$	★	★	★	★	★	★	★	★
$B_s \rightarrow \mu^+ \mu^-$	★★★★	★★★★	★★★★	★★★★	★★★★	★	★	★★★★
$K^+ \rightarrow \pi^+ \nu \bar{\nu}$	★	★	★	★	★	★★★★	★★★★	★★★★
$K_L \rightarrow \pi^0 \nu \bar{\nu}$	★	★	★	★	★	★★★★	★★★★	★★★★
$\mu \rightarrow e \gamma$	★★★★	★★★★	★★★★	★★★★	★★★★	★★★★	★★★★	★★★★
$\tau \rightarrow \mu \gamma$	★★★★	★★★★	★	★★★★	★★★★	★★★★	★★★★	★★★★
$\mu + N \rightarrow e + N$	★★★★	★★★★	★★★★	★★★★	★★★★	★★★★	★★★★	★★★★
d_n	★★★★	★★★★	★★★★	★★	★★★★	★	★★★★	★
d_e	★★★★	★★★★	★★	★	★★★★	★	★★★★	★
$(g-2)_\mu$	★★★★	★★★★	★★	★★★★	★★★★	★	?	★

2020 Vision

	NEW SM
$D^0 - \bar{D}^0$	★★
ϵ_K	★★
$S_{\psi\phi}$	★★★
$S_{\phi K_S}$	★★
$A_{\text{CP}}(B \rightarrow X_s \gamma)$	★
$A_{7,8}(B \rightarrow K^* \mu^+ \mu^-)$	★★
$A_9(B \rightarrow K^* \mu^+ \mu^-)$	★
$B \rightarrow K^{(*)} \nu \bar{\nu}$	★★★
$B_s \rightarrow \mu^+ \mu^-$	★★★
$K^+ \rightarrow \pi^+ \nu \bar{\nu}$	★★
$K_L \rightarrow \pi^0 \nu \bar{\nu}$	★★★
$\mu \rightarrow e \gamma$	★★★
$\tau \rightarrow \mu \gamma$	★★★
$\mu + N \rightarrow e + N$	★★★
d_n	★★★
d_e	★★★
$(g-2)_\mu$	★★

4th Movement

Final Messages

(TUM, IAS, EC)

Many Thanks to my Collaborators

SUSY



W. Altmannshofer



S. Gori



P. Paradisi



D. Straub

LHT



M. Blanke



B. Duling



S. Recksiegel



C. Tarantino

RS



M. Albrecht



M. Blanke



B. Duling



K. Gemmler



S. Gori



A. Weiler

4 G



B. Duling



T. Heidsieck



C. Promberger



T. Feldmann



S. Recksiegel

2 HDM



M.V. Carlucci



S. Gori



G. Isidori

ϵ_K



D. Guadagnoli

RH Currents



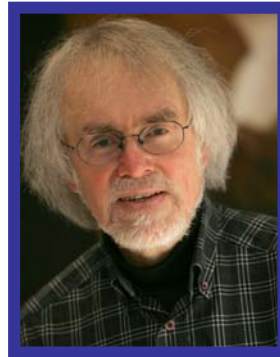
K. Gemmler



G. Isidori

Focus Group : Fundamental Physics (TUM-IAS)

**Senior
Carl von Linde
Fellow
(2008-2011)**



AJB



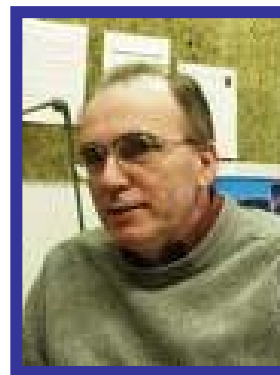
M. Gorbahn

**Junior
Carl von Linde
Fellow
(2008-2011)**

**Senior
Hans Fischer
Fellow
(2009-2012)**



G. Isidori



S. Pokorski

**Senior
Hans Fischer
Fellow
(2009-2012)**



N. Brambilla



G. Buchalla (LMU)



T. Feldmann



A. Ibarra



M. Ratz

Contacts with other Groups at TUM

Low
Energy
QCD

Effective
Theories

Lattice
QCD



W. Weise



N. Brambilla



A. Vairo



N. Kaiser



P. Hägler

Experimental Groups

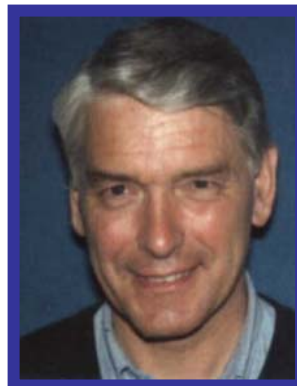
Compass
EDM's



S. Paul



T. Soldner



F. v. Feilitzsch



L. Oberauer

Neutrinos
Dark
Matter

Excellence Cluster (Universe)

MPI-Physics



C. Kiesling



F. Simon



D. Schaile



O. Biebel



J. Schieck



W. Hollik



A. Hoang



S. Bethke



G. Raffelt

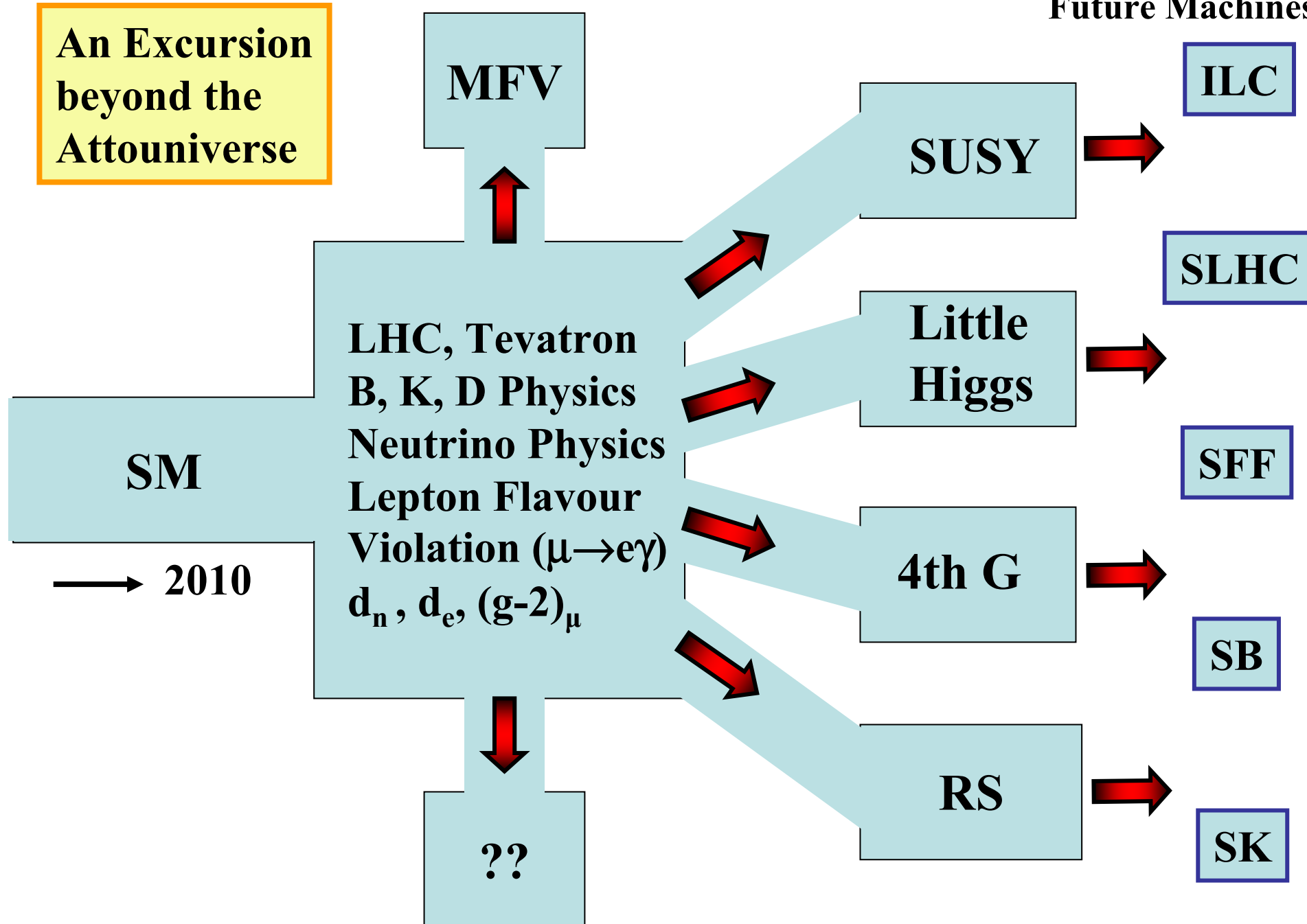


P. Fierlinger



Rondo Vivace !

An Excursion
beyond the
Attouniverse



Bombastic Expectations of Particle Physicists for this Decade

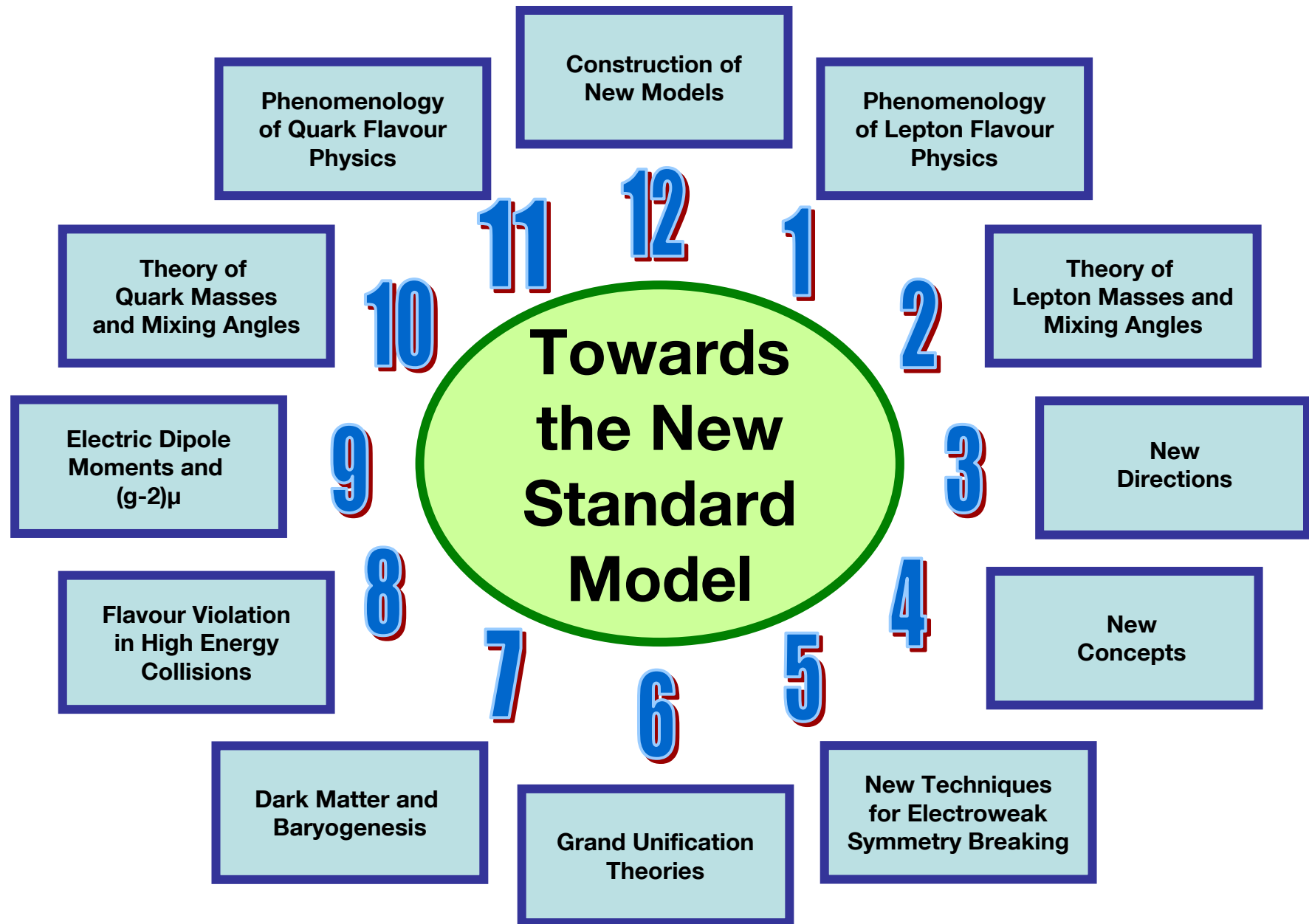
Threshold of a new and exciting era of discovery

Unprecedented accuracy

Long-standing puzzles such as the origin of mass, the matter-antimatter asymmetry of the universe and dark matter will be resolved

These results will have a profound impact on the way we see our universe







Backup