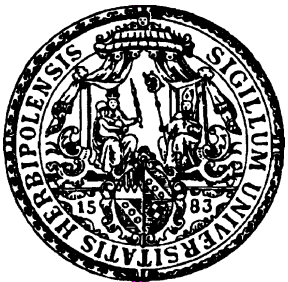
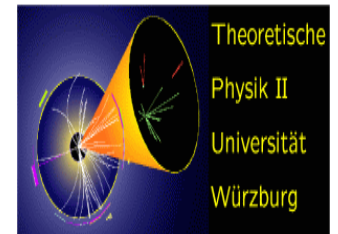


# Testing supersymmetric neutrino mass models at the LHC

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- Neutrinos & lepton flavour violation
- Signals in models with
  - Dirac neutrinos
  - Majorana neutrinos in seesaw models
- How to access the seesaw scale(s)
- Conclusions

Neutrinos: tiny masses

$$\Delta m_{atm}^2 \simeq 2.4 \cdot 10^{-3} \text{ eV}^2$$

$$\Delta m_{sol}^2 \simeq 7.6 \cdot 10^{-5} \text{ eV}^2$$

$${}^3\text{H decay: } m_\nu \lesssim 2 \text{ eV}$$

Neutrinos: large mixings

$$|\tan \theta_{atm}|^2 \simeq 1$$

$$|\tan \theta_{sol}|^2 \simeq 0.4$$

$$|U_{e3}|^2 \lesssim 0.05$$

strong bounds for charged leptons

$$BR(\mu \rightarrow e\gamma) \lesssim 1.2 \cdot 10^{-11}$$

$$BR(\mu^- \rightarrow e^- e^+ e^-) \lesssim 10^{-12}$$

$$BR(\tau \rightarrow e\gamma) \lesssim 1.1 \cdot 10^{-7}$$

$$BR(\tau \rightarrow \mu\gamma) \lesssim 4.5 \cdot 10^{-8}$$

$$BR(\tau \rightarrow lll') \lesssim O(10^{-8}) \quad (l, l' = e, \mu)$$

$$|d_e| \lesssim 10^{-27} \text{ e cm}, \quad |d_\mu| \lesssim 1.5 \cdot 10^{-18} \text{ e cm}, \quad |d_\tau| \lesssim 1.5 \cdot 10^{-16} \text{ e cm}$$

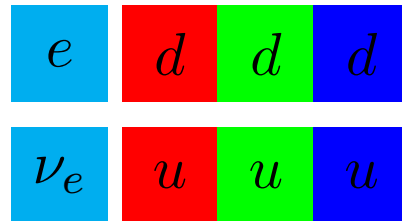
SUSY contributions to anomalous magnetic moments

$$|\Delta a_e| \leq 10^{-12}, \quad 0 \leq \Delta a_\mu \leq 43 \cdot 10^{-10}, \quad |\Delta a_\tau| \leq 0.058$$

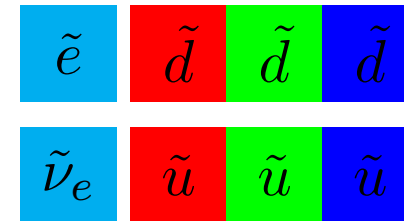
Standard Model

MSSM

matter:



$\Leftrightarrow$



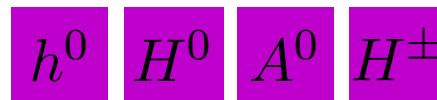
gauge sector:



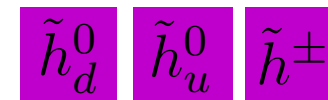
$\Leftrightarrow$



Higgs sector:



$\Leftrightarrow$



assume at the moment conserved  $R$ -Parity:  $(-1)^{(3(B-L)+2s)}$

$(\tilde{\gamma}, \tilde{z}^0, \tilde{h}_d^0, \tilde{h}_u^0) \rightarrow \tilde{\chi}_i^0, (\tilde{w}^\pm, \tilde{h}^\pm) \rightarrow \tilde{\chi}_j^\pm$

analog to leptons or quarks

$$Y_\nu H \bar{\nu}_L \nu_R \rightarrow Y_\nu v \bar{\nu}_L \nu_R = m_\nu \bar{\nu}_L \nu_R$$

requires  $Y_\nu \ll Y_e$

⇒ no impact for present or future collider experiments

Exception:  $\tilde{\nu}_R$  is LSP and thus a candidate for dark matter

⇒ long lived NLSP, e.g.  $\tilde{t}_1 \rightarrow l^+ b \tilde{\nu}_R$

Remark:  $m_{\tilde{\nu}_R}$  hardly runs ⇒ e.g.  $m_{\tilde{\nu}_R} \simeq m_0$  in mSUGRA

$m_{\tilde{\nu}_R} \simeq 0$  in GMSB

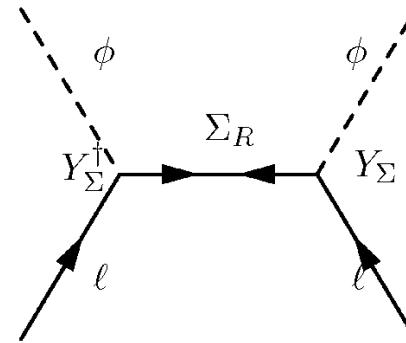
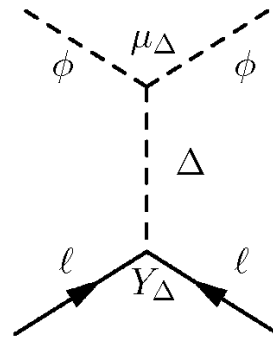
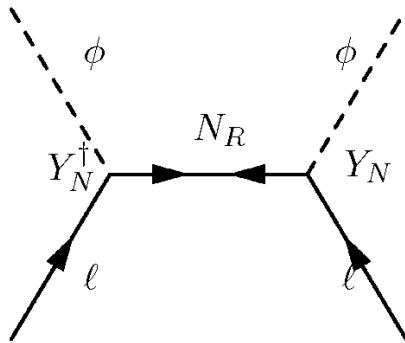
S. Gopalakrishna, A. de Gouvea and W. P., JHEP **0611** (2006) 050

S. K. Gupta, B. Mukhopadhyaya, S. K. Rai, PRD **75** (2007) 075007

D. Choudhury, S. K. Gupta, B. Mukhopadhyaya, PRD **78** (2008) 015023

Neutrino masses due to

$$\frac{f}{\Lambda}(HL)(HL)$$



\* P. Minkowski, Phys. Lett. B **67** (1977) 421; T. Yanagida, KEK-report 79-18 (1979);  
M. Gell-Mann, P. Ramond, R. Slansky, in *Supergravity*, North Holland (1979), p. 315;  
R.N. Mohapatra and G. Senjanovic, *Phys. Rev. Lett.* **44** 912 (1980); M. Magg and C. Wetterich,  
Phys. Lett. B **94** (1980) 61; G. Lazarides, Q. Shafi and C. Wetterich, Nucl. Phys. B **181**  
(1981) 287; J. Schechter and J. W. F. Valle, Phys. Rev. **D25**, 774 (1982);  
R. Foot, H. Lew, X. G. He and G. C. Joshi, Z. Phys. C **44** (1989) 441.

Relevant SU(5) invariant parts of the superpotentials at  $M_{GUT}$

● Type-I

$$W_{\text{RHN}} = \mathbf{Y}_N^{\text{I}} N^c \bar{\mathbf{5}}_M \cdot \mathbf{5}_H + \frac{1}{2} M_R N^c N^c$$

● Type-II

$$W_{15\text{H}} = \frac{1}{\sqrt{2}} \mathbf{Y}_N^{\text{II}} \bar{\mathbf{5}}_M \cdot 15 \cdot \bar{\mathbf{5}}_M + \frac{1}{\sqrt{2}} \lambda_1 \bar{\mathbf{5}}_H \cdot 15 \cdot \bar{\mathbf{5}}_H + \frac{1}{\sqrt{2}} \lambda_2 \mathbf{5}_H \cdot \bar{15} \cdot \mathbf{5}_H \\ + M_{15} 15 \cdot \bar{15}$$

● Type-III

$$W_{24\text{H}} = \mathbf{5}_H 24_M \mathbf{Y}_N^{\text{III}} \bar{\mathbf{5}}_M + \frac{1}{2} 24_M M_{24} 24_M$$

Under  $SU(3) \times SU_L(2) \times U(1)_Y$

- The  $\mathbf{5}$ ,  $\mathbf{10}$  and  $\mathbf{5}_H$  contain

$$\bar{\mathbf{5}}_M = (\widehat{D}^c, \widehat{L}), \quad \mathbf{10} = (\widehat{U}^c, \widehat{E}^c, \widehat{Q}), \quad \mathbf{5}_H = (\widehat{H}^c, \widehat{H}_u), \quad \bar{\mathbf{5}}_H = (\widehat{H}^c, \widehat{H}_d)$$

- The  $\mathbf{15}$  decomposes as

$$\mathbf{15} = \widehat{S}(6, 1, -\frac{2}{3}) + \widehat{T}(1, \mathbf{3}, 1) + \widehat{Z}(3, 2, \frac{1}{6})$$

- The  $\mathbf{24}$  decomposes as

$$\begin{aligned} \mathbf{24}_M = & \widehat{W}_M(1, \mathbf{3}, 0) + \widehat{B}_M(1, 1, 0) + \widehat{X}_M(3, 2, -\frac{5}{6}) \\ & + \widehat{X}_M(\bar{\mathbf{3}}, 2, \frac{5}{6}) + \widehat{G}_M(8, 1, 0) \end{aligned}$$



Postulate very heavy right-handed neutrinos yielding the following superpotential below  $M_{GUT}$ :

$$W_I = W_{MSSM} + W_\nu ,$$

$$W_\nu = \hat{N}^c Y_\nu \hat{L} \cdot \hat{H}_u + \frac{1}{2} \hat{N}^c M_R \hat{N}^c ,$$

Neutrino mass matrix

$$m_\nu = -\frac{v_u^2}{2} Y_\nu^T M_R^{-1} Y_\nu$$

Inverting the seesaw equation gives  $Y_\nu$  a la Casas & Ibarra

$$Y_\nu = \sqrt{2} \frac{i}{v_u} \sqrt{\hat{M}_R} \cdot R \cdot \sqrt{\hat{m}_\nu} \cdot U^\dagger$$

$\hat{m}_\nu, \hat{M}_R$  ... diagonal matrices containing the corresponding eigenvalues

$U$  ..... neutrino mixing matrix

$R$  ..... complex orthogonal matrix.

Below  $M_{GUT}$  the superpotential reads

$$\begin{aligned}
 W_{II} &= W_{MSSM} + \frac{1}{\sqrt{2}} (Y_T \hat{L} \hat{T}_1 \hat{L} + Y_S \hat{D}^c \hat{S}_1 \hat{D}^c) + Y_Z \hat{D}^c \hat{Z}_1 \hat{L} \\
 &+ \frac{1}{\sqrt{2}} (\lambda_1 \hat{H}_d \hat{T}_1 \hat{H}_d + \lambda_2 \hat{H}_u \hat{T}_2 \hat{H}_u) + M_T \hat{T}_1 \hat{T}_2 + M_Z \hat{Z}_1 \hat{Z}_2 + M_S \hat{S}_1 \hat{S}_2
 \end{aligned}$$

fields with index 1 (2) originate from the 15-plet ( $\overline{15}$ -plet).

The effective mass matrix is

$$m_\nu = -\frac{v_u^2}{2} \frac{\lambda_2}{M_T} Y_T.$$

Note that

$$\hat{Y}_T = U^T \cdot Y_T \cdot U,$$

In the  $SU(5)$  broken phase the superpotential becomes

$$\begin{aligned}
 W_{III} = & W_{MSSM} + \hat{H}_u (\hat{W}_M Y_N - \sqrt{\frac{3}{10}} \hat{B}_M Y_B) \hat{L} + \hat{H}_u \hat{X}_M Y_X \hat{D}^c \\
 & + \frac{1}{2} \hat{B}_M M_B \hat{B}_M + \frac{1}{2} \hat{G}_M M_G \hat{G}_M + \frac{1}{2} \hat{W}_M M_W \hat{W}_M + \hat{X}_M M_X \hat{X}_M
 \end{aligned}$$

giving

$$m_\nu = -\frac{v_u^2}{2} \left( \frac{3}{10} Y_B^T M_B^{-1} Y_B + \frac{1}{2} Y_W^T M_W^{-1} Y_W \right) \simeq -v_u^2 \frac{4}{10} Y_W^T M_W^{-1} Y_W$$

last step: valid if  $M_B \simeq M_W$  and  $Y_B \simeq Y_W$

⇒ Casas-Ibarra decomposition for  $Y_W$  as in type-I up to factor 4/5

MSSM:  $(b_1, b_2, b_3) = (33/5, 1, -3)$

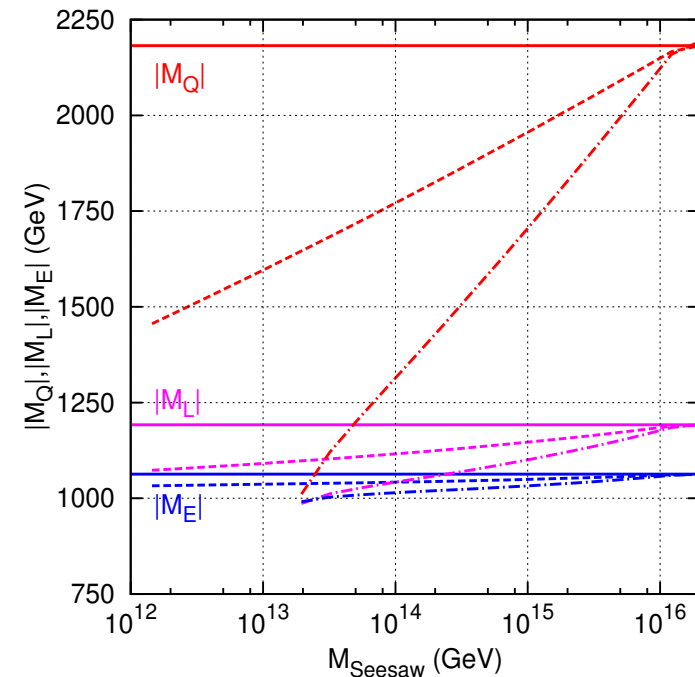
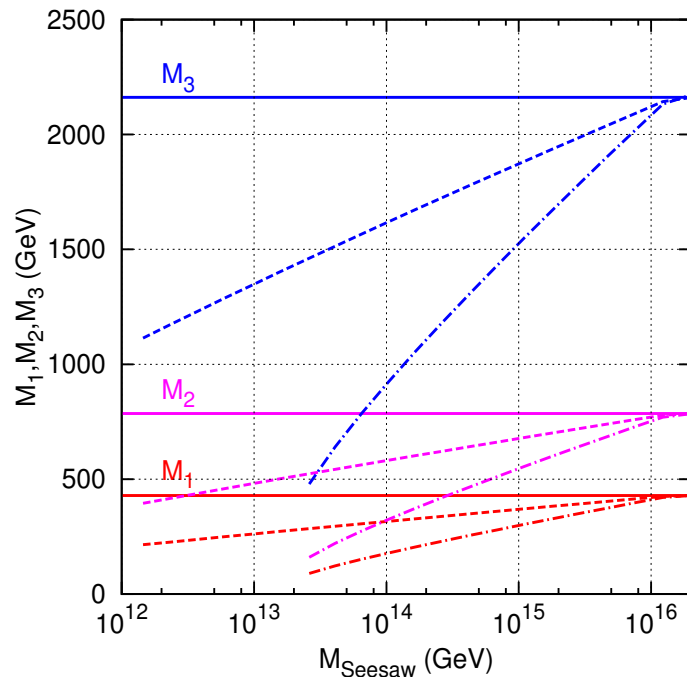
per 15-plet  $\Delta b_i = 7/2 \Rightarrow$  type II model  $\Delta b_i = 7$

per 24-plet  $\Delta b_i = 5 \Rightarrow$  type III model  $\Delta b_i = 15$

MSSM:  $(b_1, b_2, b_3) = (33/5, 1, -3)$

per 15-plet  $\Delta b_i = 7/2 \Rightarrow$  type II model  $\Delta b_i = 7$

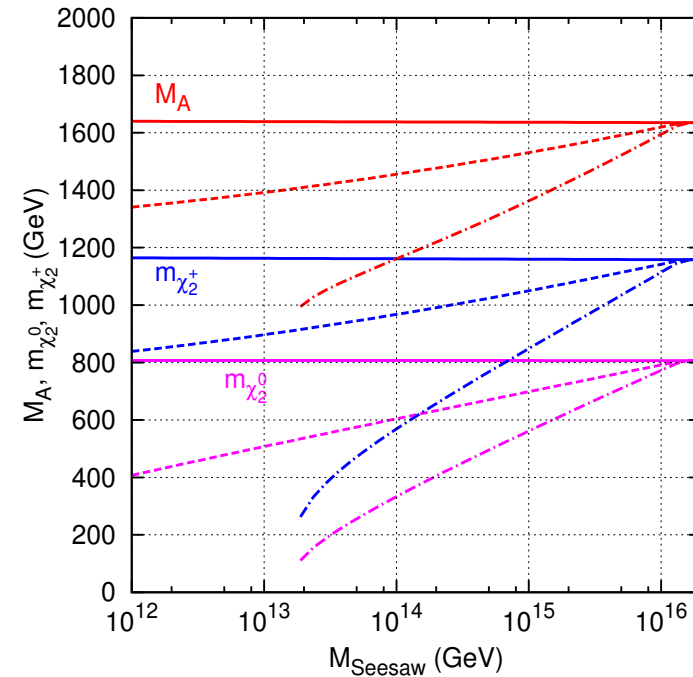
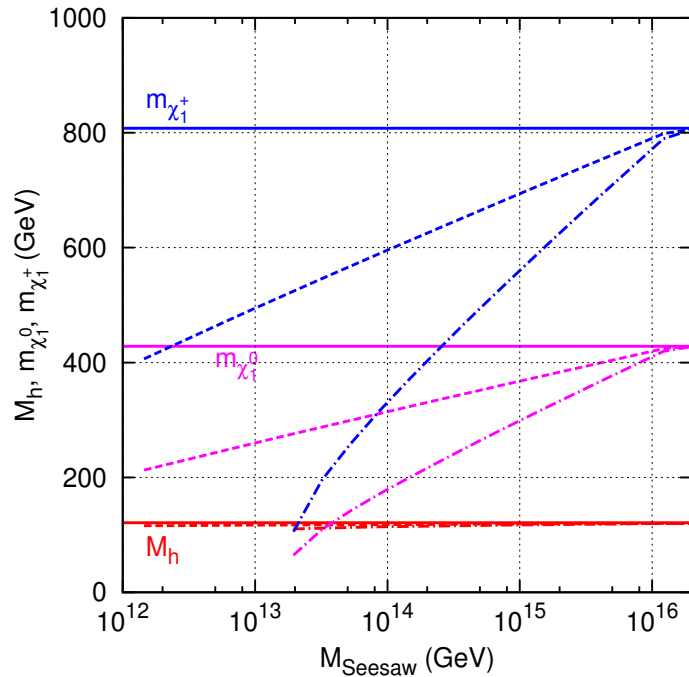
per 24-plet  $\Delta b_i = 5 \Rightarrow$  type III model  $\Delta b_i = 15$



$Q = 1 \text{ TeV}, m_0 = M_{1/2} = 1 \text{ TeV}, A_0 = 0, \tan \beta = 10$  and  $\mu > 0$

Full lines ... seesaw type I, dashed lines ... type II, dash-dotted lines ... type III  
degenerate spectrum of the seesaw particles

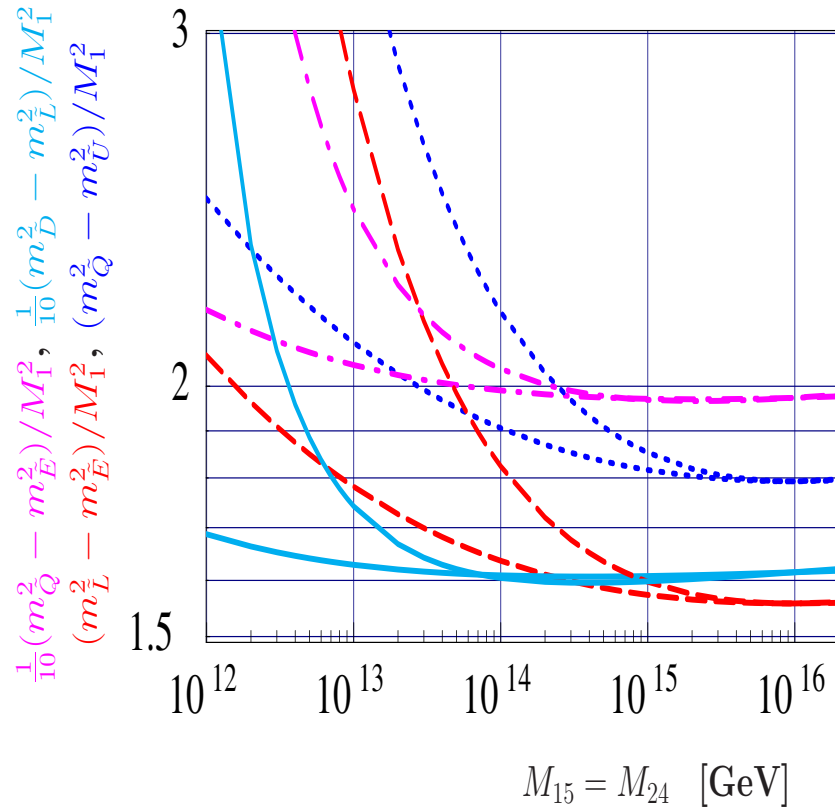
J. N. Esteves, M. Hirsch, W.P., J. C. Romao, F. Staub, Phys. Rev. D83 (2011) 013003



$$m_0 = M_{1/2} = 1 \text{ TeV}, A_0 = 0, \tan \beta = 10 \text{ and } \mu > 0$$

Full lines ... seesaw type I, dashed lines ... type II, dash-dotted lines ... type III  
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J. N. Esteves, M. Hirsch, W.P., J. C. Romao, F. Staub, Phys. Rev. D83 (2011) 013003



Seesaw I ( $\simeq$  MSSM)

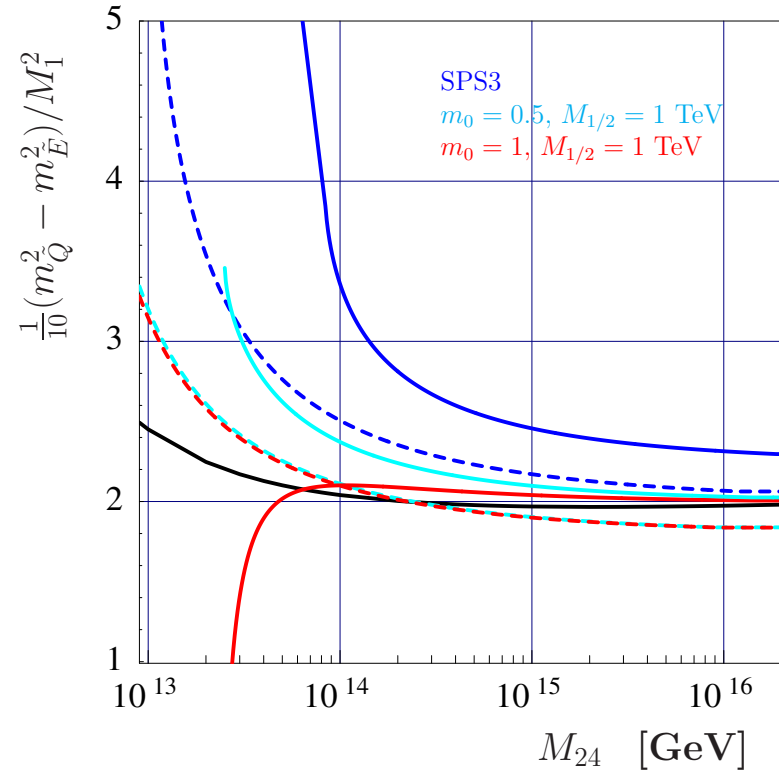
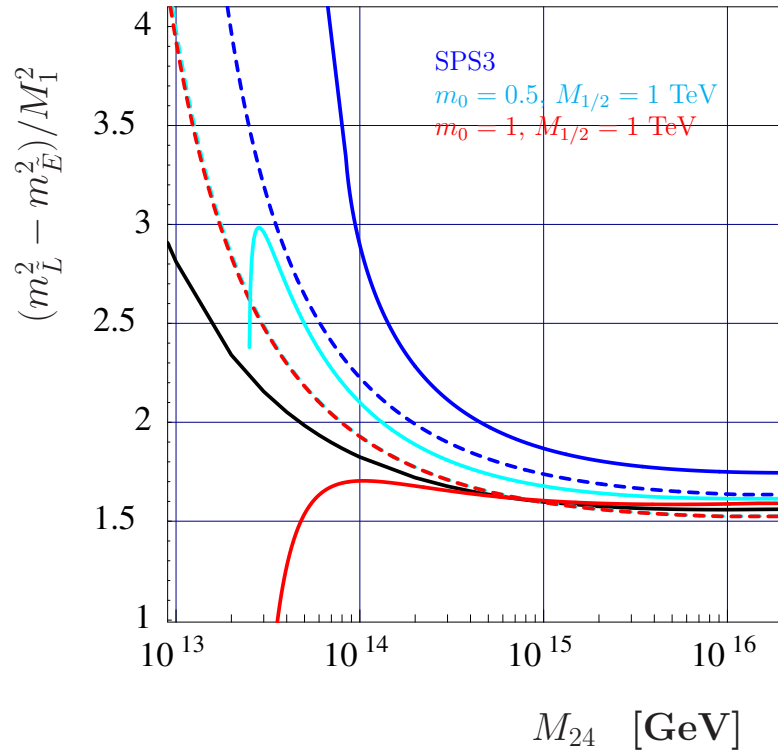
$$\frac{m_Q^2 - m_E^2}{M_1^2} \simeq 20, \quad \frac{m_D^2 - m_L^2}{M_1^2} \simeq 18$$

$$\frac{m_L^2 - m_E^2}{M_1^2} \simeq 1.6, \quad \frac{m_Q^2 - m_U^2}{M_1^2} \simeq 1.55$$

(solution of 1-loop RGEs)

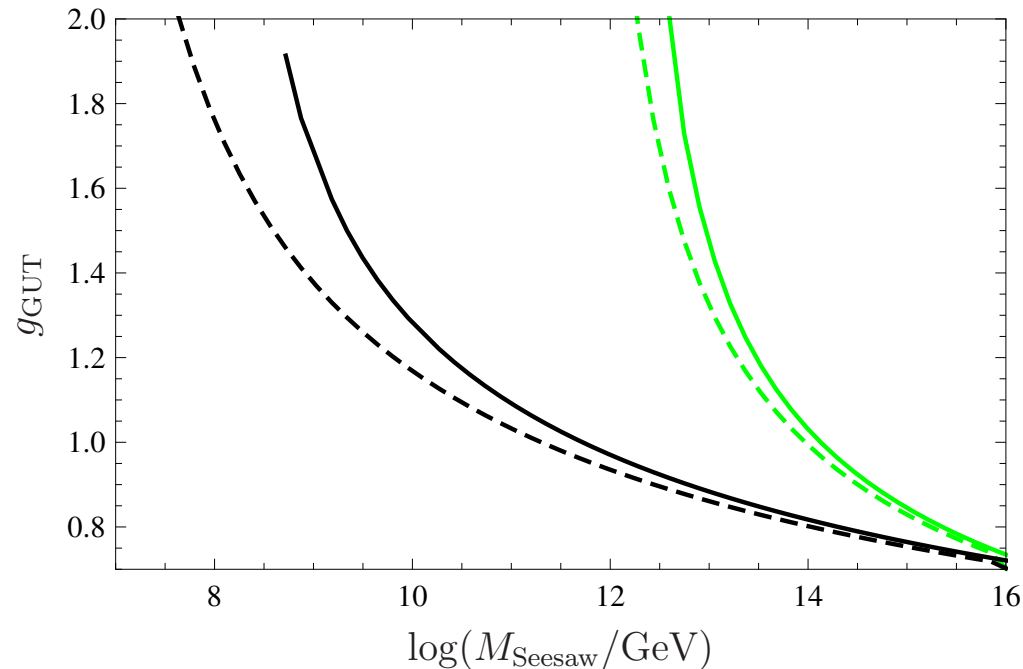
M. Hirsch, S. Kaneko, W. P., Phys. Rev. D **78** (2008) 093004

J. Esteves, M. Hirsch, J. Romão, W. P., F. Staub, Phys. Rev. D **83** (2011) 013003



- blue lines ... SPS3
- light blue lines ...  $m_0 = 500 \text{ GeV}$  and  $M_{1/2} = 1 \text{ TeV}$
- red lines ...  $m_0 = M_{1/2} = 1 \text{ TeV}$
- black line ... analytical approximation
- full (dashed) lines ... 2-loop (1-loop) results





$m_0 = M_{1/2} = 1 \text{ TeV}$ ,  $A_0 = 0$ ,  $\tan \beta = 10$  and  $\mu > 0$

$M_{GUT} = 2 \times 10^{16} \text{ GeV}$

black lines ... seesaw type-II

green lines ... seesaw type-III with three **24**-plets with degenerate mass spectrum

full (dashed) lines ... 2-loop (1-loop) results

one-step integration of the RGEs assuming mSUGRA boundary

$$\Delta M_{L,ij}^2 \simeq - \frac{a_k}{8\pi^2} (3m_0^2 + A_0^2) \left( Y_N^{k,\dagger} L Y_N^k \right)_{ij}$$

$$\Delta A_{l,ij} \simeq - a_k \frac{3}{16\pi^2} A_0 \left( Y_e Y_N^{k,\dagger} L Y_N^k \right)_{ij}$$

$$\Delta M_{E,ij}^2 \simeq 0$$

$$L_{ij} = \ln(M_{GUT}/M_i) \delta_{ij}$$

for  $i \neq j$  with  $Y_e$  diagonal

$$a_I = 1, \quad a_{II} = 6 \quad \text{and} \quad a_{III} = \frac{9}{5}$$

$(\Delta M_{\tilde{L}}^2)_{ij}$  and  $(\Delta A_l)_{ij}$  induce

$$\begin{aligned} l_j &\rightarrow l_i \gamma, l_i l_k^+ l_r^- \\ \tilde{l}_j &\rightarrow l_i \tilde{\chi}_s^0 \\ \tilde{\chi}_s^0 &\rightarrow l_i \tilde{l}_k \end{aligned}$$

Neglecting  $L$ - $R$  mixing:

$$\begin{aligned} Br(l_i \rightarrow l_j \gamma) &\propto \alpha^3 m_{l_i}^5 \frac{|(\delta m_L^2)_{ij}|^2}{\tilde{m}^8} \tan^2 \beta \\ \frac{Br(\tilde{\tau}_2 \rightarrow e + \chi_1^0)}{Br(\tilde{\tau}_2 \rightarrow \mu + \chi_1^0)} &\simeq \left( \frac{(\Delta M_L^2)_{13}}{(\Delta M_L^2)_{23}} \right)^2 \end{aligned}$$

Moreover, in most of the parameter space

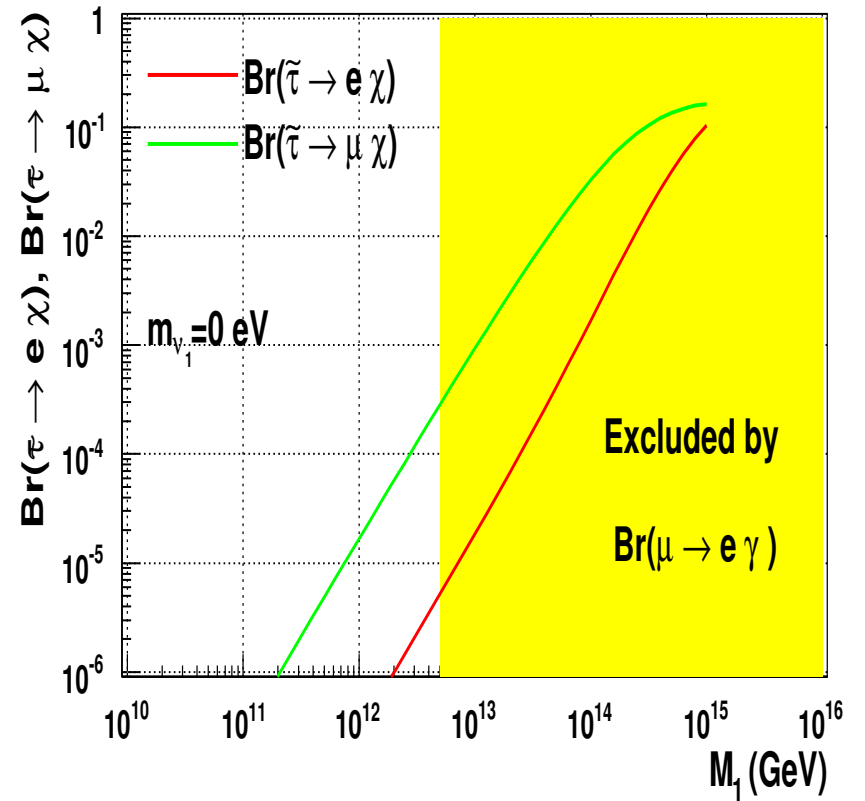
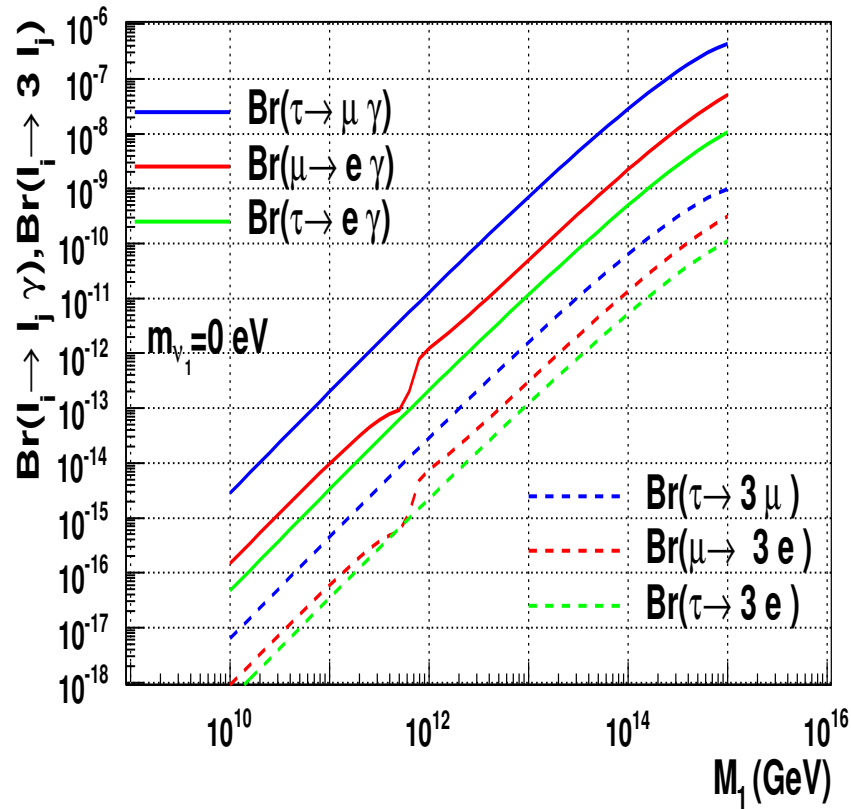
$$\frac{Br(l_i \rightarrow 3l_j)}{Br(l_i \rightarrow l_j + \gamma)} \simeq \frac{\alpha}{3\pi} \left( \log\left(\frac{m_{l_i}^2}{m_{l_j}^2}\right) - \frac{11}{4} \right)$$

take all parameters real

$$U = U_{\text{TBM}} = \begin{pmatrix} \sqrt{\frac{2}{3}} & \frac{1}{\sqrt{3}} & 0 \\ -\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \\ \frac{1}{\sqrt{6}} & -\frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \end{pmatrix}$$

$$R = \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

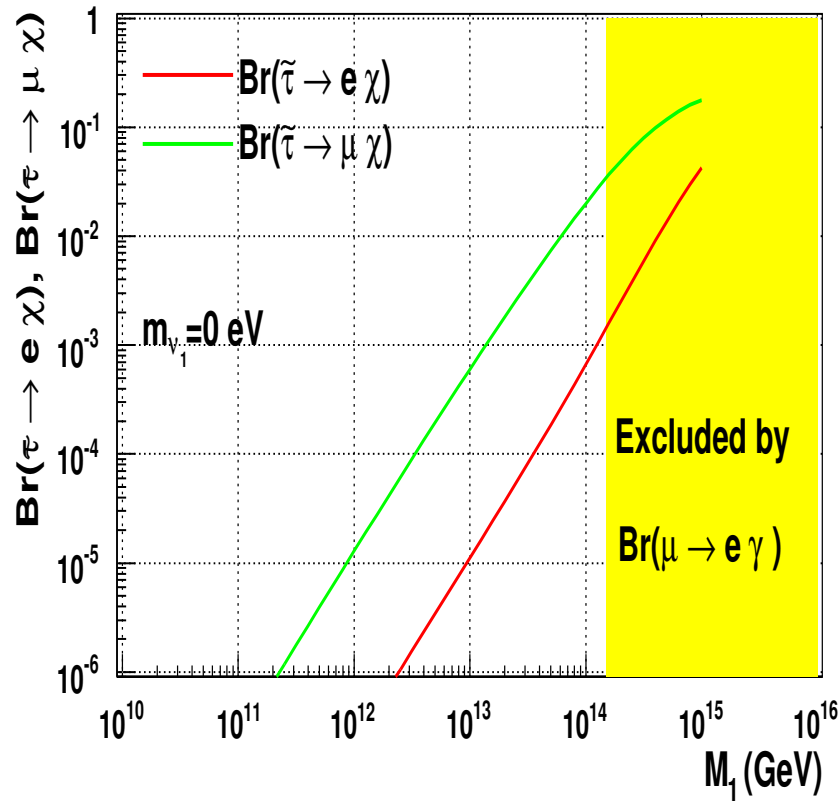
Use 2-loop RGEs and 1-loop corrections including flavour effects



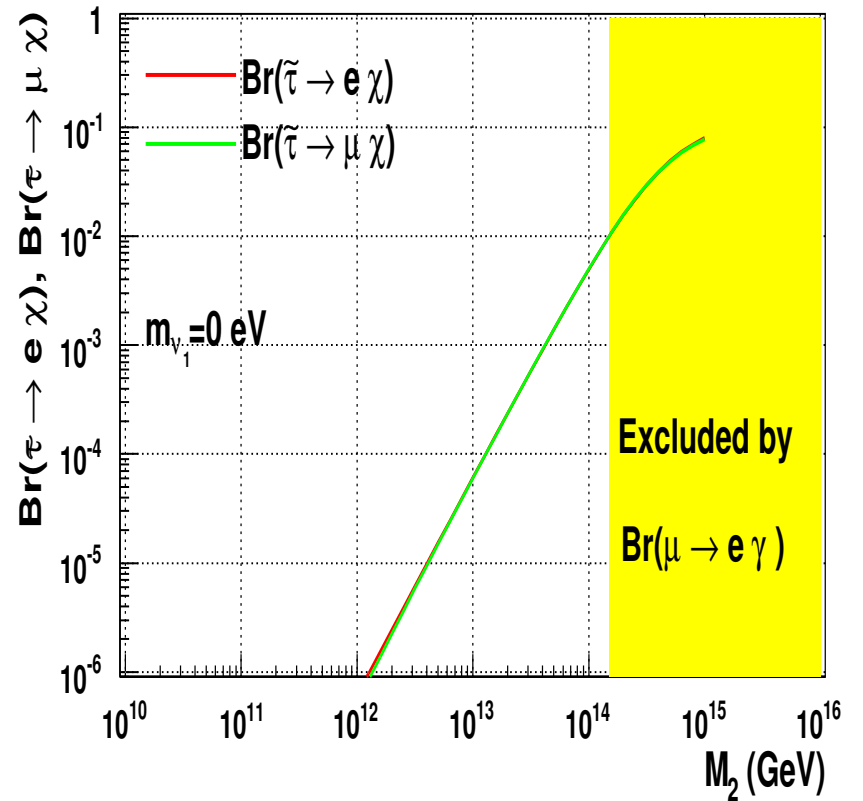
degenerate  $\nu_R$

SPS1a' ( $M_0 = 70$  GeV,  $M_{1/2} = 250$  GeV,  $A_0 = -300$  GeV,  $\tan \beta = 10$ ,  $\mu > 0$ )

M. Hirsch et al. Phys. Rev. D 78 (2008) 013006



degenerate  $\nu_R$



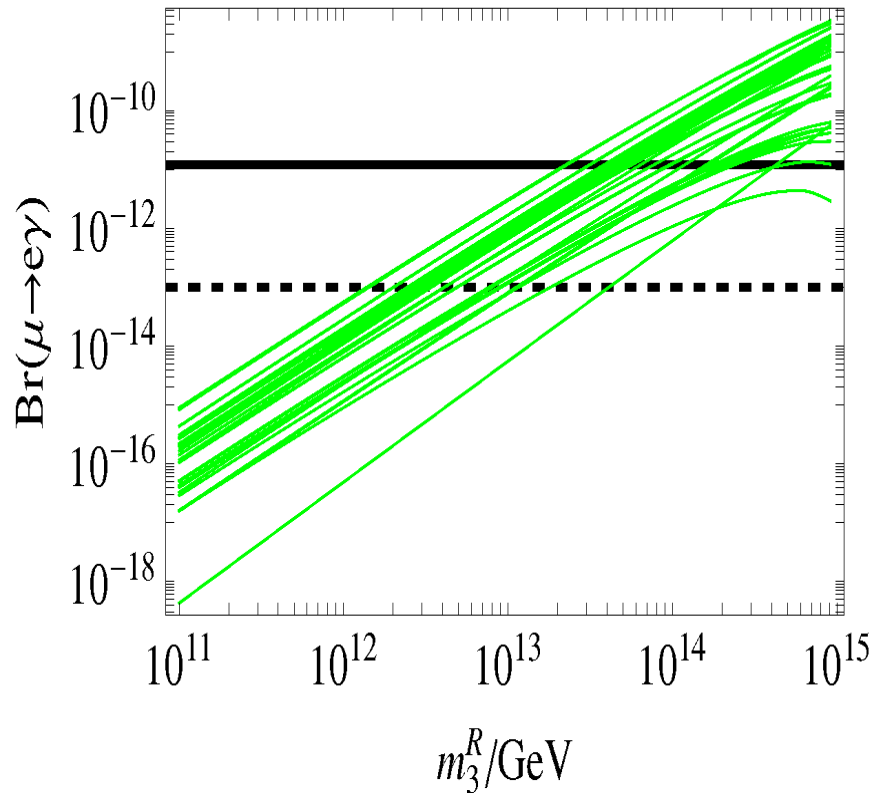
hierarchical  $\nu_R$

$(M_1 = M_3 = 10^{10} \text{ GeV})$

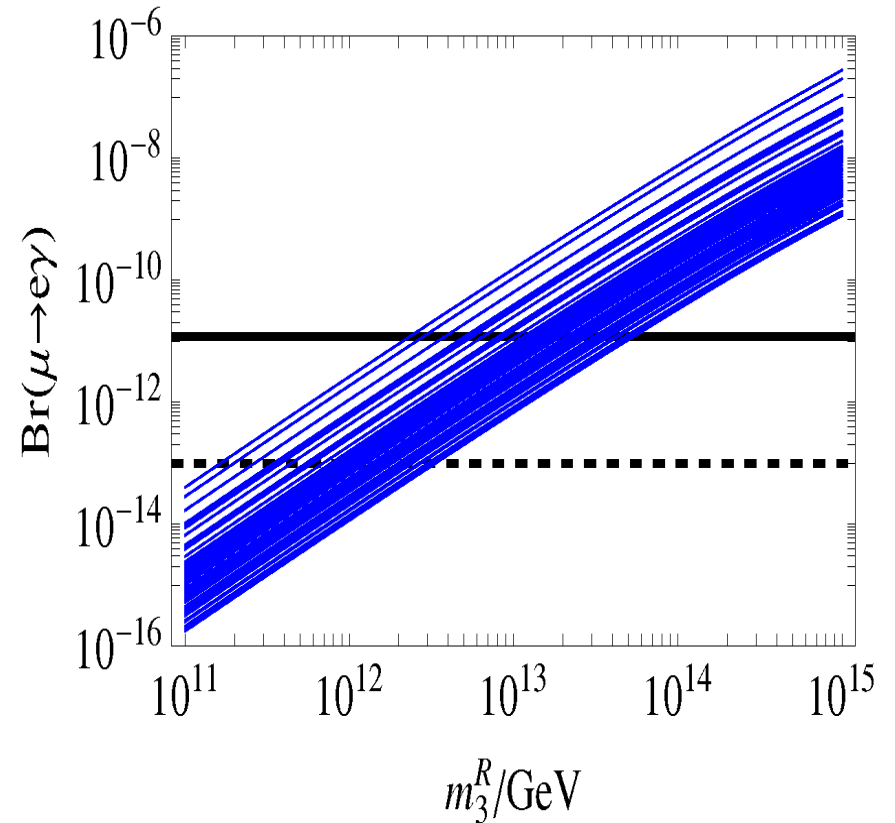
SPS3 ( $M_0 = 90 \text{ GeV}$ ,  $M_{1/2} = 400 \text{ GeV}$ ,  $A_0 = 0 \text{ GeV}$ ,  $\tan \beta = 10$ ,  $\mu > 0$ )

M. Hirsch et al. Phys. Rev. D 78 (2008) 013006

Texture models, hierarchical  $\nu_R$   
real textures

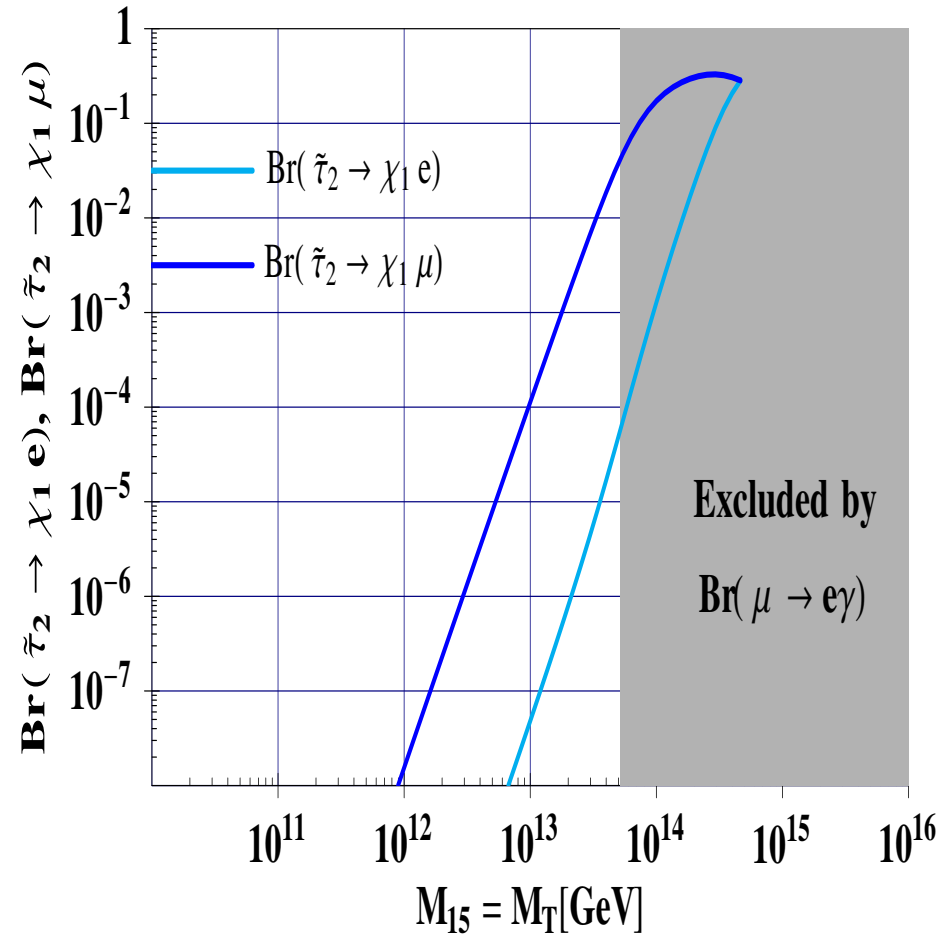
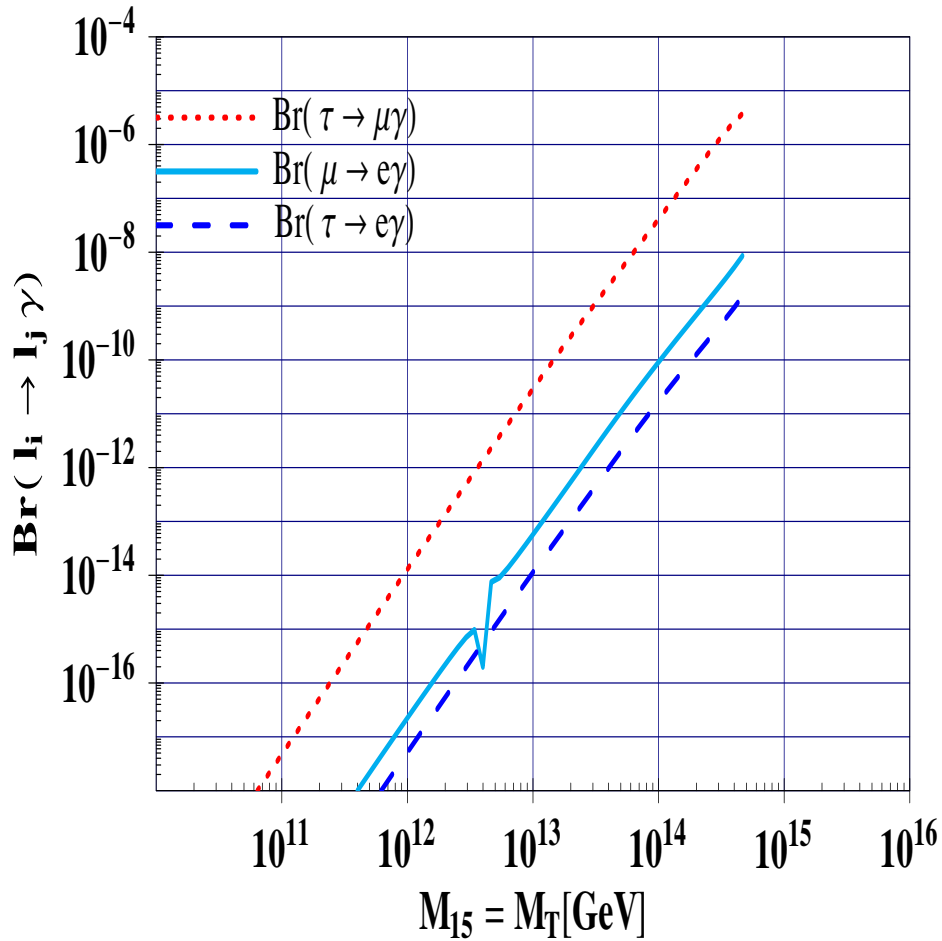


"complexification" of one texture



SPS1a' ( $M_0 = 70$  GeV,  $M_{1/2} = 250$  GeV,  $A_0 = -300$  GeV,  $\tan \beta = 10$ ,  $\mu > 0$ )

F. Deppisch, F. Plentinger, G. Seidl, JHEP 1101 (2011) 004

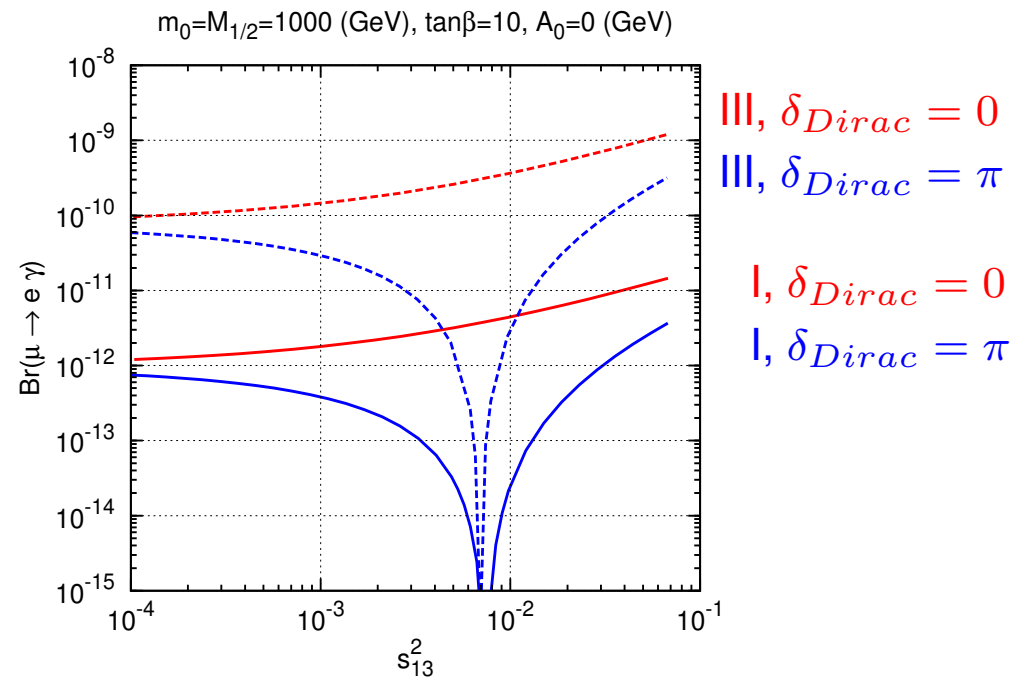
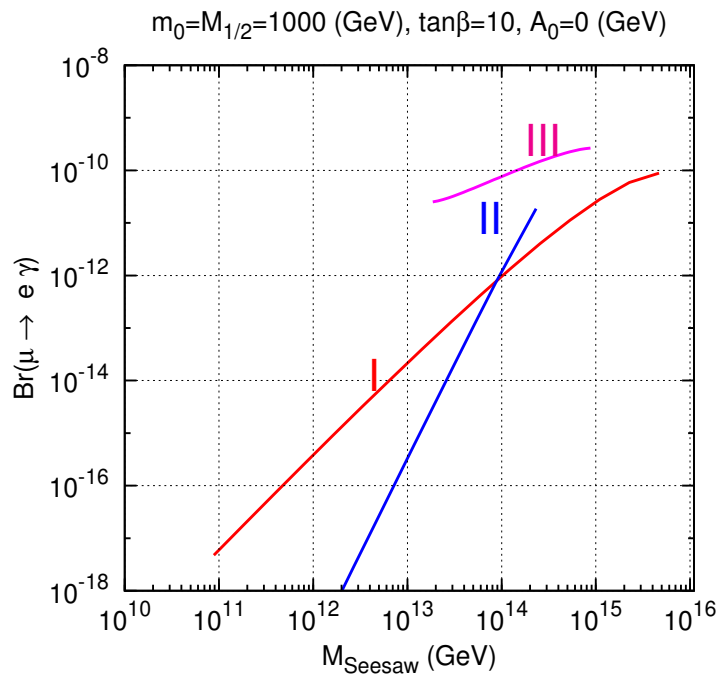


$$\lambda_1 = \lambda_2 = 0.5$$

$$\text{SPS3 } (M_0 = 90 \text{ GeV}, M_{1/2} = 400 \text{ GeV}, A_0 = 0 \text{ GeV}, \tan \beta = 10, \mu > 0)$$

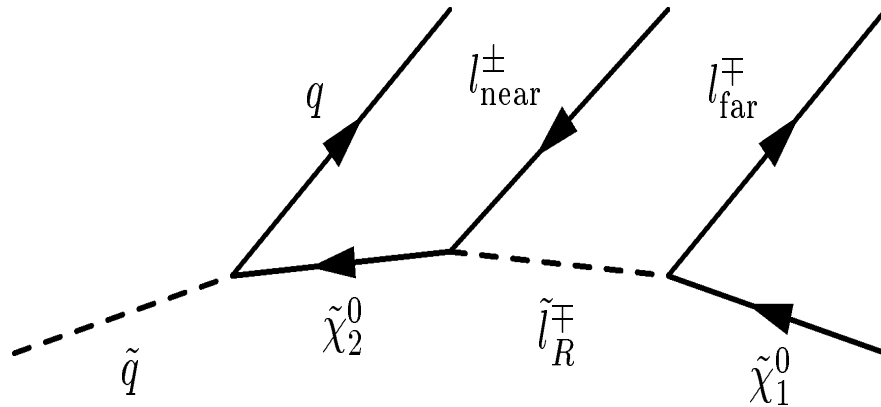
M. Hirsch, S. Kaneko, W. P., Phys. Rev. D 78 (2008) 093004.



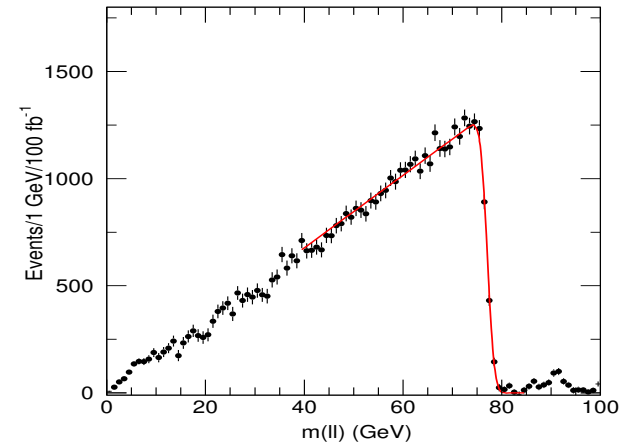


degenerate spectrum of the seesaw particles,  $M_{seesaw} = 10^{14}$  GeV

J. Esteves, M.Hirsch, J. Romão, W.P., F. Staub, Phys. Rev. D83 (2011) 013003



G. Polesello



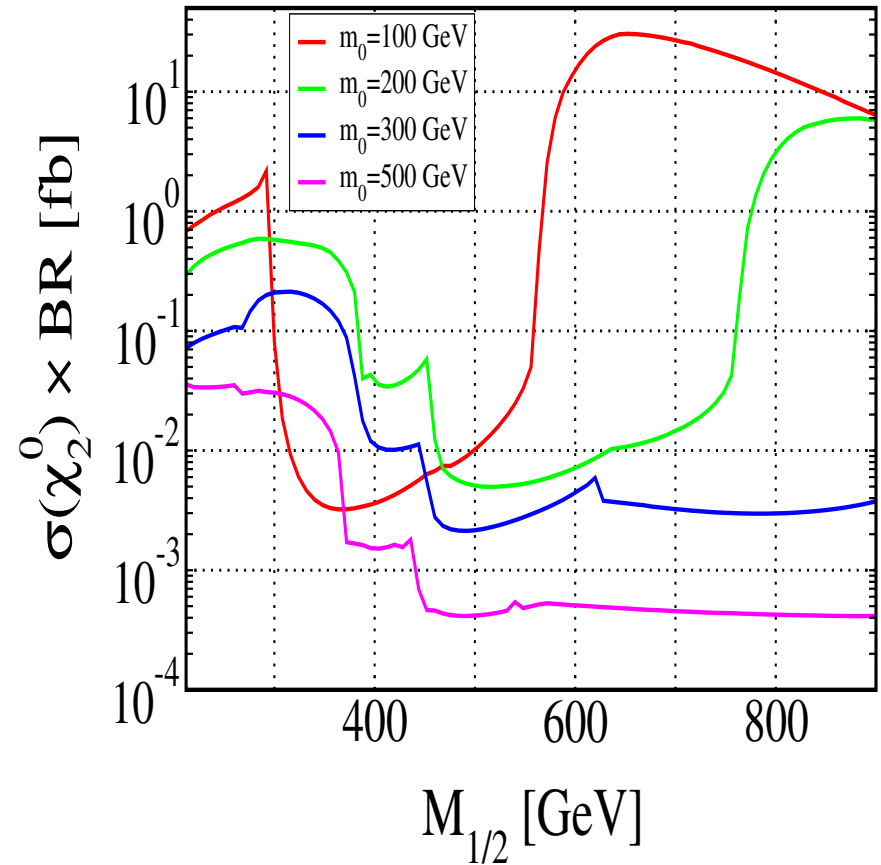
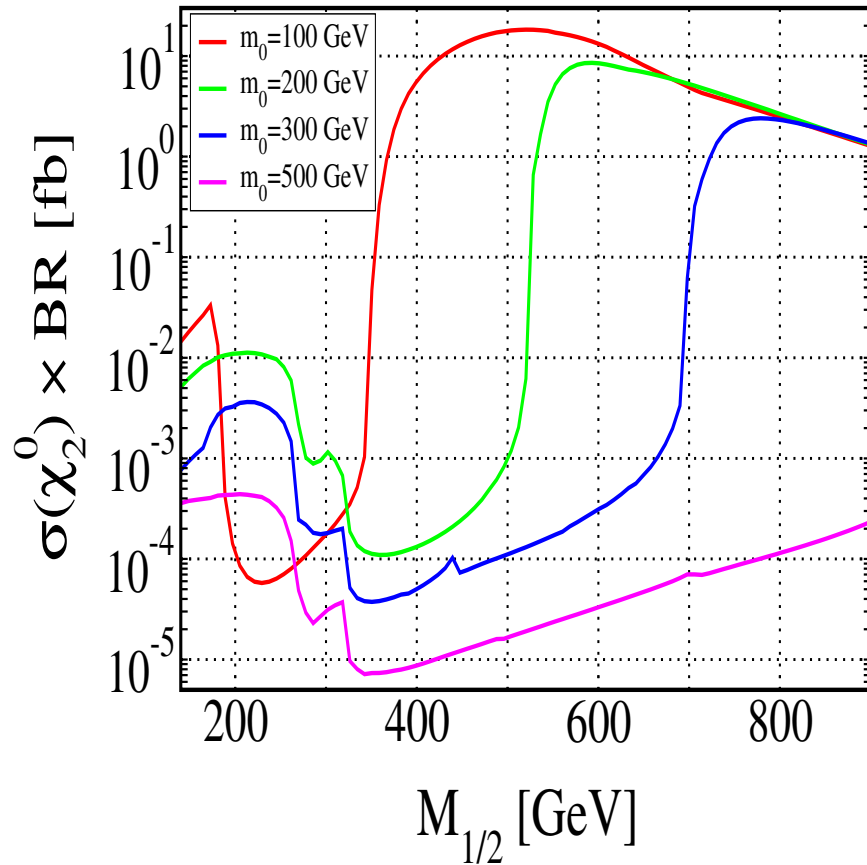
5 kinematical observables depending on 4 SUSY masses

e.g.:  $m(ll) = 77.02 \pm 0.05 \pm 0.08$

$\Rightarrow$  mass determination within 2-5%

For background suppression

$$N(e^+e^-) + N(\mu^+\mu^-) - N(e^+\mu^-) - N(\mu^+e^-)$$



$$\sigma(pp \rightarrow \tilde{\chi}_2^0) \times BR(\chi_2^0 \rightarrow \sum_{i,j} \tilde{l}_i l_j \rightarrow \mu^\pm \tau^\mp \tilde{\chi}_1^0)$$

$$A_0 = 0, \tan \beta = 10, \mu > 0 \text{ (Seesaw II: } \lambda_1 = 0.02, \lambda_2 = 0.5)$$

J.N. Esteves et al., JHEP 0905, 003 (2009)

Exp. input:

- MEG, BELLE, ...:  $BR(\mu \rightarrow e\gamma)$ , ...
- LHC: mass differences , edge variables (including LFV signals)
- ILC/CLIC: mass spectrum, LFV branching ratios, cross sections

Theo. input

- High scale model, e.g. mSUGRA, GMSB
- Seesaw type

	Mass, ideal	“LHC”	“LC”	“LHC+LC”
$h^0$	116.0	0.25	0.05	0.05
$H^0$	425.0		1.5	1.5
$\tilde{\chi}_1^0$	97.7	4.8	0.05	0.05
$\tilde{\chi}_2^0$	183.9	4.7	1.2	0.08
$\tilde{\chi}_4^0$	413.9	5.1	3-5	2.5
$\tilde{\chi}_1^\pm$	183.7		0.55	0.55
$\tilde{e}_R$	125.3	4.8	0.05	0.05
$\tilde{e}_L$	189.9	5.0	0.18	0.18
$\tilde{\tau}_1$	107.9	5-8	0.24	0.24
$\tilde{q}_R$	547.2	7-12	-	5-11
$\tilde{q}_L$	564.7	8.7	-	4.9
$\tilde{t}_1$	366.5		1.9	1.9
$\tilde{b}_1$	506.3	7.5	-	5.7
$\tilde{g}$	607.1	8.0	-	6.5

$$m_0 = 70 \text{ GeV}$$

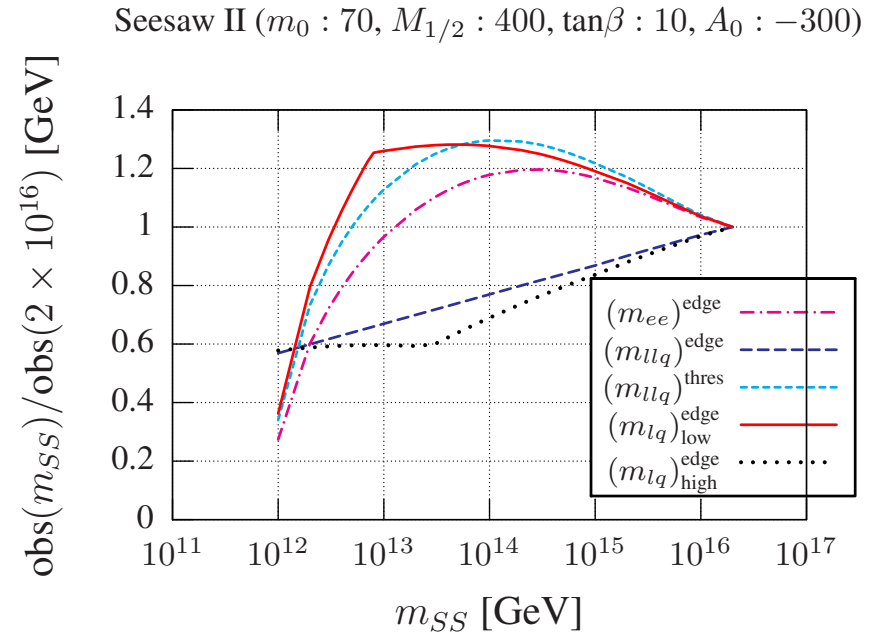
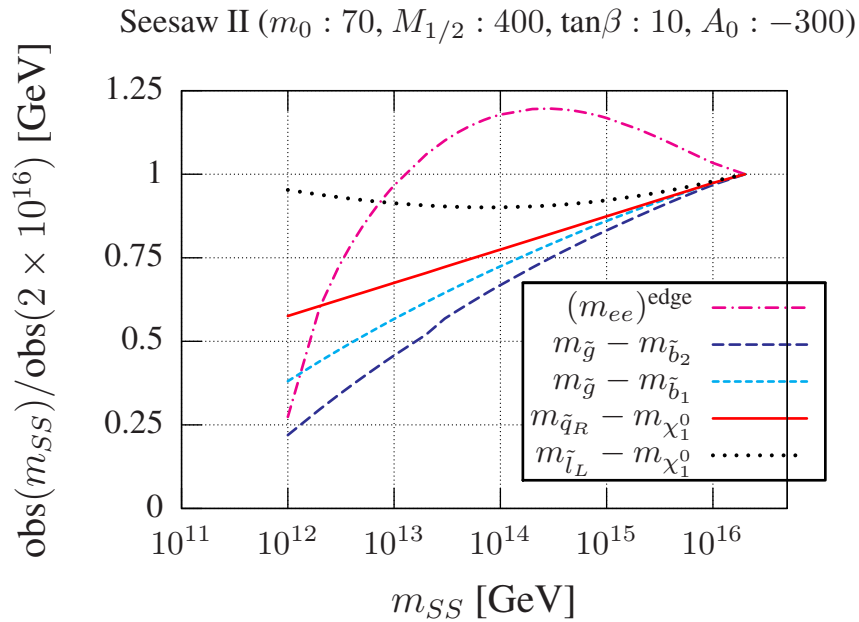
$$m_{1/2} = 250 \text{ GeV}$$

$$A_0 = -300 \text{ GeV}$$

$$\tan \beta = 10$$

$$\text{sign}(\mu) = +$$

G. Weiglein *et al.*, Phys. Rept. 426 (2006) 47; J.A. Aguilar-Saavedra *et al.*, EPJC 46 (2006) 43



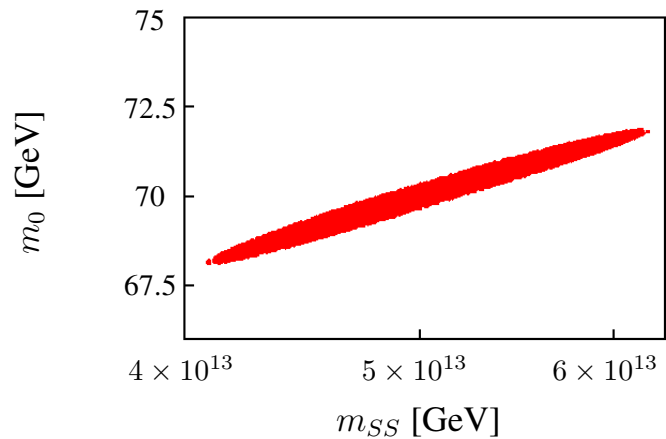
**Kinks**

$$(m_{lq})_{\text{high}}^{\text{edge}} = \max[(m_{l_{\text{near}q}}^{\text{max}})^2, (m_{l_{\text{far}q}}^{\text{max}})^2]$$

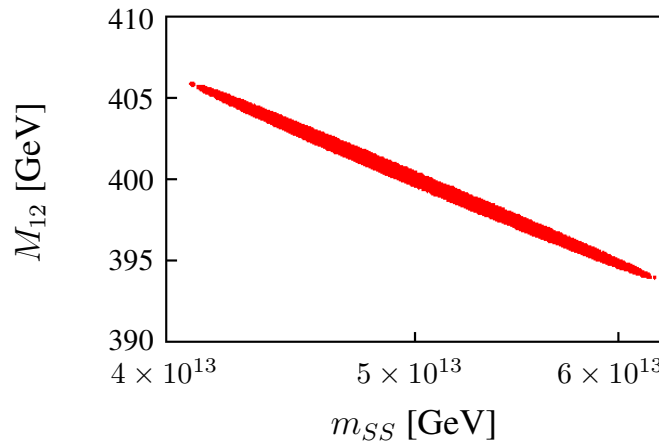
$$(m_{lq})_{\text{low}}^{\text{edge}} = \min[(m_{l_{\text{near}q}}^{\text{max}})^2, (m_{\tilde{q}}^2 - m_{\chi_2^0}^2)(m_{\tilde{l}_R}^2 - m_{\chi_1^0}^2)/(2m_{\tilde{l}_R}^2 - m_{\chi_1^0}^2)]$$

M. Hirsch, W.P., L. Reichert, arXiv:1101.2140 [hep-ph]

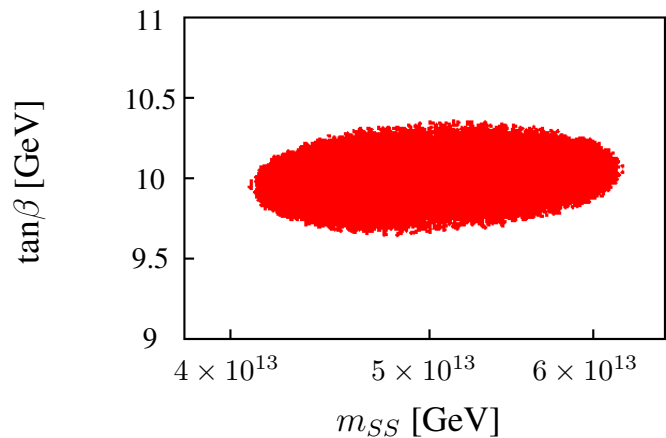
Seesaw II ( $m_0 : 70, M_{1/2} : 400, \tan\beta : 10, A_0 : -300$ )



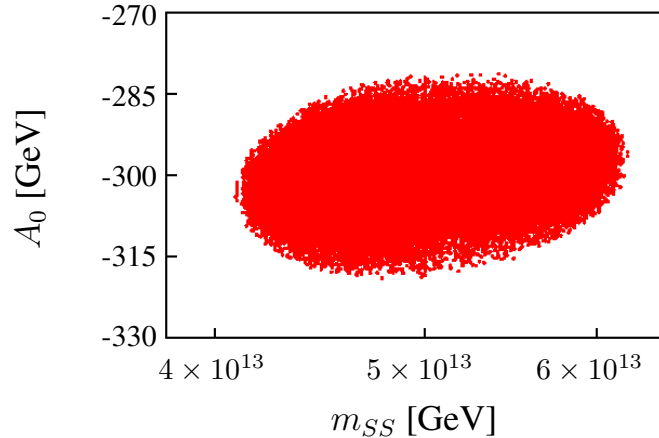
Seesaw II ( $m_0 : 70, M_{1/2} : 400, \tan\beta : 10, A_0 : -300$ )



Seesaw II ( $m_0 : 70, M_{1/2} : 400, \tan\beta : 10, A_0 : -300$ )

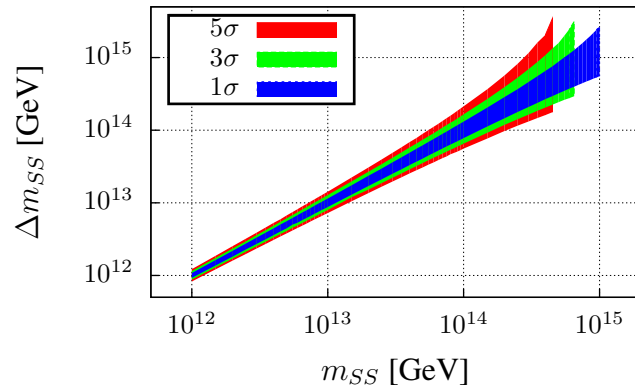


Seesaw II ( $m_0 : 70, M_{1/2} : 400, \tan\beta : 10, A_0 : -300$ )

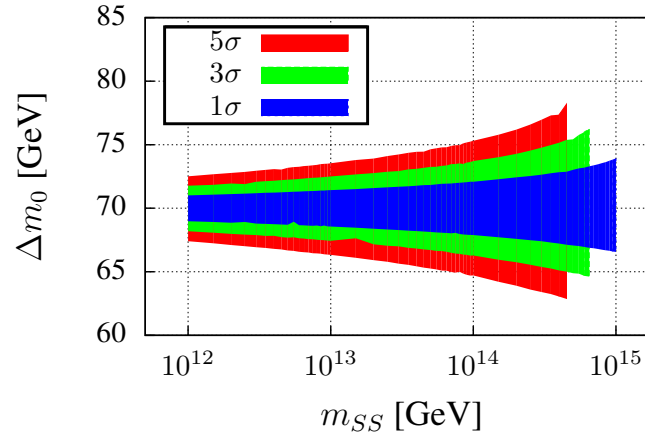


original scale  $5 \times 10^{13}$  GeV; M. Hirsch, W.P., L. Reichert, arXiv:1101.2140 [hep-ph]

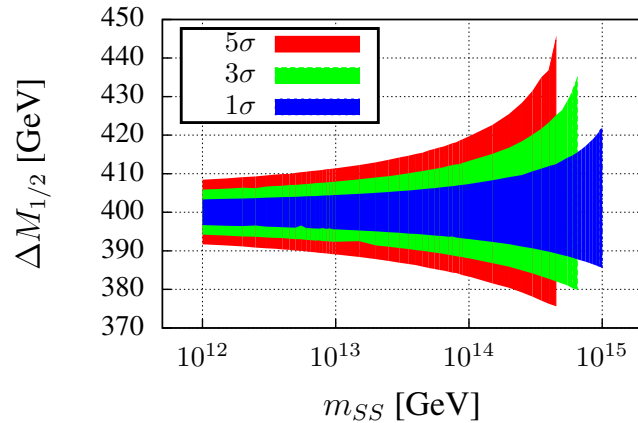
Seesaw II ( $M_0 : 70, M_{1/2} : 400, \tan\beta : 10, A_0 : -300$ )



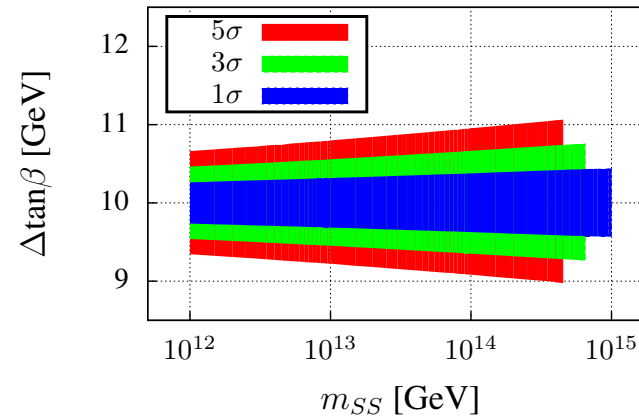
Seesaw II ( $M_0 : 70, M_{1/2} : 400, \tan\beta : 10, A_0 : -300$ )



Seesaw II ( $M_0 : 70, M_{1/2} : 400, \tan\beta : 10, A_0 : -300$ )

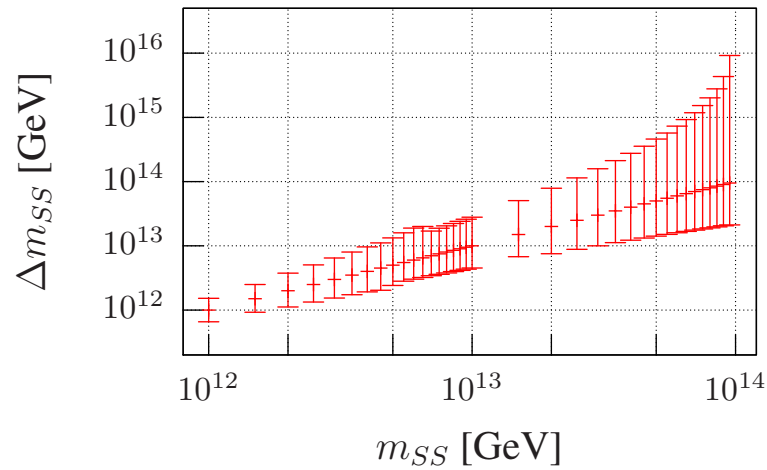


Seesaw II ( $M_0 : 70, M_{1/2} : 400, \tan\beta : 10, A_0 : -300$ )

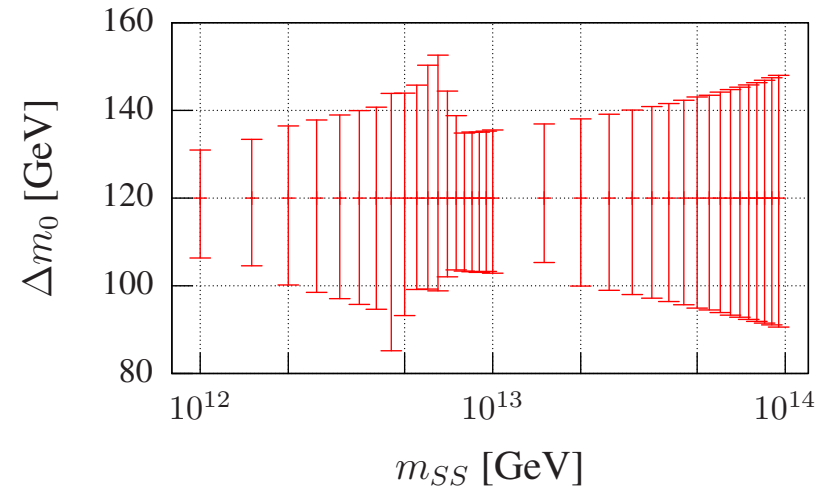




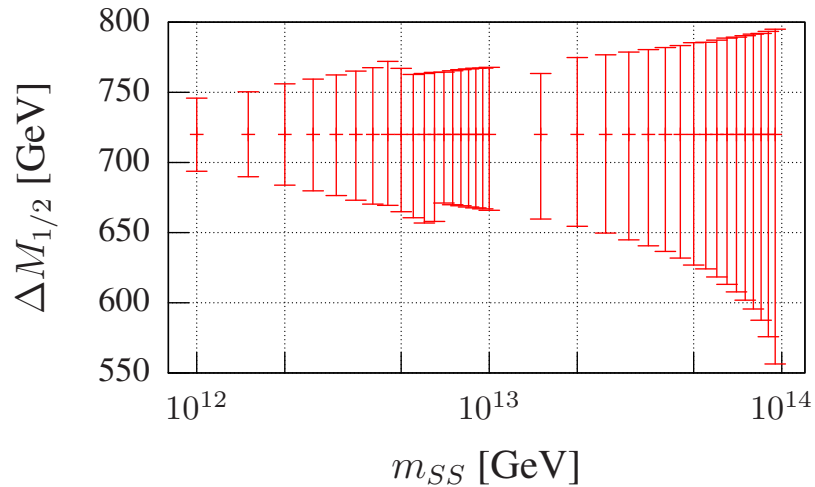
Seesaw II ( $m_0 : 120, M_{1/2} : 720, \tan\beta : 10, A_0 : 0$ )



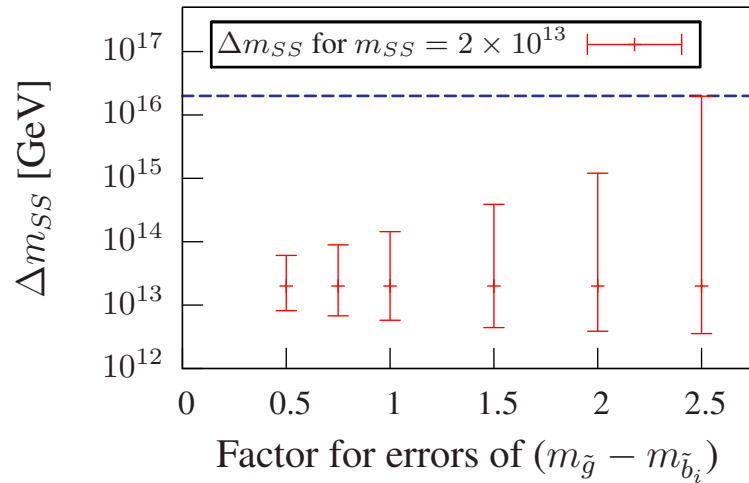
Seesaw II ( $m_0 : 120, M_{1/2} : 720, \tan\beta : 10, A_0 : 0$ )



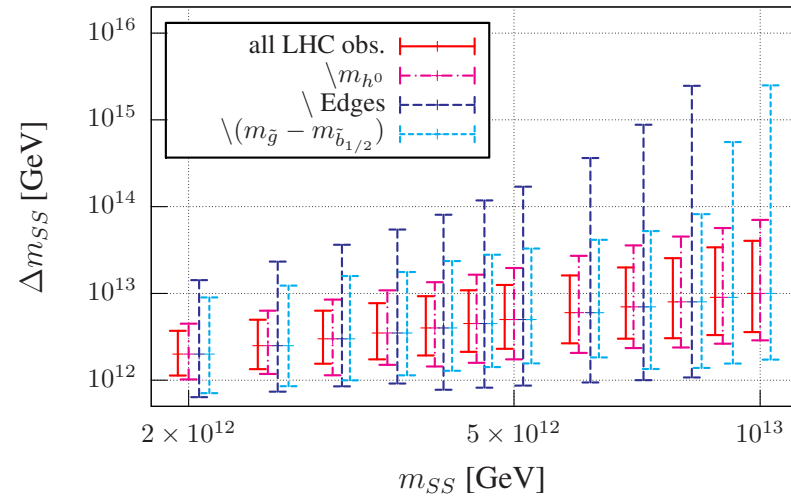
Seesaw II ( $m_0 : 120, M_{1/2} : 720, \tan\beta : 10, A_0 : 0$ )



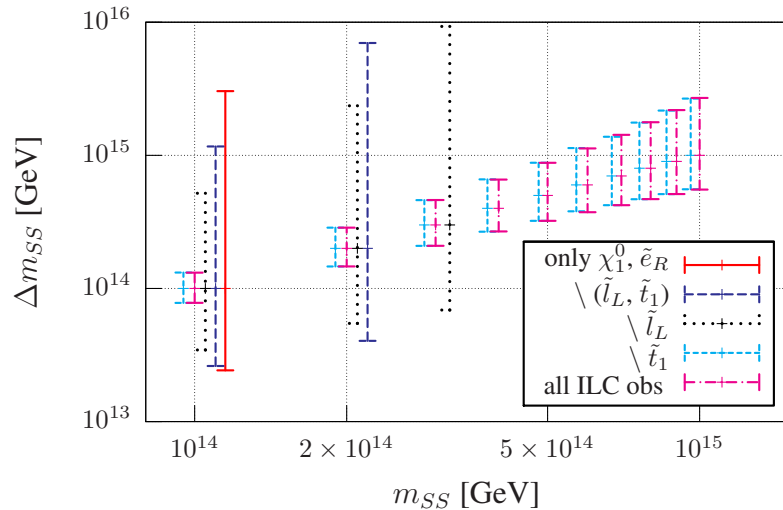
Seesaw II ( $m_0 : 70, M_{1/2} : 400, \tan\beta : 10, A_0 : -300$ )



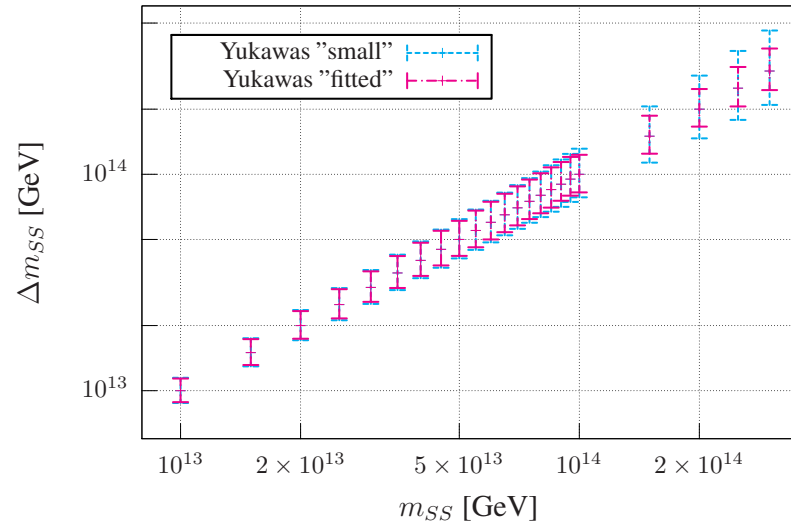
Seesaw II ( $m_0 : 70, M_{1/2} : 400, \tan\beta : 10, A_0 : -300$ )



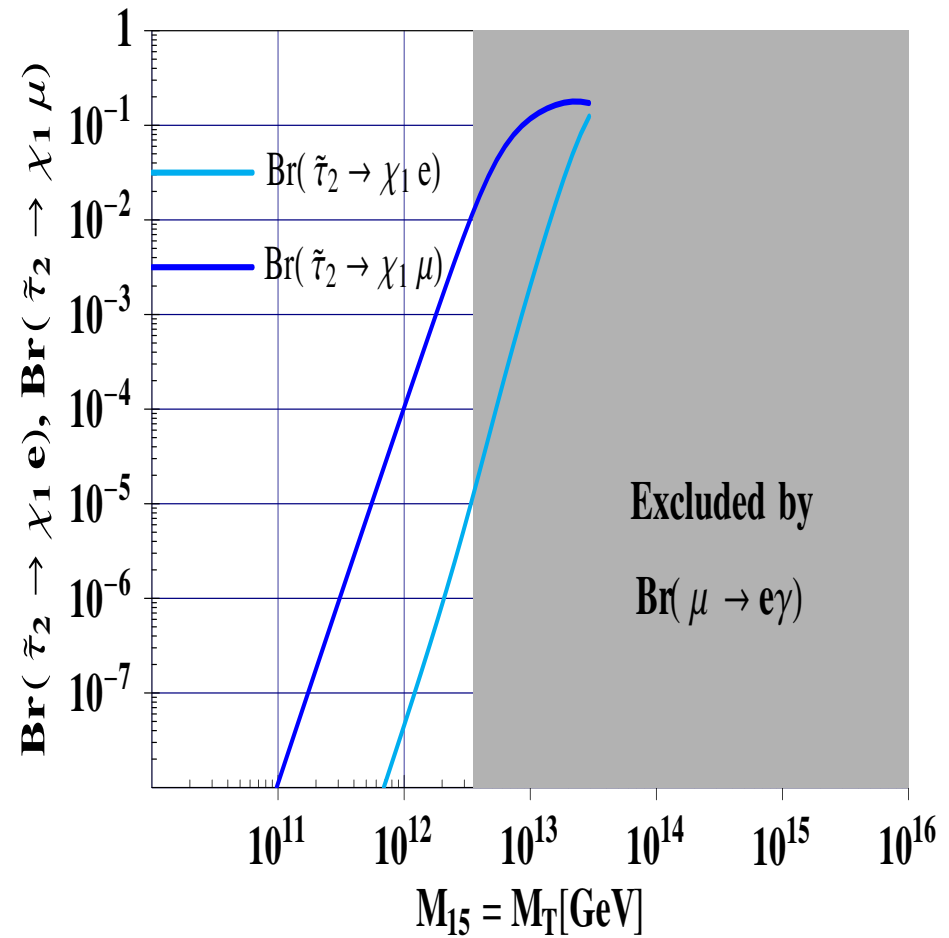
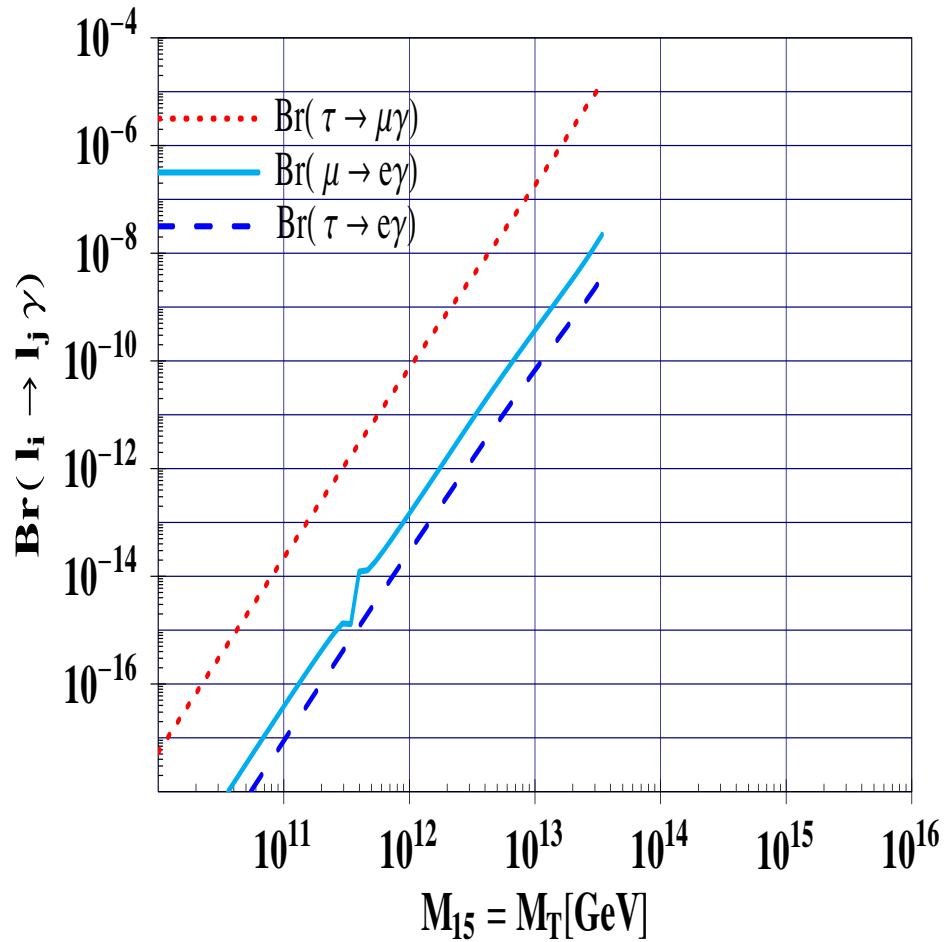
Seesaw II ( $m_0 : 70, M_{1/2} : 400, \tan\beta : 10, A_0 : -300$ )



Seesaw II ( $m_0 : 70, M_{1/2} : 400, \tan\beta : 10, A_0 : -300$ )



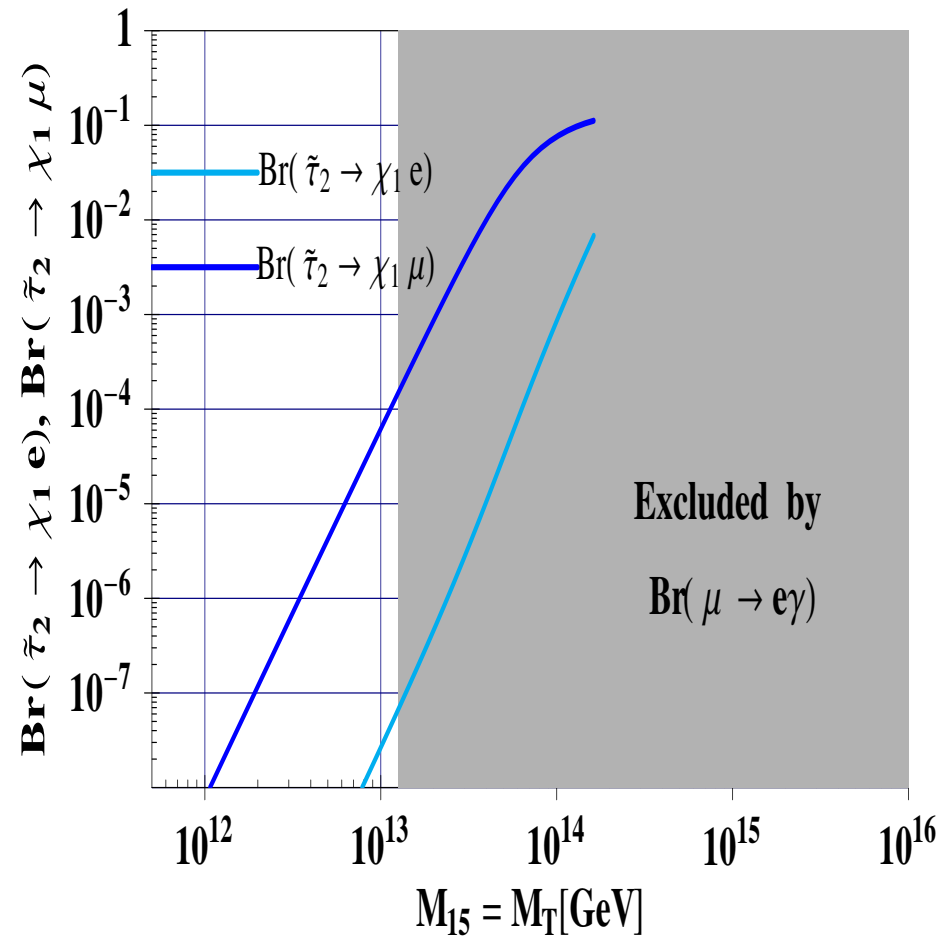
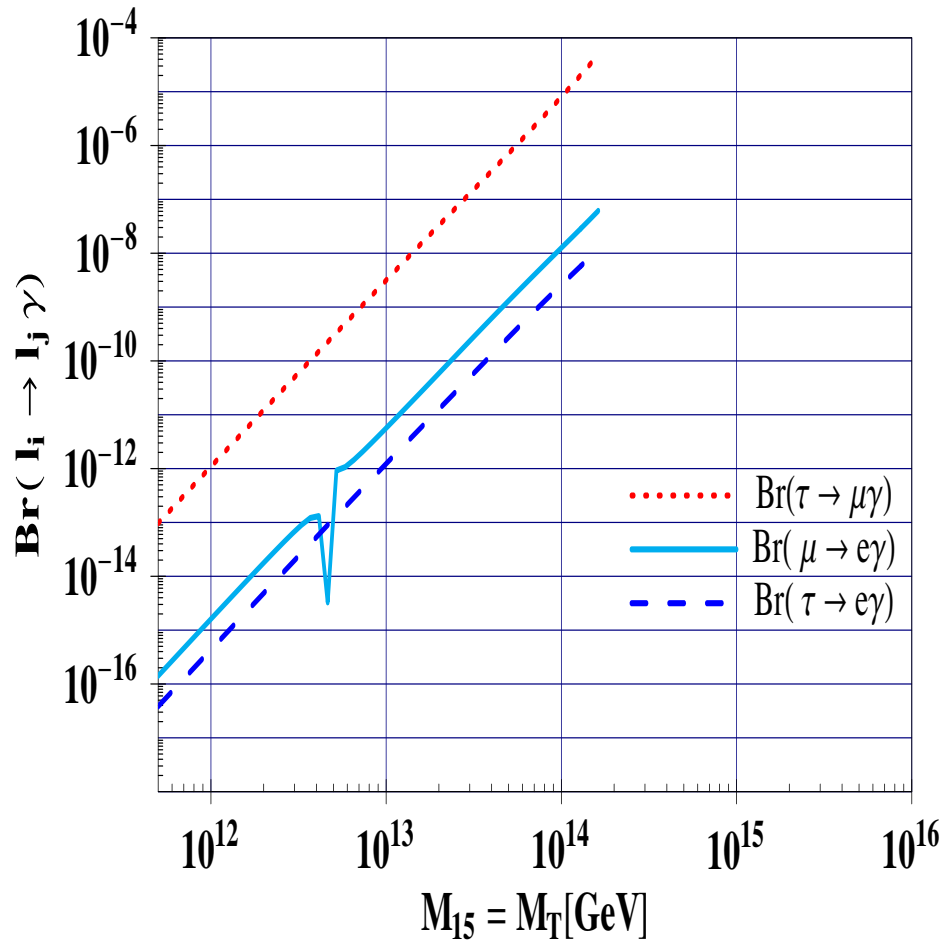
- Dirac neutrinos: displaced vertices if  $\tilde{\nu}_R$  LSP, e.g.  $\tilde{t}_1 \rightarrow lb\tilde{\nu}_R$   
(but NMSSM:  $\tilde{t}_1 \rightarrow lb\nu\tilde{\chi}_1^0$ )
- Seesaw models:
  - in case of seesaw II, III: different mass ratios
  - promoting:  $\tilde{\tau}_2$  decays
  - LFV signals difficult to test at LHC, of O(10 fb) or below
- Model reconstruction for seesaw II, III
  - LHC: possible in favourable parts of parameter space; might improve if additional observables can be used, e.g. lepton and jet spectra
  - LHC+ILC: possible for large part of parameter space, however still some model dependence
  - to distinguish between type II and III: need in addition LFV signals



$$\lambda_1 = \lambda_2 = 0.05$$

SPS3 ( $M_0 = 90 \text{ GeV}$ ,  $M_{1/2} = 400 \text{ GeV}$ ,  $A_0 = 0 \text{ GeV}$ ,  $\tan \beta = 10$ ),  $\mu > 0$

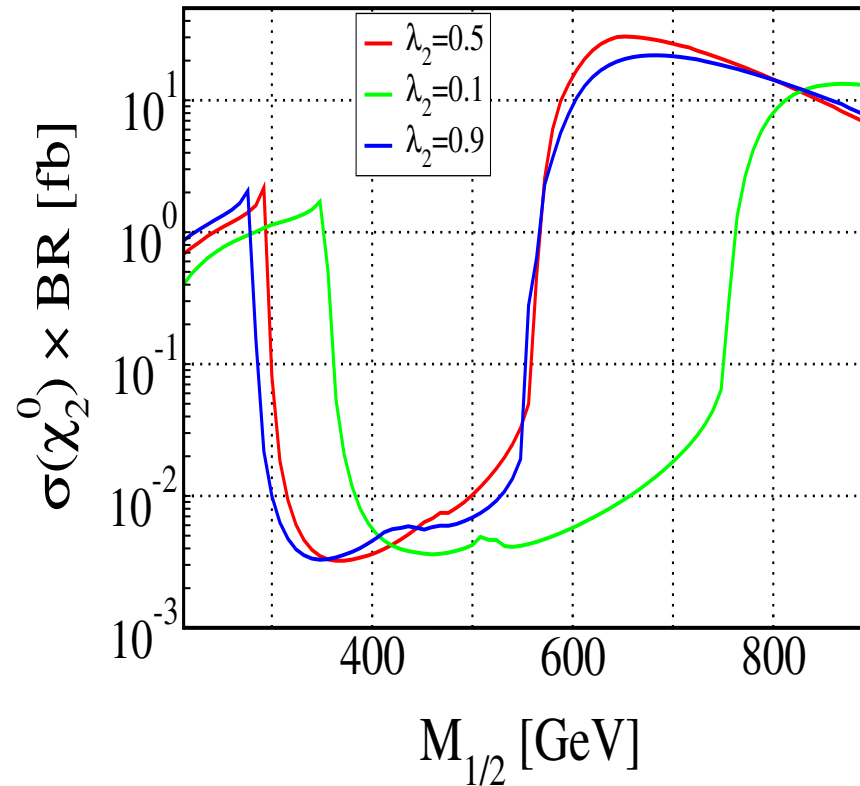
M. Hirsch, S. Kaneko, W. P., Phys. Rev. D 78 (2008) 093004.



$$\lambda_1 = \lambda_2 = 0.5$$

SPS1a' ( $M_0 = 70 \text{ GeV}$ ,  $M_{1/2} = 250 \text{ GeV}$ ,  $A_0 = -300 \text{ GeV}$ ,  $\tan \beta = 10$ ),  $\mu > 0$

M. Hirsch, S. Kaneko, W. P., Phys. Rev. D 78 (2008) 093004.



$$\sigma(pp \rightarrow \tilde{\chi}_2^0) \times BR(\chi_2^0 \rightarrow \sum_{i,j} \tilde{l}_i l_j \rightarrow \mu^\pm \tau^\mp \tilde{\chi}_1^0)$$

$m_0 = 100 \text{ GeV}$   $A_0 = 0$ ,  $\tan \beta = 10$ ,  $\mu > 0$ ,  $\lambda_1 = 0.02$

J.N. Esteves et al., JHEP 0905, 003 (2009)

talk by I. Borjanovic at 'Flavour in the era of LHC', Nov.'05, CERN

**L=100 fb<sup>-1</sup>**

**Fit results**

Edge	Nominal Value	Fit Value	Syst. Error Energy Scale	Statistical Error
$m(ll)^{edge}$	77.077	77.024	0.08	0.05
$m(qll)^{edge}$	431.1	431.3	4.3	2.4
$m(ql)^{edge}_{min}$	302.1	300.8	3.0	1.5
$m(ql)^{edge}_{max}$	380.3	379.4	3.8	1.8
$m(qll)^{thres}$	203.0	204.6	2.0	2.8

**Mass reconstruction**

5 endpoints measurements, 4 unknown masses

$$\chi^2 = \sum \chi_j^2 = \sum \left[ \frac{E_j^{theory}(\vec{m}) - E_j^{exp}}{\sigma_j^{exp}} \right]^2$$

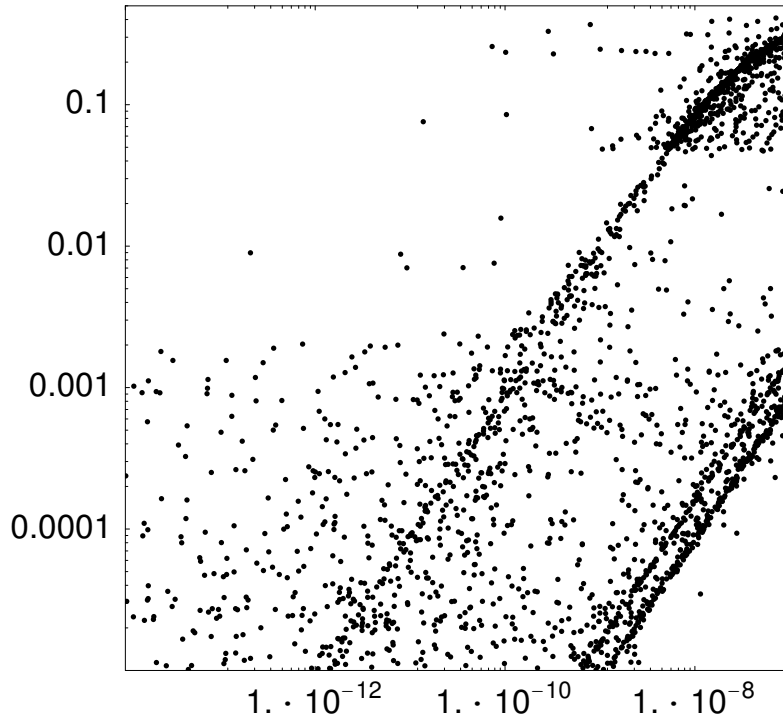
$$E_j^i = E_j^{nom} + a_j^i \sigma_j^{fit} + b_j^i E_{scale}$$

$m(\chi_1^0) = 96 \text{ GeV}$   
 $m(l_R) = 143 \text{ GeV}$   
 $m(\chi_2^0) = 177 \text{ GeV}$   
 $m(q_L) = 540 \text{ GeV}$

$\Delta m(\chi_1^0) = 4.8 \text{ GeV}, \quad \Delta m(\chi_2^0) = 4.7 \text{ GeV},$   
 $\Delta m(l_R) = 4.8 \text{ GeV}, \quad \Delta m(q_L) = 8.7 \text{ GeV}$

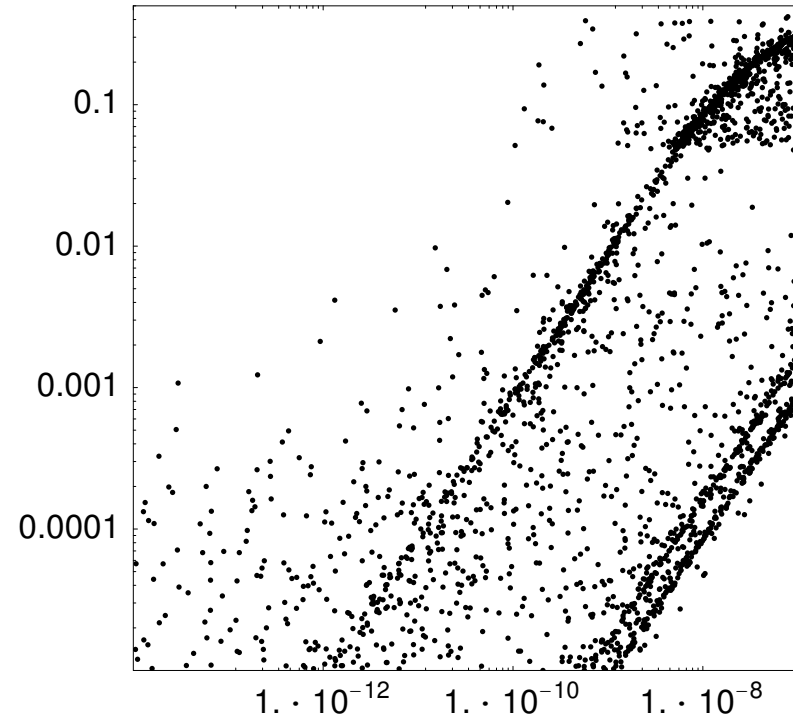
Gjelsten, Lytken, Miller, Osland, Polesello, ATL-PHYS-2004-007

$$\text{BR}(\tilde{\chi}_2^0 \rightarrow \tilde{\chi}_1^0 e^\pm \tau^\mp)$$



$$\text{BR}(\tau \rightarrow e\gamma)$$

$$\text{BR}(\tilde{\chi}_2^0 \rightarrow \tilde{\chi}_1^0 \mu^\pm \tau^\mp)$$



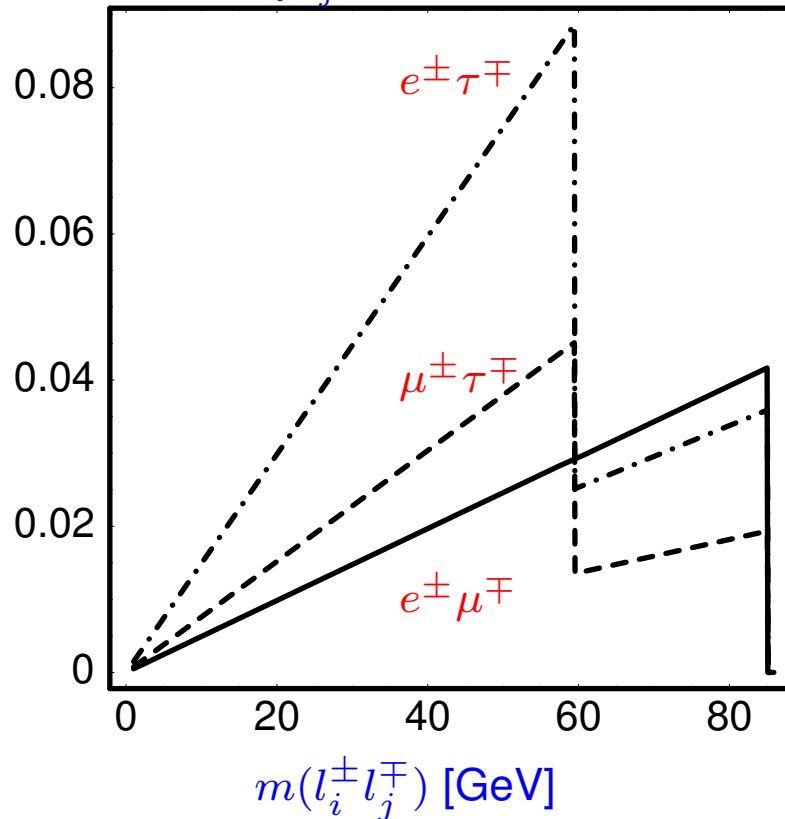
$$\text{BR}(\tau \rightarrow \mu\gamma)$$

Variations around SPS1a

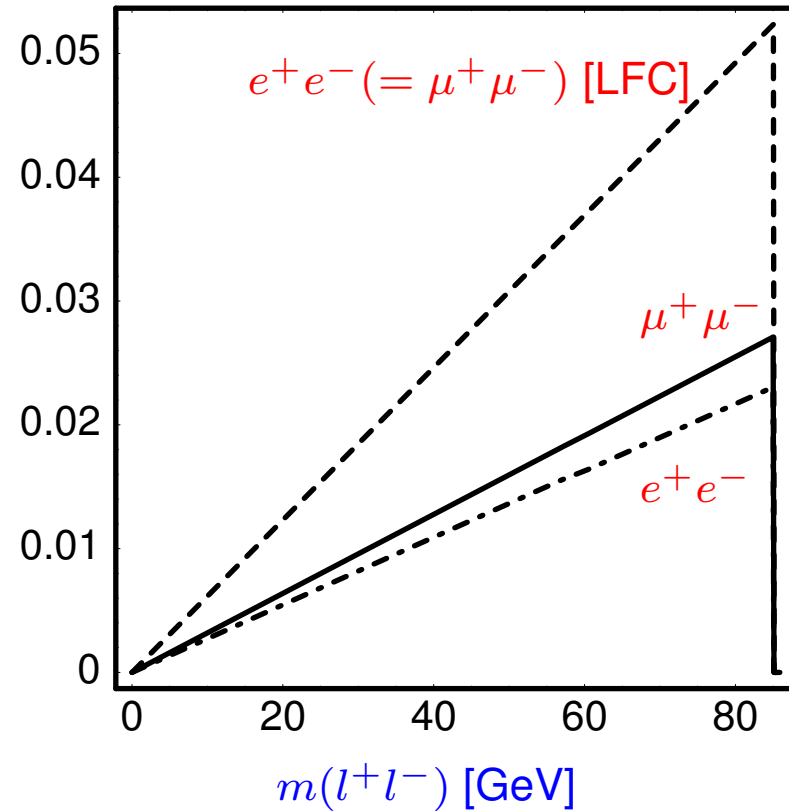
$$(M_0 = 100 \text{ GeV}, M_{1/2} = 250 \text{ GeV}, A_0 = -100 \text{ GeV}, \tan \beta = 10)$$



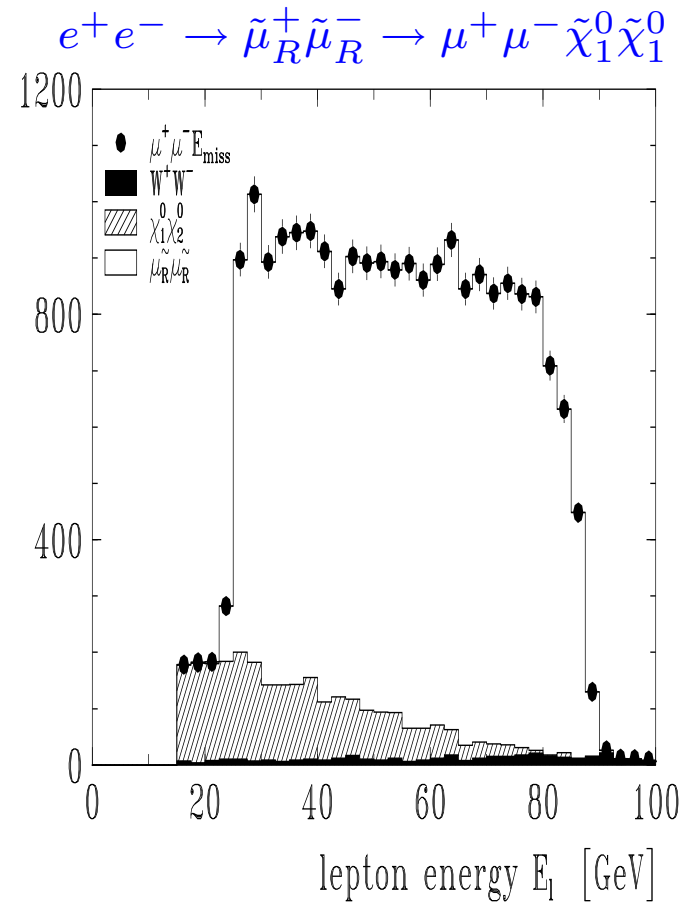
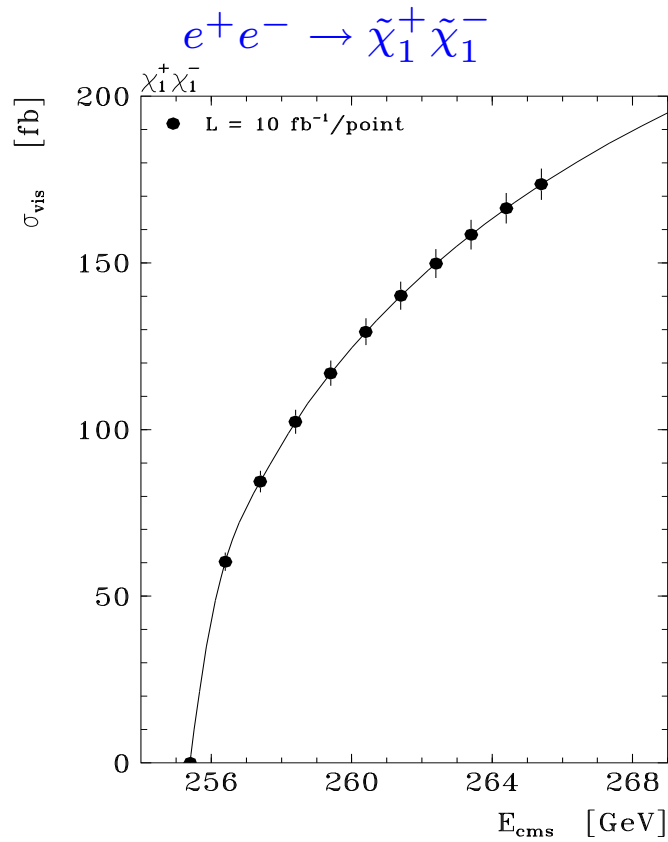
$$\frac{100}{\Gamma_{tot}} \frac{d\Gamma(\tilde{\chi}_2^0 \rightarrow l_i^\pm l_j^\mp \tilde{\chi}_1^0)}{dm(l_i^\pm l_j^\mp)}$$



$$\frac{100}{\Gamma_{tot}} \frac{d\Gamma(\tilde{\chi}_2^0 \rightarrow l^+ l^- \tilde{\chi}_1^0)}{dm(l^+ l^-)}$$

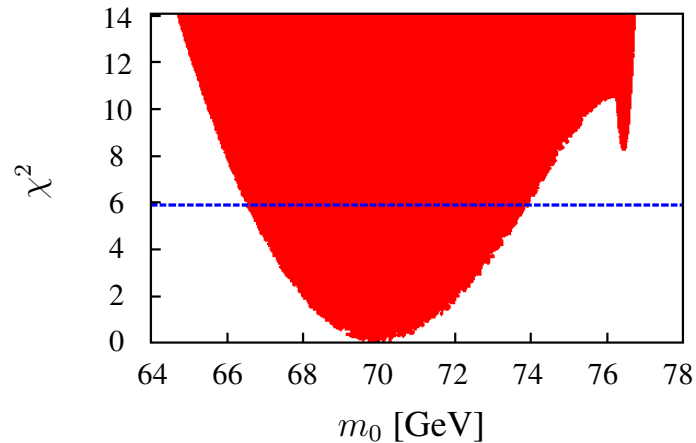


A. Bartl et al., Eur. Phys. J. C 46 (2006) 783

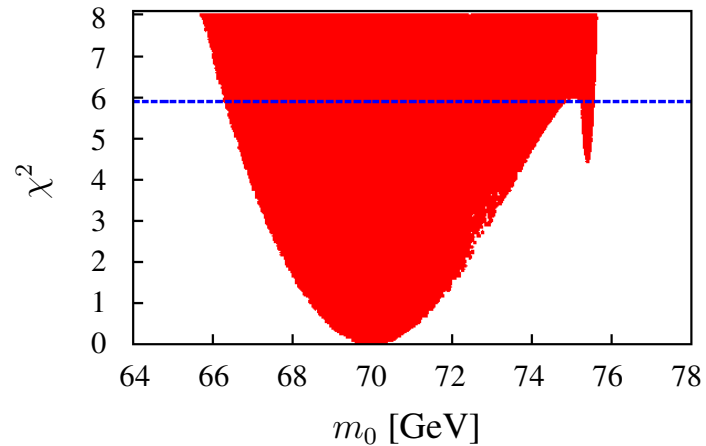


U. Martyn

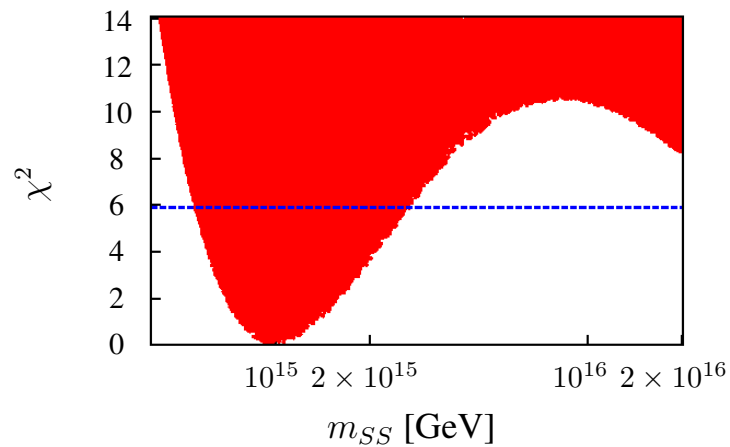
Seesaw II ( $m_0 : 70, M_{1/2} : 400, \tan\beta : 10, A_0 : -300$ )



Seesaw II ( $m_0 : 70, M_{1/2} : 400, \tan\beta : 10, A_0 : -300$ )

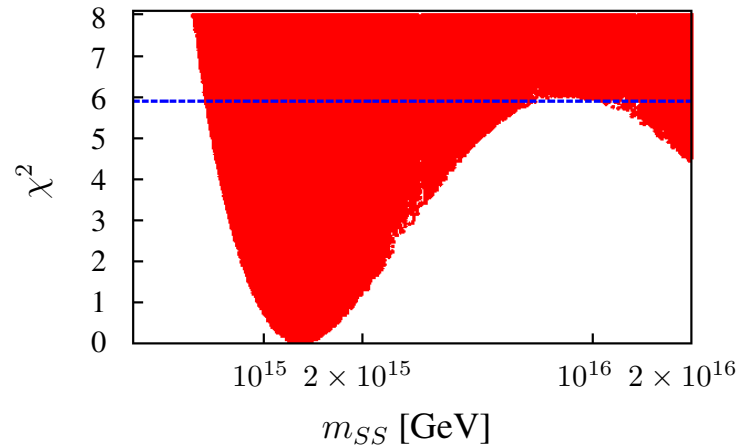


Seesaw II ( $m_0 : 70, M_{1/2} : 400, \tan\beta : 10, A_0 : -300$ )



original scale  $1 \times 10^{15}$  GeV

Seesaw II ( $m_0 : 70, M_{1/2} : 400, \tan\beta : 10, A_0 : -300$ )



original scale  $1.3 \times 10^{15}$  GeV